

The modern view of planet formation



Wladimir Lyra
New Mexico State University

Planet Formation is an active and evolving field of research

Google planet formation

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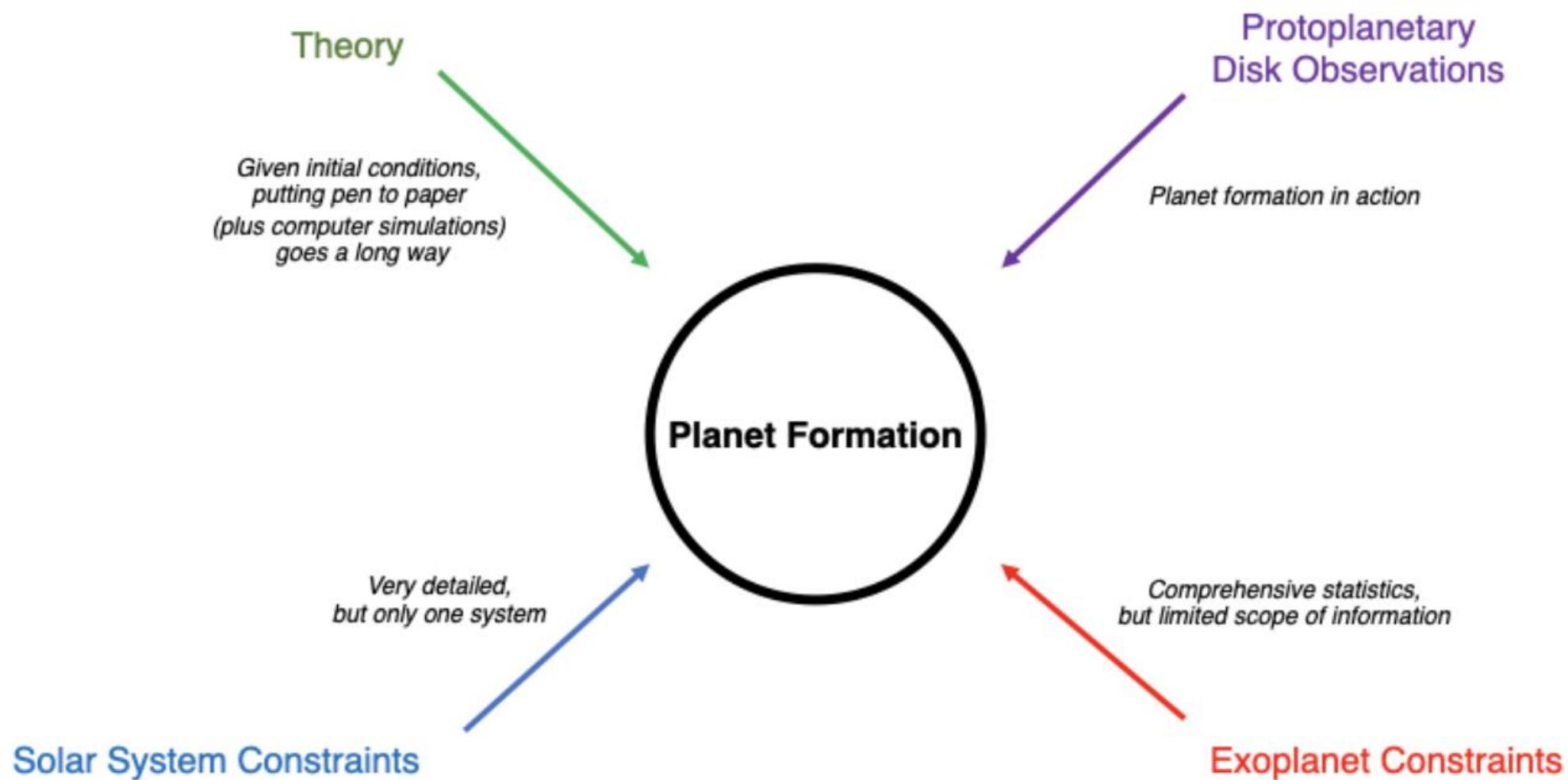
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Universe Today Scientists Gain a New Understanding of How Stars and Planets Form As young stars form, they exert a powerful influence on their surroundings and create complex interactions between them and their... 1 day ago

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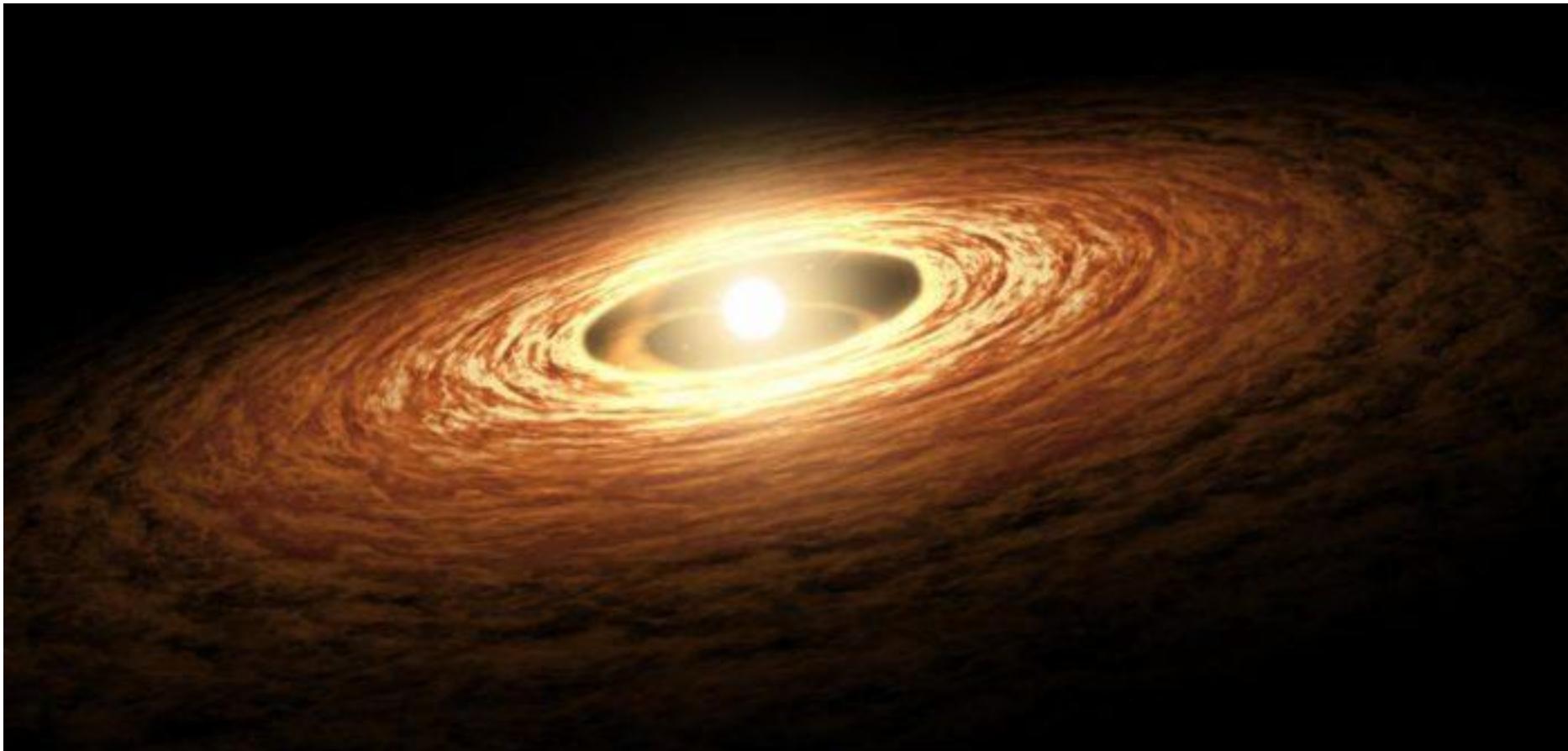




Voyager 10 approach sequence

The Solar Nebula

Nebular hypothesis – planets form in disks of gas and dust
(Kant 1755, Laplace 1794)





Betelgeuse

Bellatrix

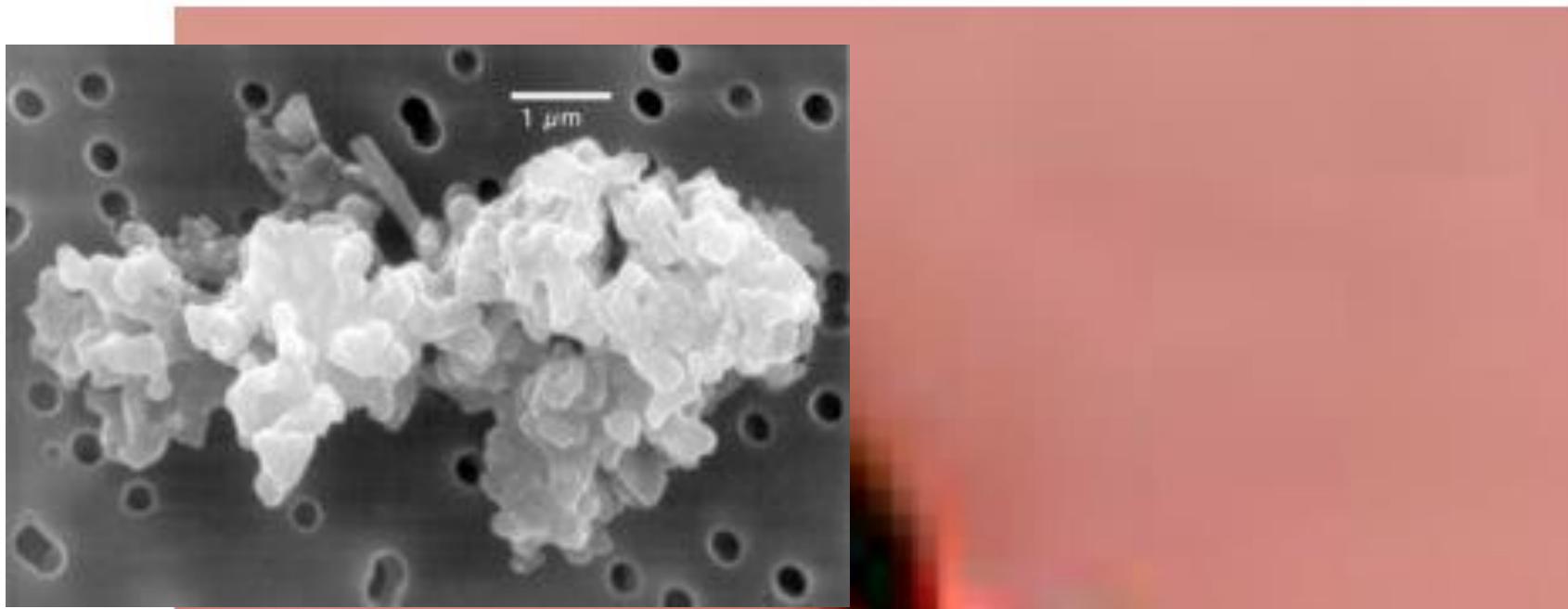
Orion's Belt

Orion Nebula

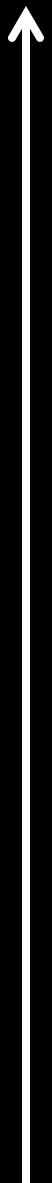
Rigel

Saiph









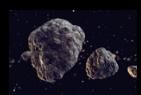
100 *Earth masses*



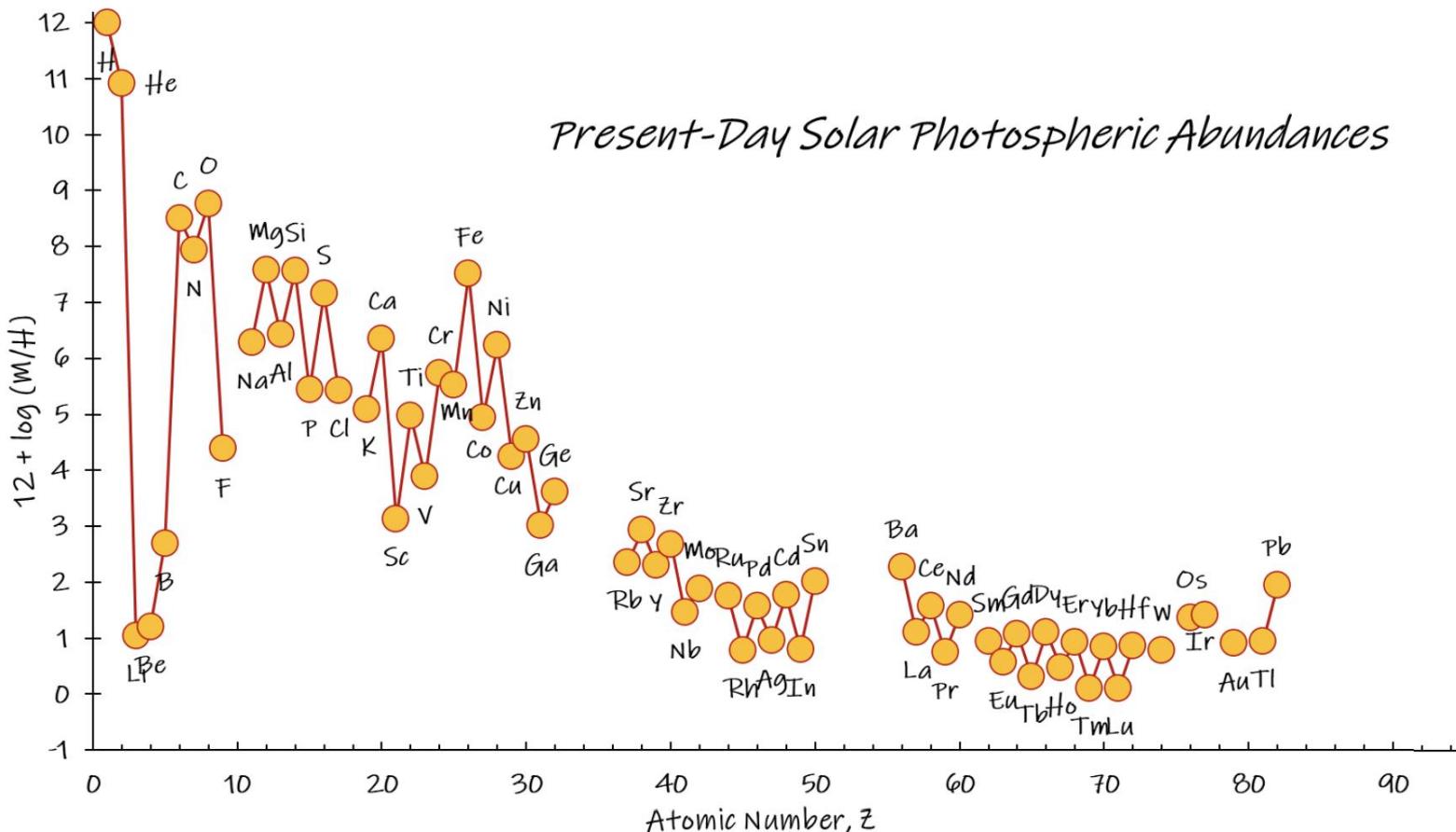
10 *Earth masses*



1 *Earth mass*



Chemical Composition of the Sun



Most abundant elements, in order:

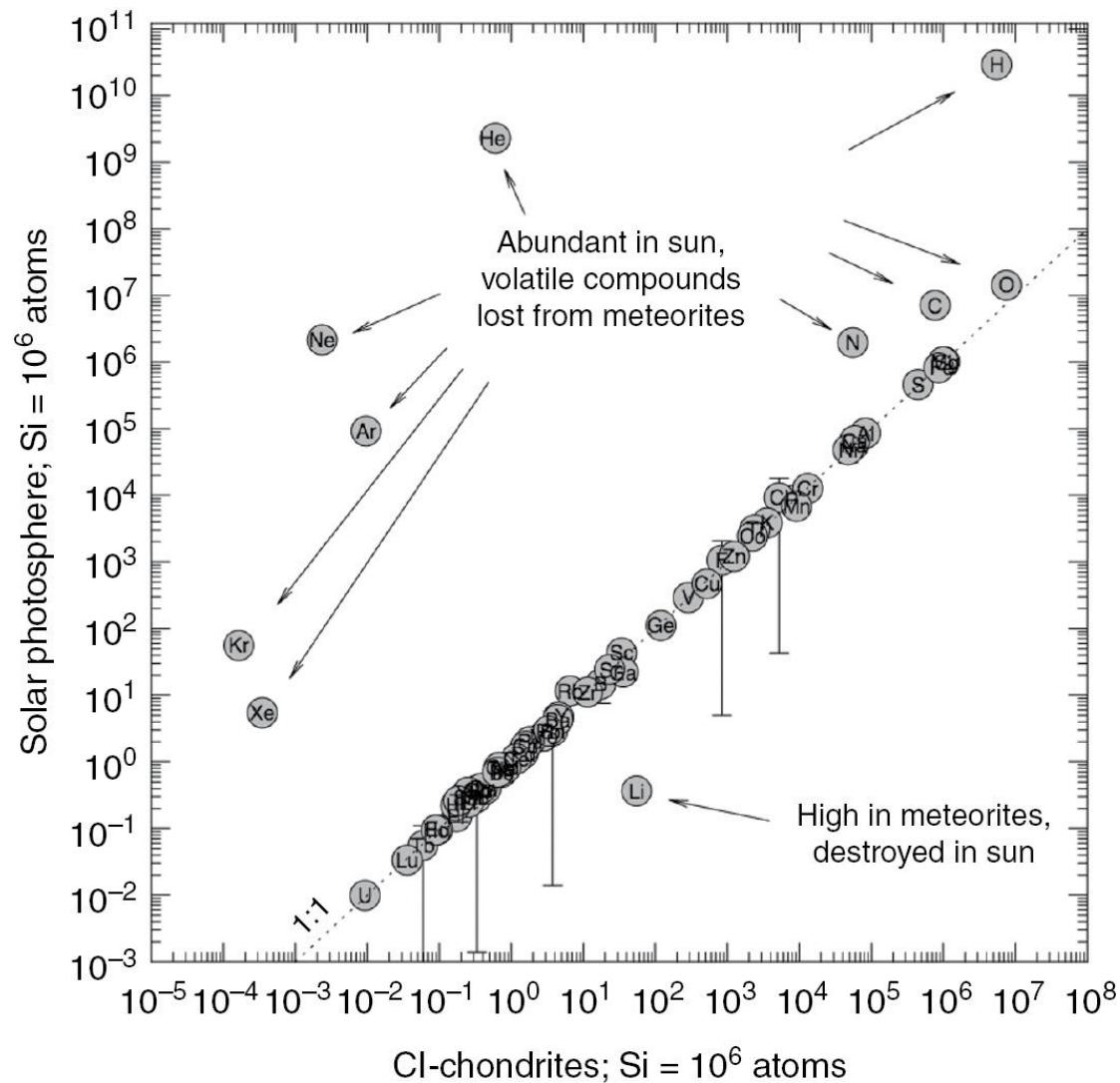
H (71%) He (27%)

O (1.04%) C (0.46%)

Ne (0.13%) Fe (0.11%) N (0.1%)

Si (0.06%), Mg (0.05%), S (0.04%)

Refractories in meteorites: Solar Composition



Periodic Table

1	H	Hydrogen		2	He	Helium																																										
3	Li	Lithium	4	Be	Beryllium																																											
11	Na	Sodium	12	Mg	Magnesi...																																											
19	K	Potassium	20	Ca	Calcium	21	Sc	Scandium																																								
37	Rb	Rubidium	38	Sr	Strontium	39	Y	Yttrium																																								
55	Cs	Caesium	56	Ba	Barium	57	La	Lanthan...																																								
87	Fr	Francium	88	Ra	Radium	89	Ac	Actinium																																								
			104	Rf	Rutherford...	105	Db	Dubnium	106	Sg	Seaborg...	107	Bh	Bohrium	108	Hs	Hassium	109	Mt	Meitneri...	110	Ds	Darmsta...	111	Rg	Roentge...	112	Cn	Coperni...	113	Nh	Nihonium	114	Fl	Flerovium	115	Mc	Moscovi...	116	Lv	Livermor...	117	Ts	Tenness...	118	Og	Oganess...	
			22	Ti	Titanium	23	V	Vanadium	24	Cr	Chromium	25	Mn	Mangan...	26	Fe	Iron	27	Co	Cobalt	28	Ni	Nickel	29	Cu	Copper	30	Zn	Zinc	31	Ga	Gallium	32	Ge	Germani...	33	As	Arsenic	34	Se	Selenium	35	Br	Bromine	36	Kr	Krypton	
			40	Zr	Zirconium	41	Nb	Niobium	42	Mo	Molybde...	43	Tc	Technetiu...	44	Ru	Ruthenium	45	Rh	Rhodium	46	Pd	Palladium	47	Ag	Silver	48	Cd	Cadmium	49	In	Indium	50	Sn	Tin	51	Sb	Antimony	52	Te	Tellurium	53	I	Iodine	54	Xe	Xenon	
			72	Hf	Hafnium	73	Ta	Tantalum	74	W	Tungsten	75	Re	Rhenium	76	Os	Osmium	77	Ir	Iridium	78	Pt	Platinum	79	Au	Gold	80	Hg	Mercury	81	Tl	Thallium	82	Pb	Lead	83	Bi	Bismuth	84	Po	Polonium	85	At	Astatine	86	Rn	Radon	

Alkali metals

○ Metalloids

○ Actinides

○ Alkaline earth metals

Reactive nonmetals

○ Unknown properties

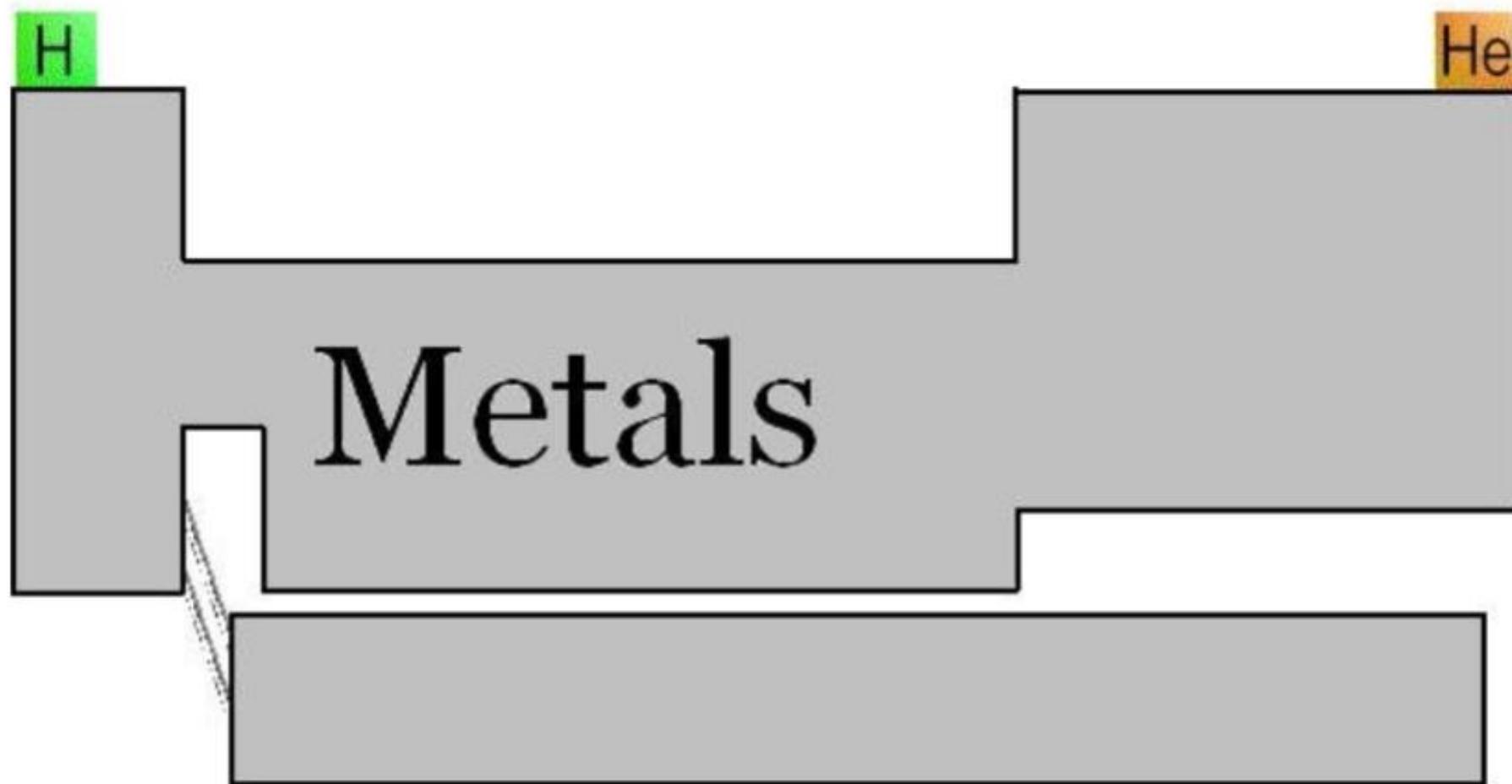
○ Transition metals

Noble gases

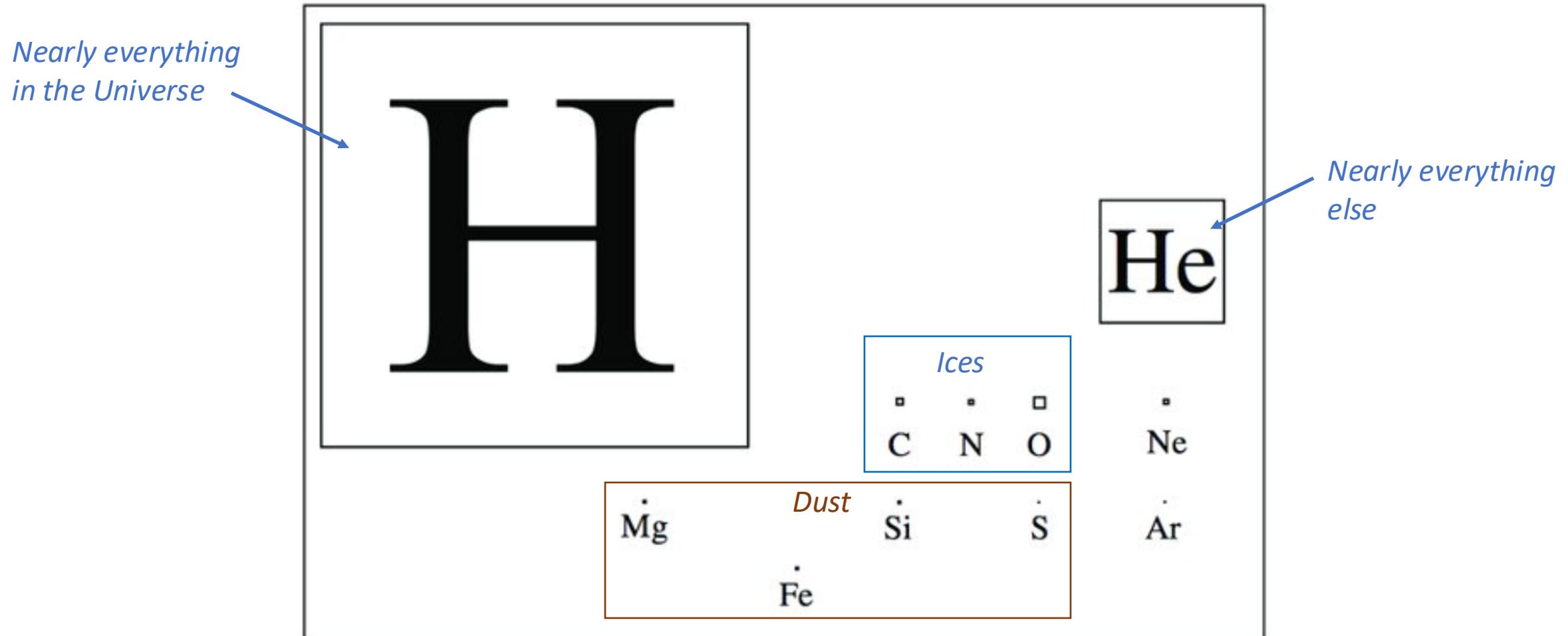
● Post-transition metals

○ Lanthanides

The Astronomer's Periodic Table



The Planetary Astronomer's Periodic Table

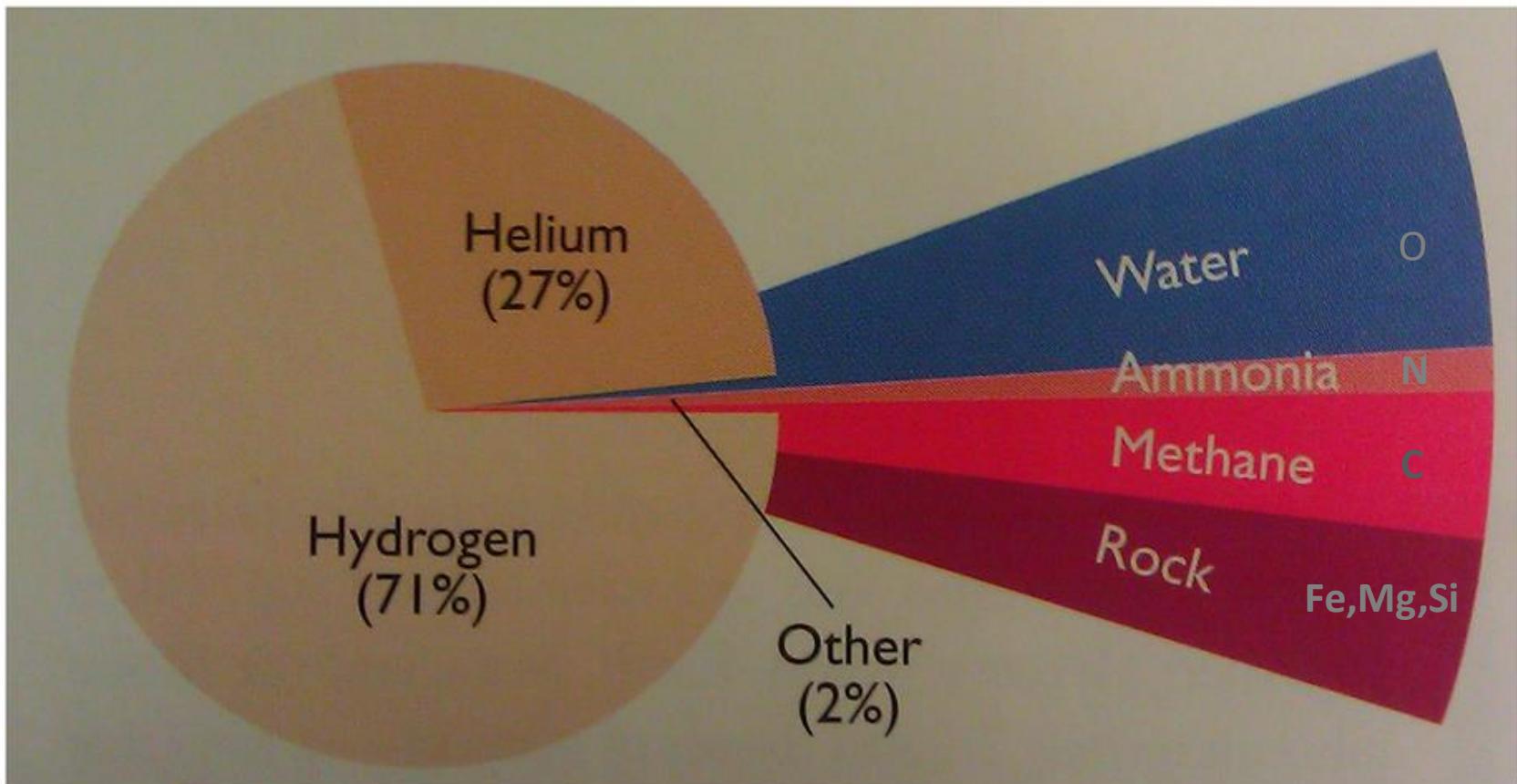


Solar composition:

H ~ 0.71

He ~ 0.27

Metals ~ 0.02



Classes of planets

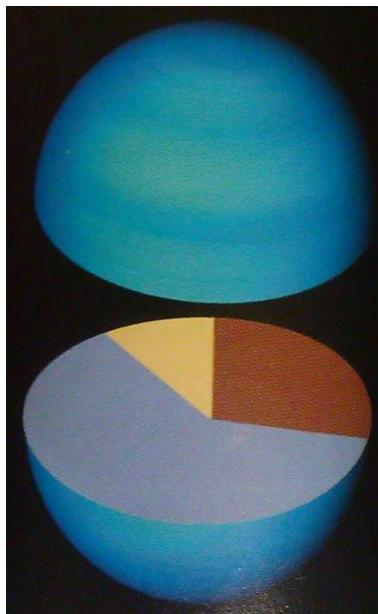
Rocky Planets

Earth



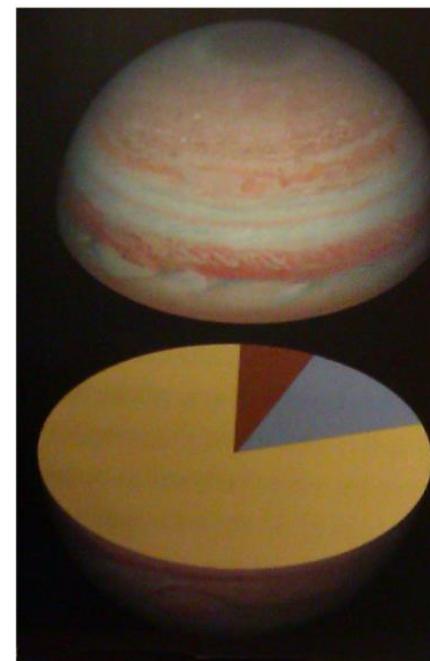
Ice Giants

Uranus/Neptune

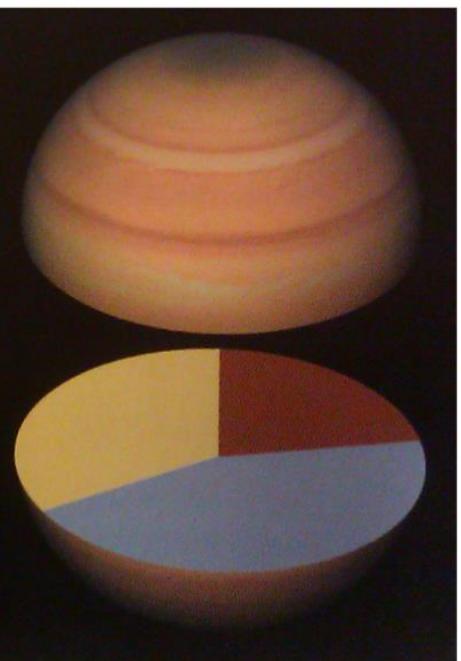


Gas Giants

Jupiter



Saturn



Rock

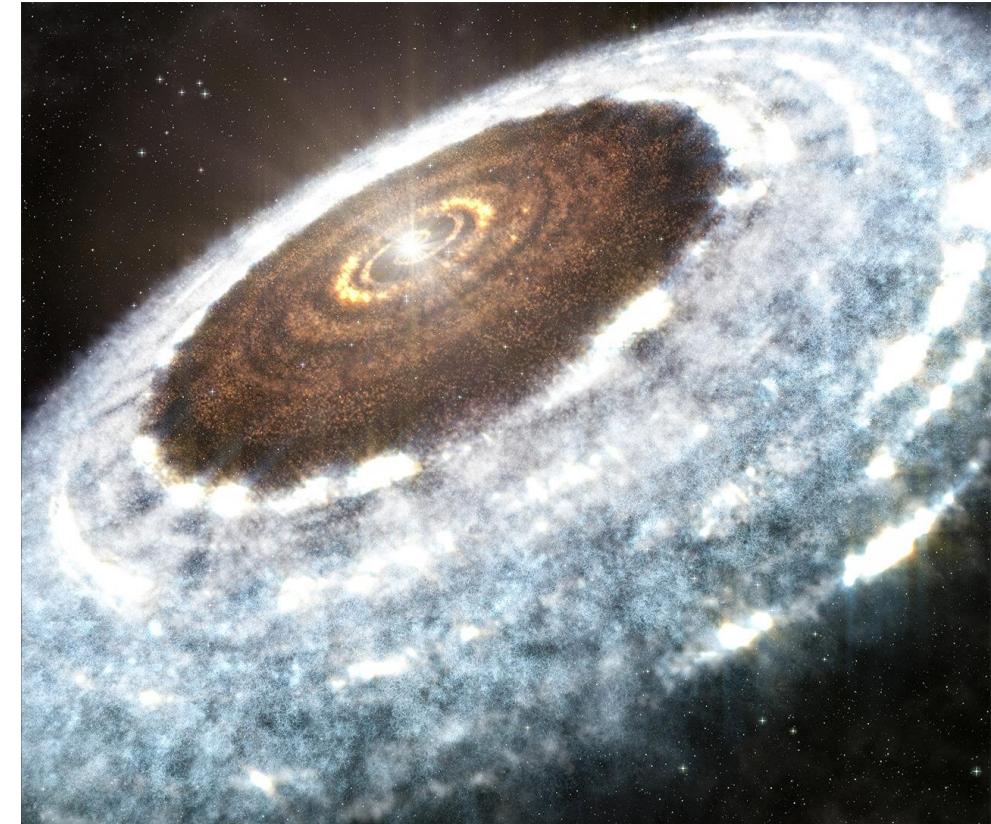


Ice



Gas

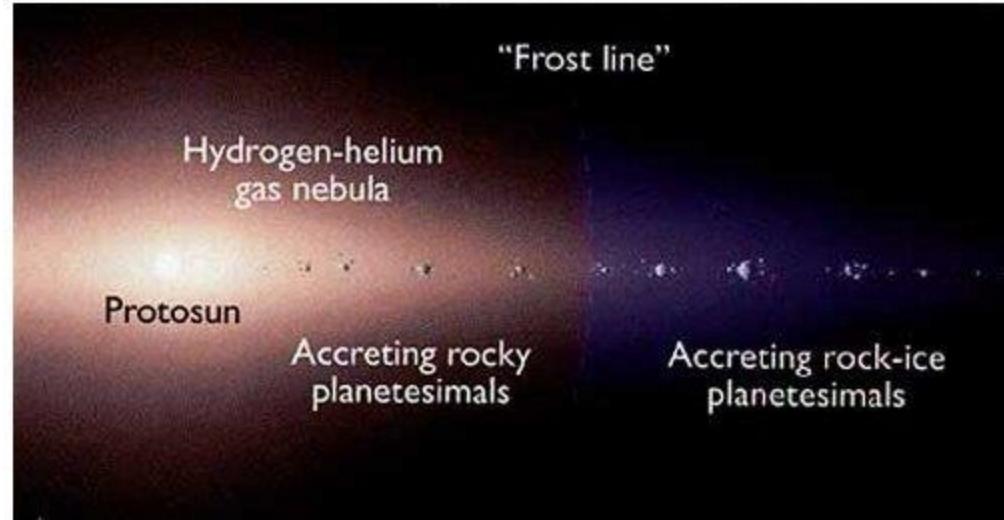
Snowline



The idea, *roughly*

Inward of snowline

Rocks only
(small)



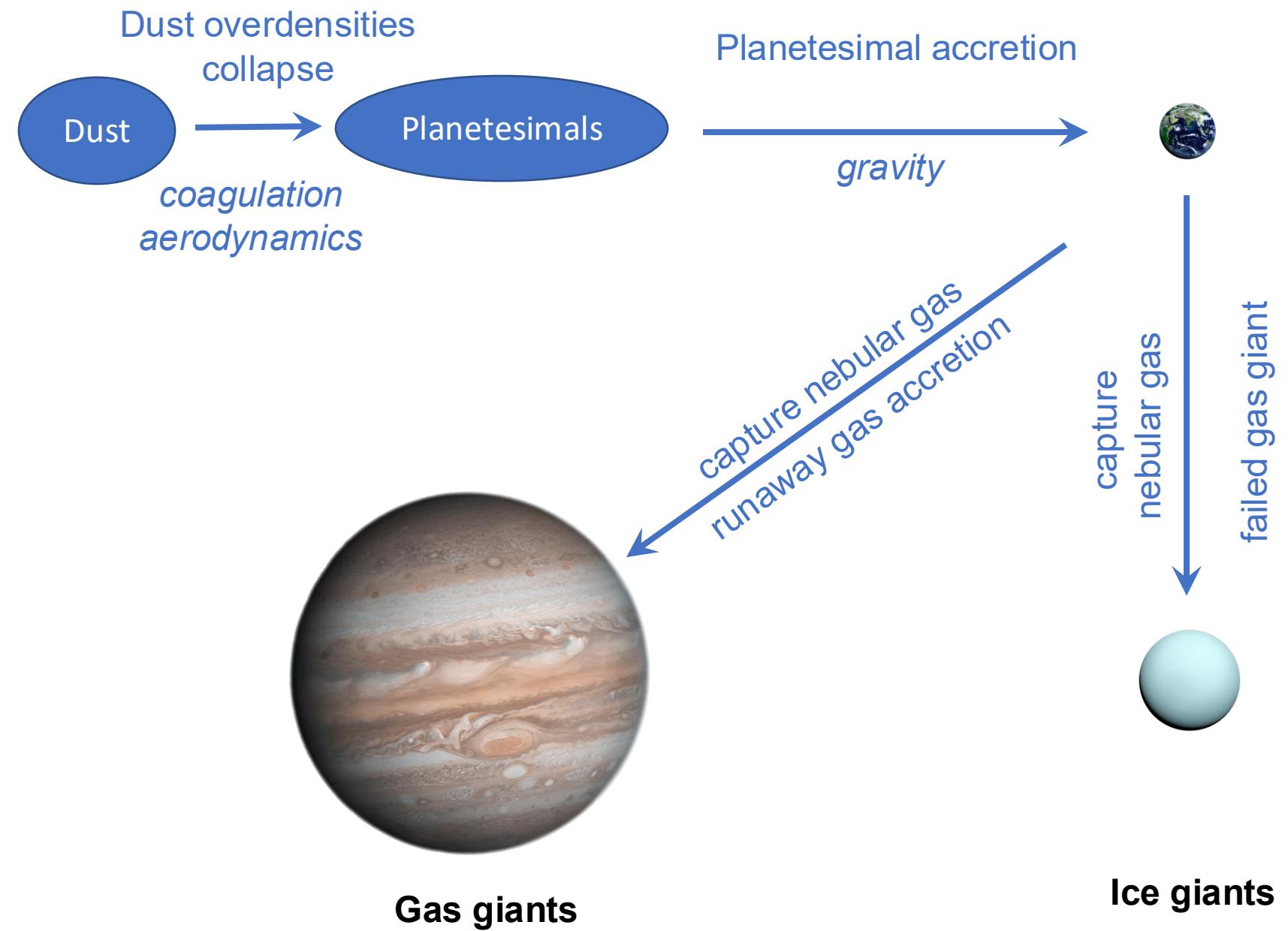
Outward of snowline

Ice comes to aid!
Growing big
icy/rocky cores.



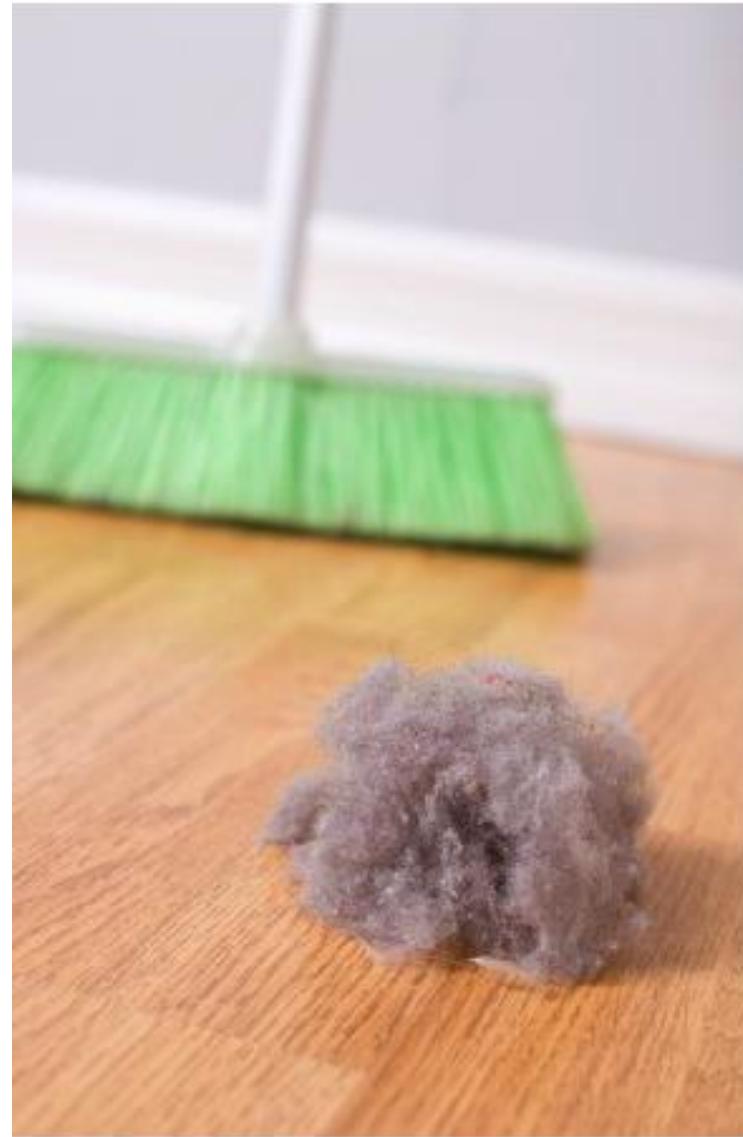
“Core Accretion”

Core
Accretion



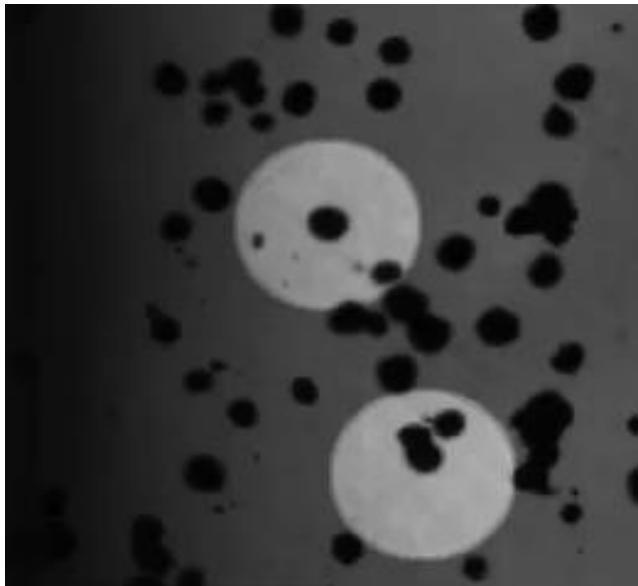


“Bottom-Up”: Dust Growth and Core Accretion

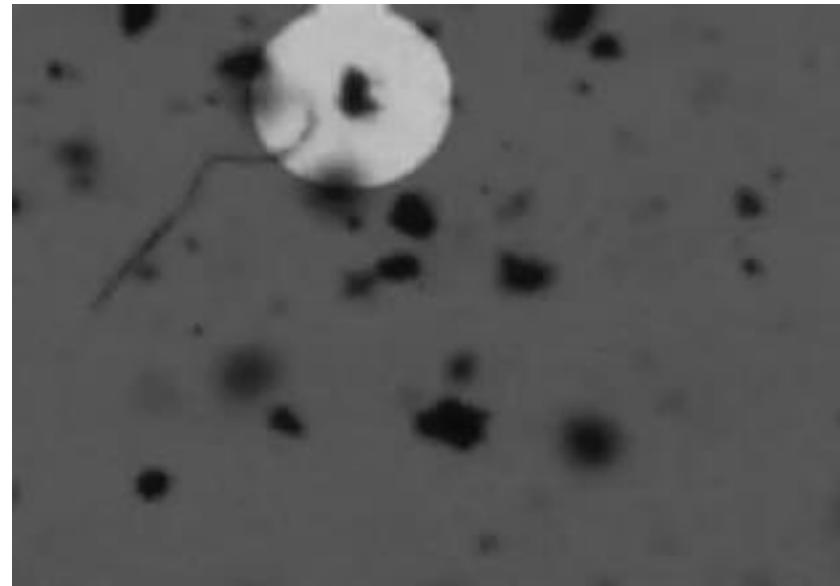


Grain collision outcomes

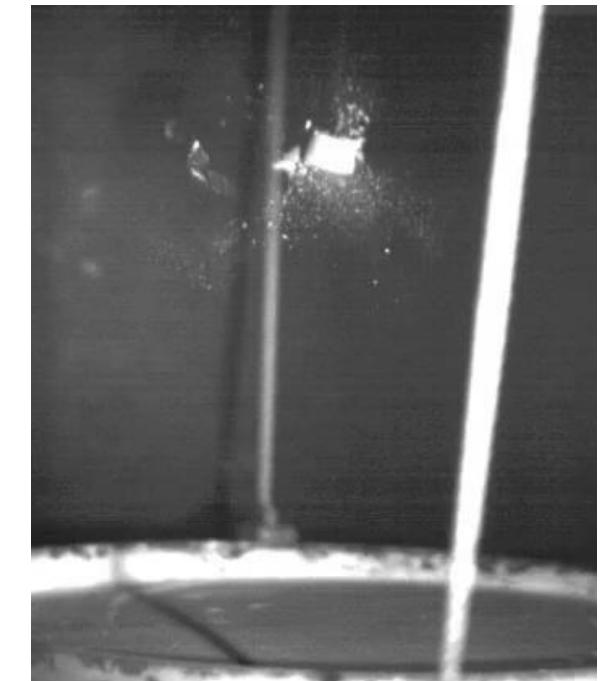
Bouncing



Sticking



Destruction



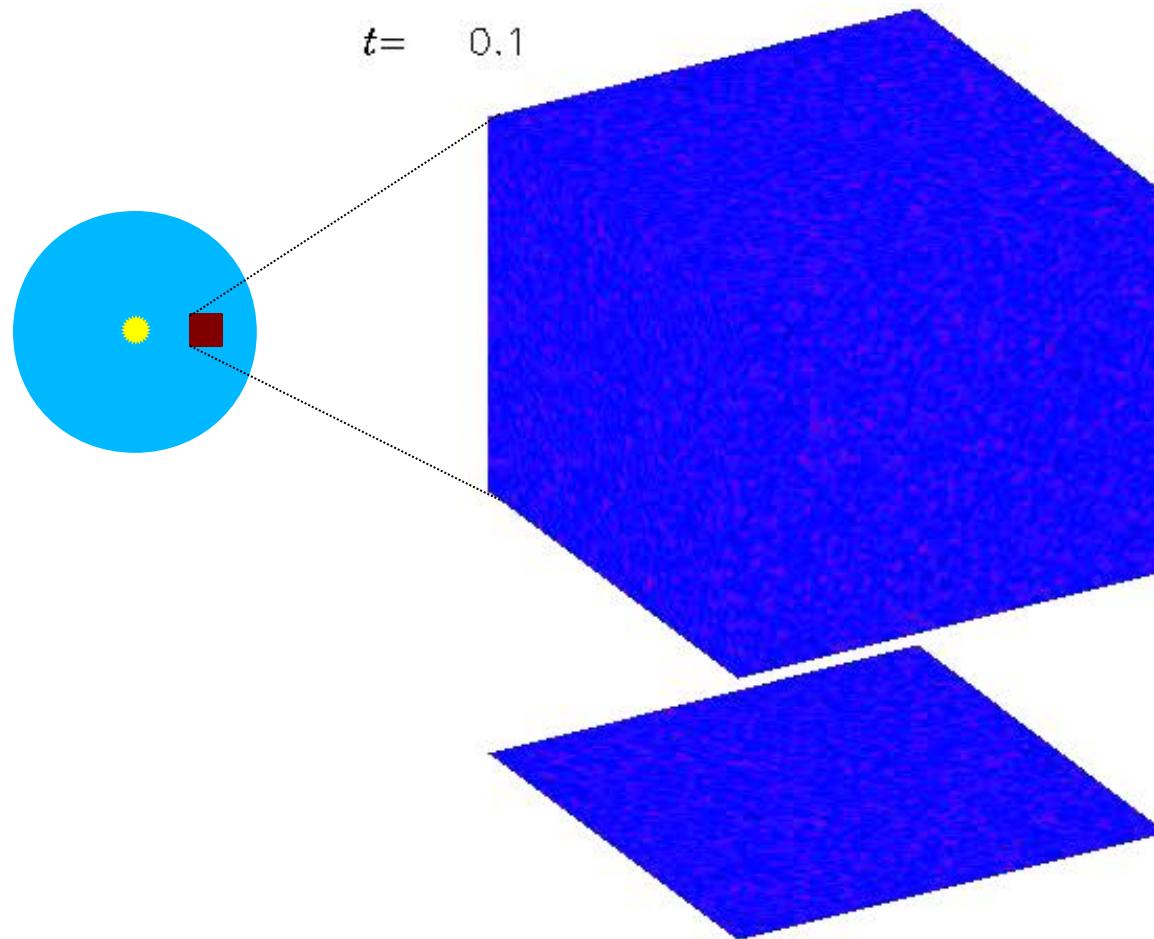
Dust Growth and Drift

Dust particle
coagulation
and radial drift

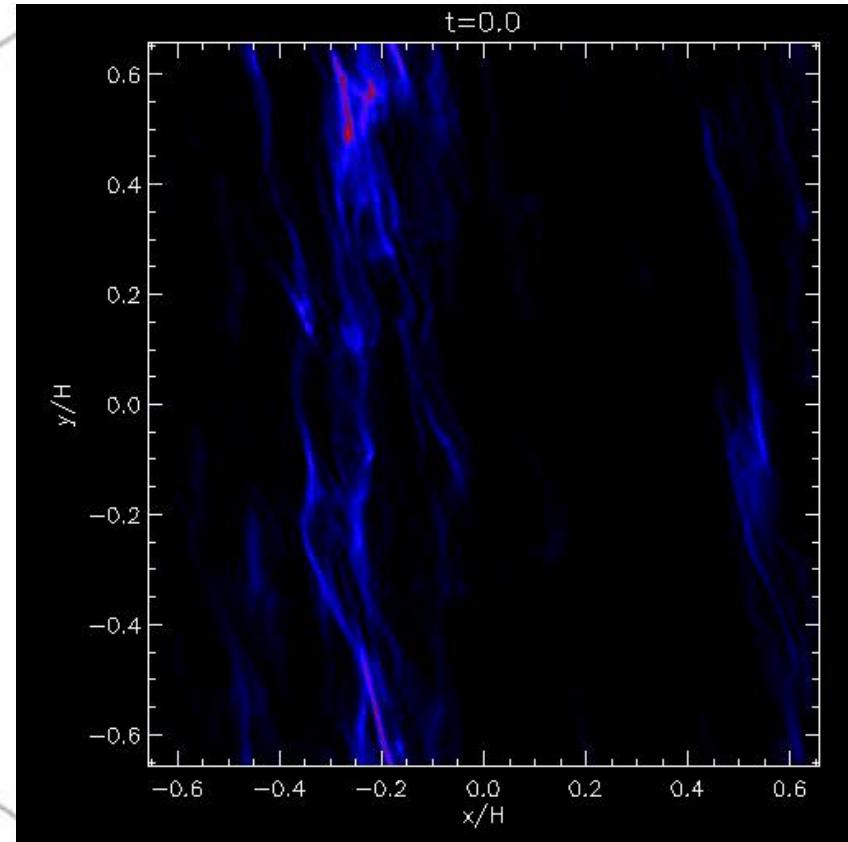
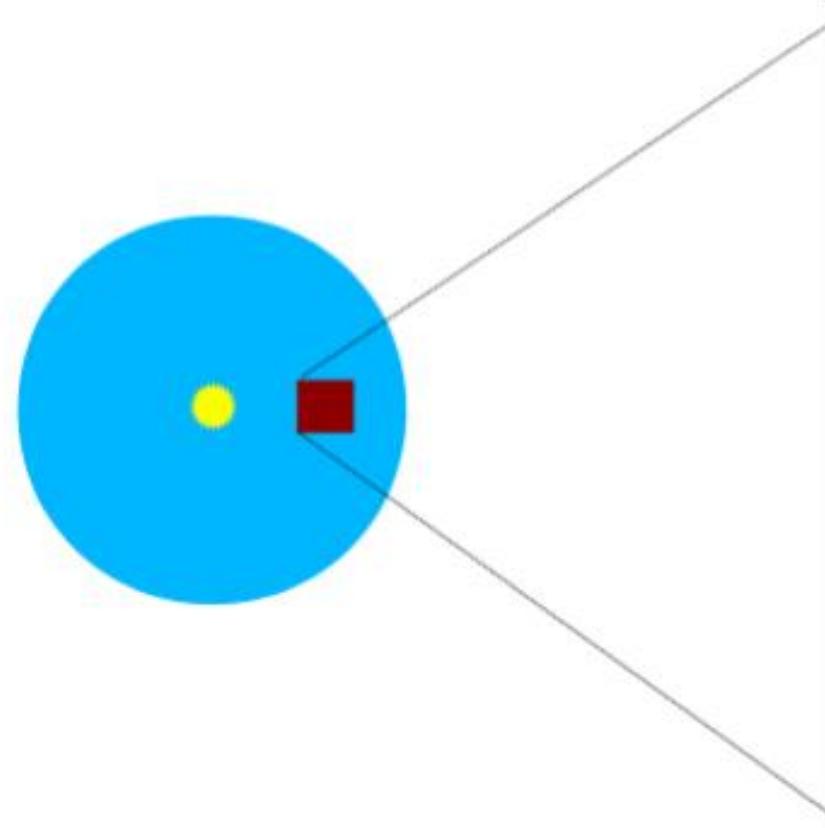
F. Brauer, C.P. Dullemond
Th. Henning

“Streaming” Instability

The dust drift is not orderly, but turbulent



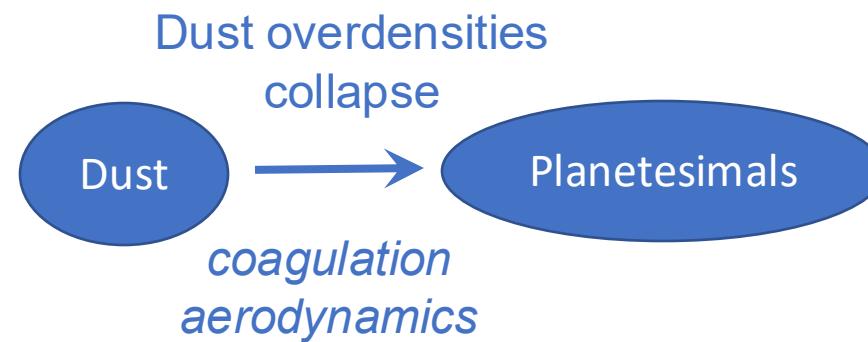
Gravitational collapse into km-sized objects (“planetesimals”)



Johansen et al. (2007)

Core Accretion (...30 yrs ago)

Core
Accretion

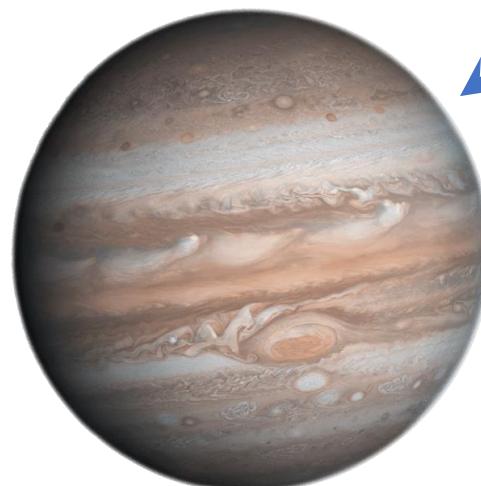


Rocky planets

Planetesimal accretion
Oligarchs
Isolation mass
gravity



capture nebular gas
runaway gas accretion



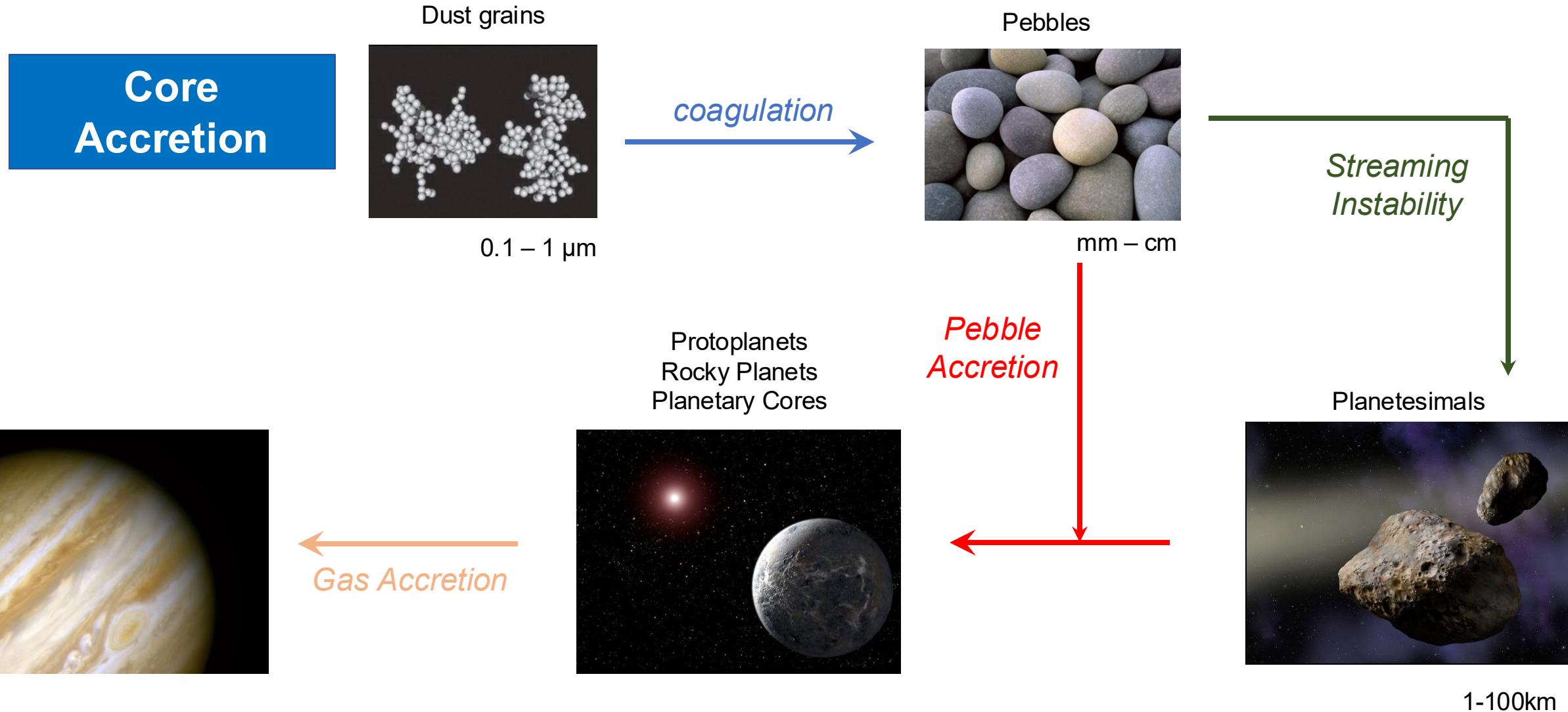
Gas giants

capture nebular gas
failed gas giant



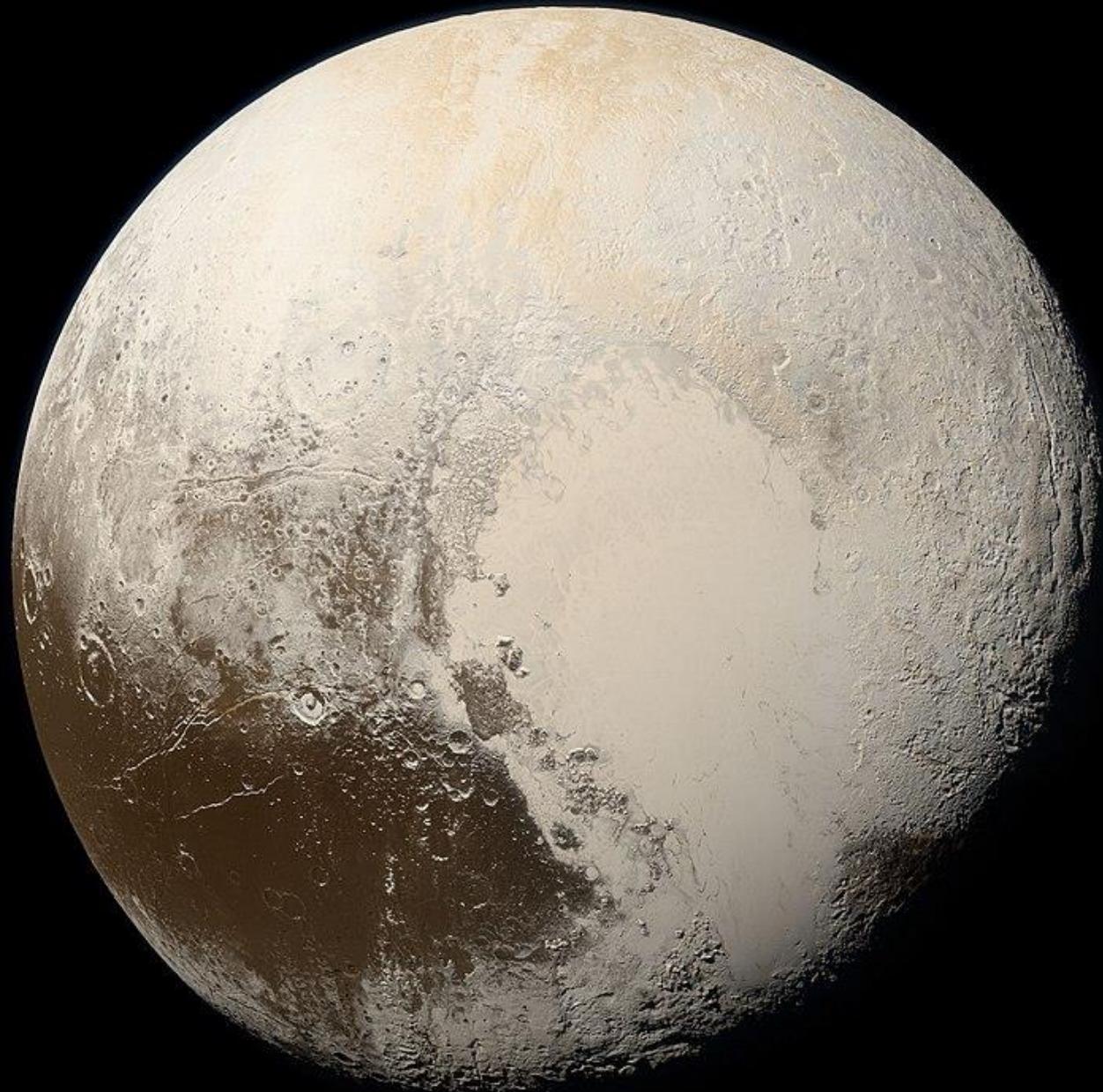
Ice giants

Core Accretion



Fingerprints of
streaming instability

**How can we verify the
streaming instability hypothesis?**



EVENT PARTNERS



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*Center for
Nonlinear Studies*

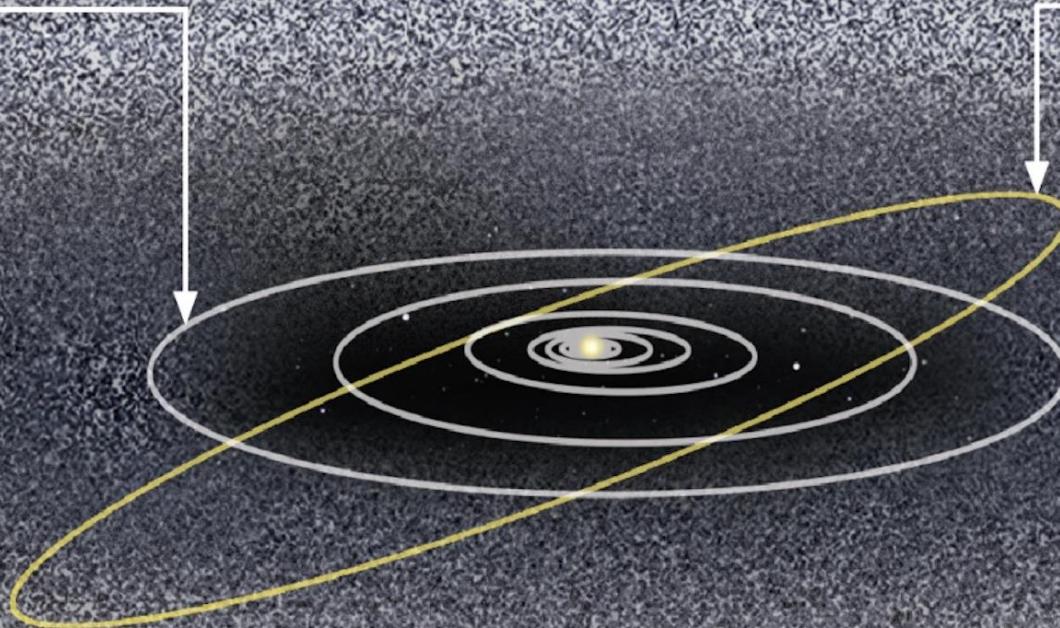


Pluto sponsor?



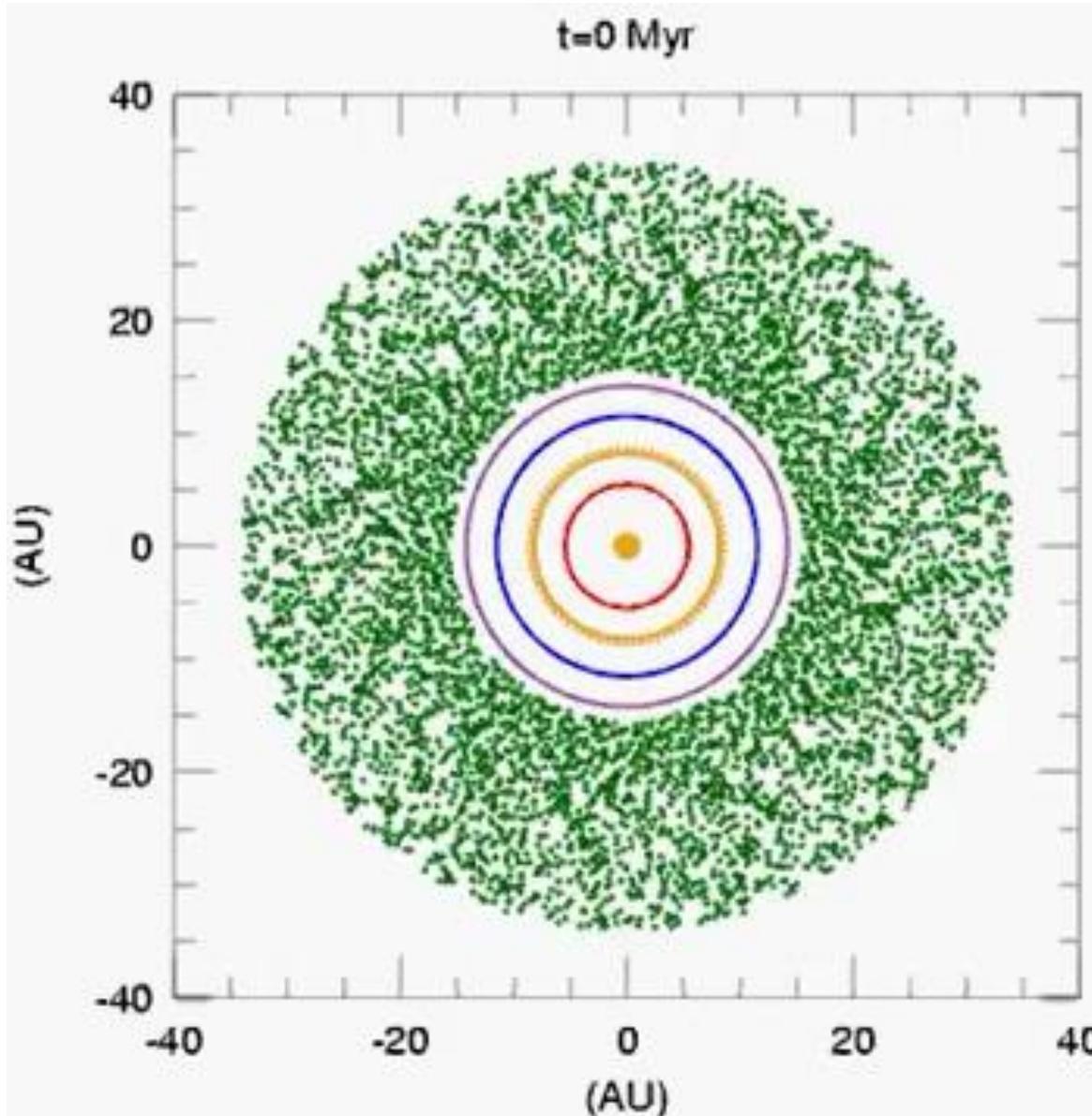
Neptune's orbit

Pluto's orbit

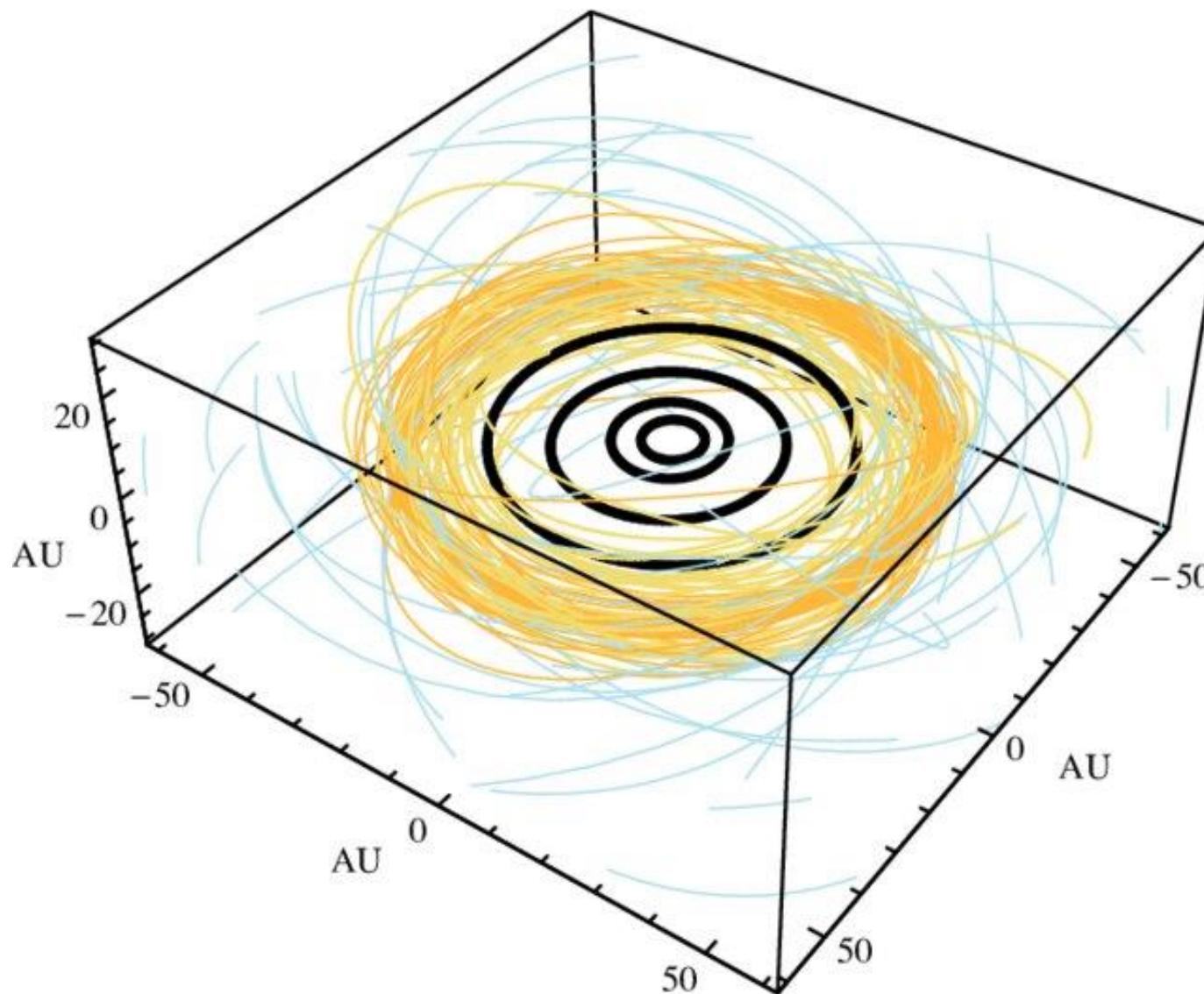


THE KUIPER BELT

Evolution of the orbits of the giant planets

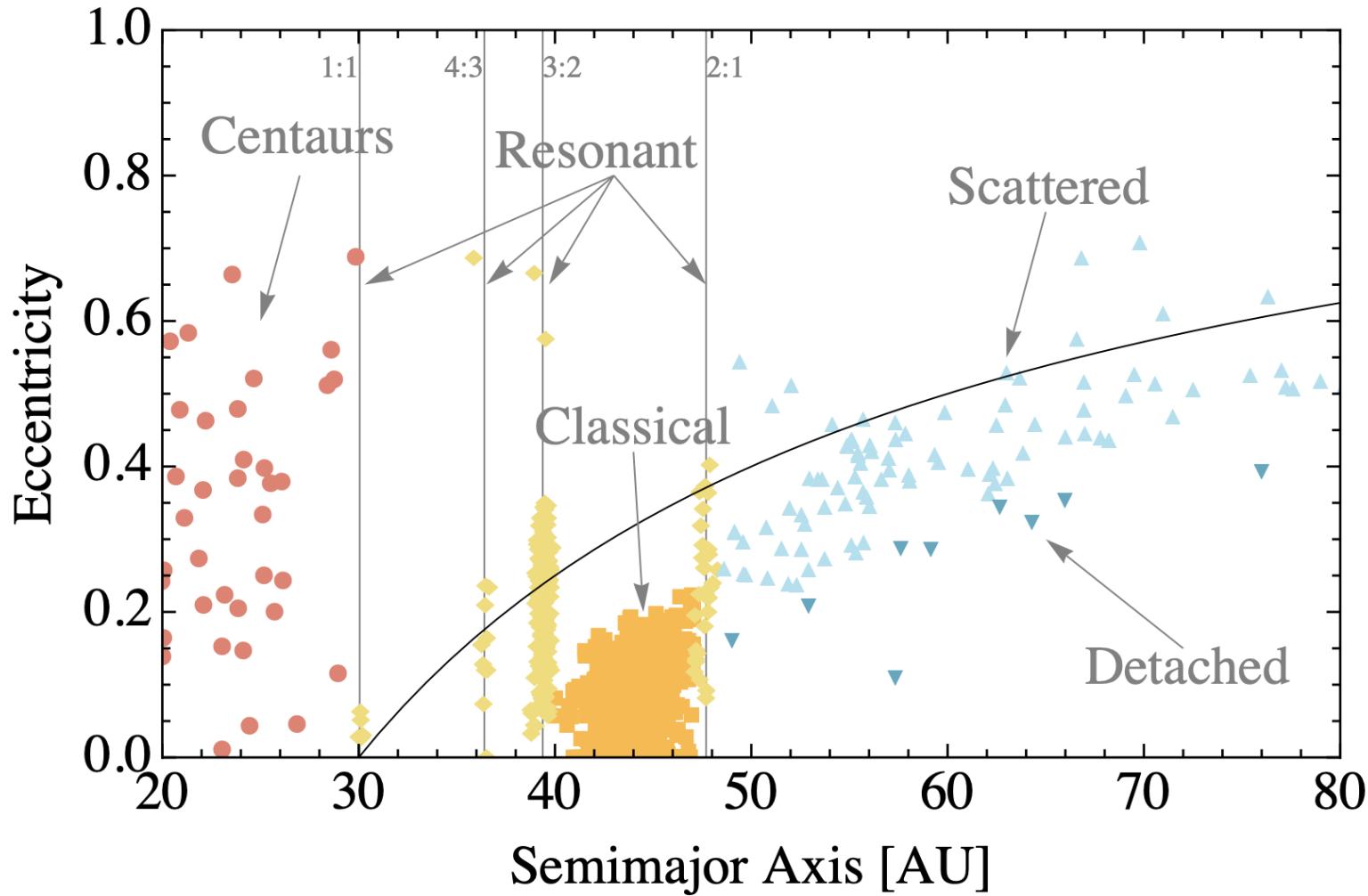


Structure of the Kuiper Belt



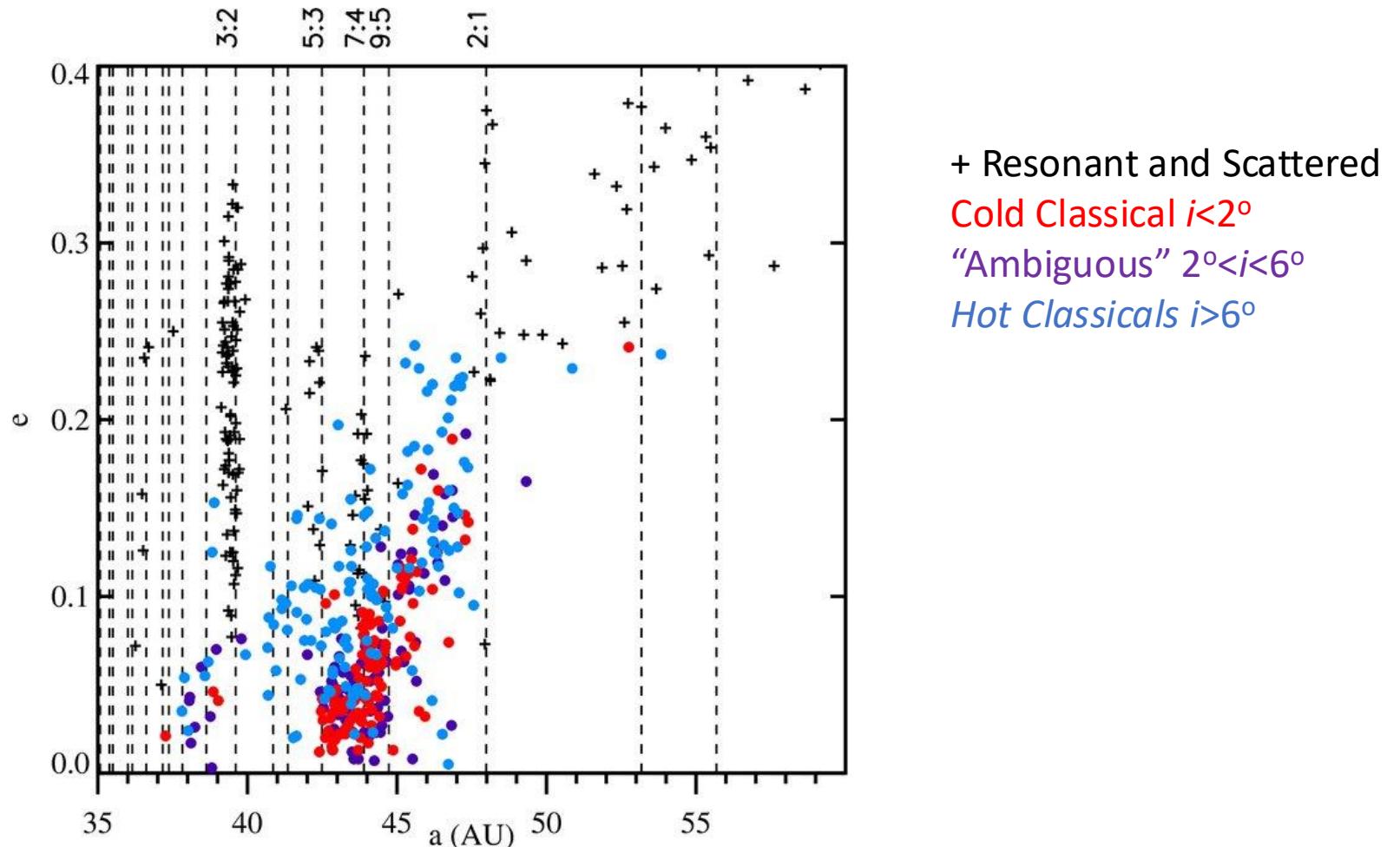
Classicals
Scattered
Resonants

Structure of the Kuiper Belt



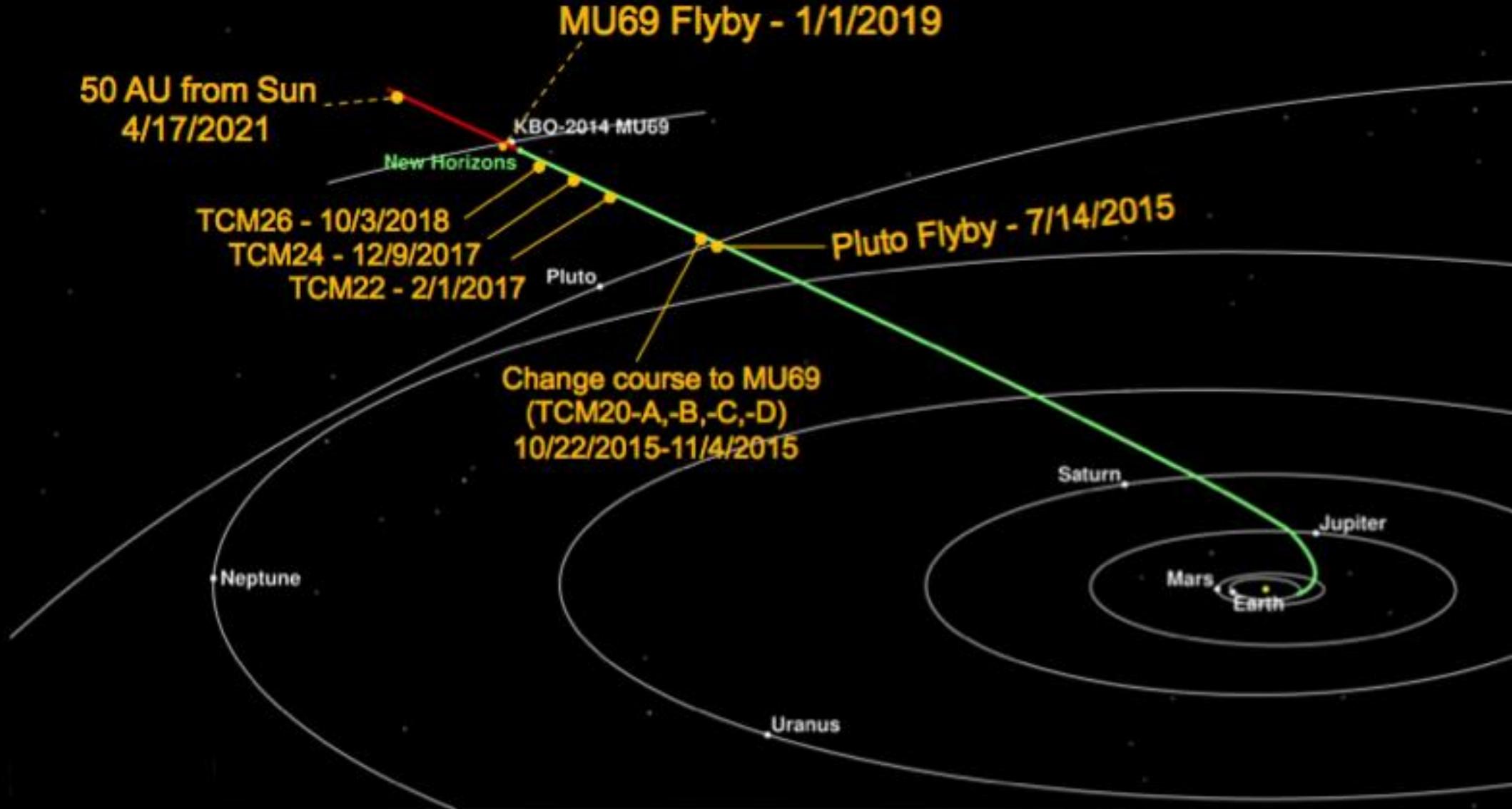
Classicals:
Objects between 40 and 50 au

“Cold” Classical Kuiper Belt Object

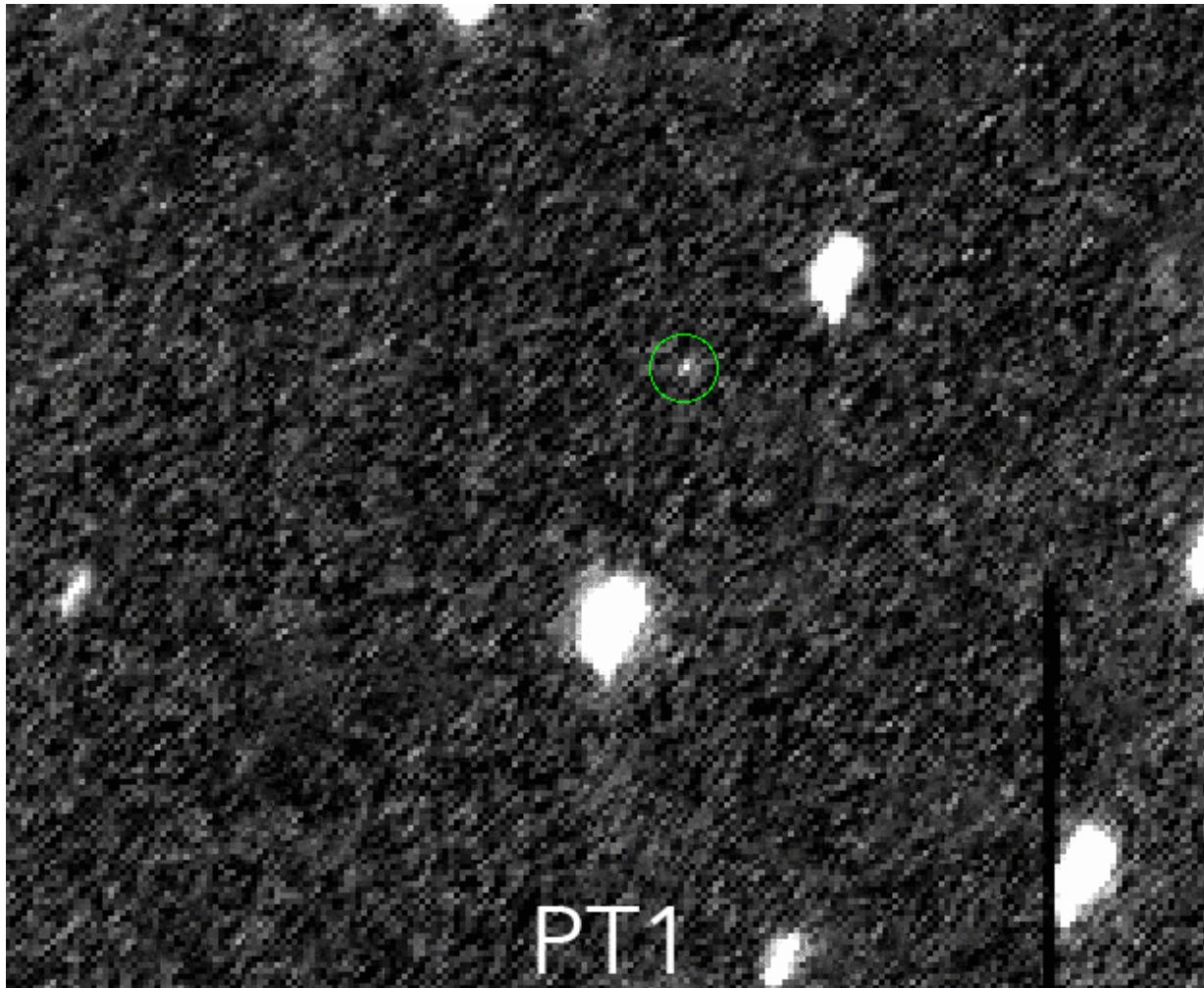


A dynamically cold population: low inclination and low eccentricity orbits, like the planets.
Presumably pristine planetesimals

New Horizons Trajectory

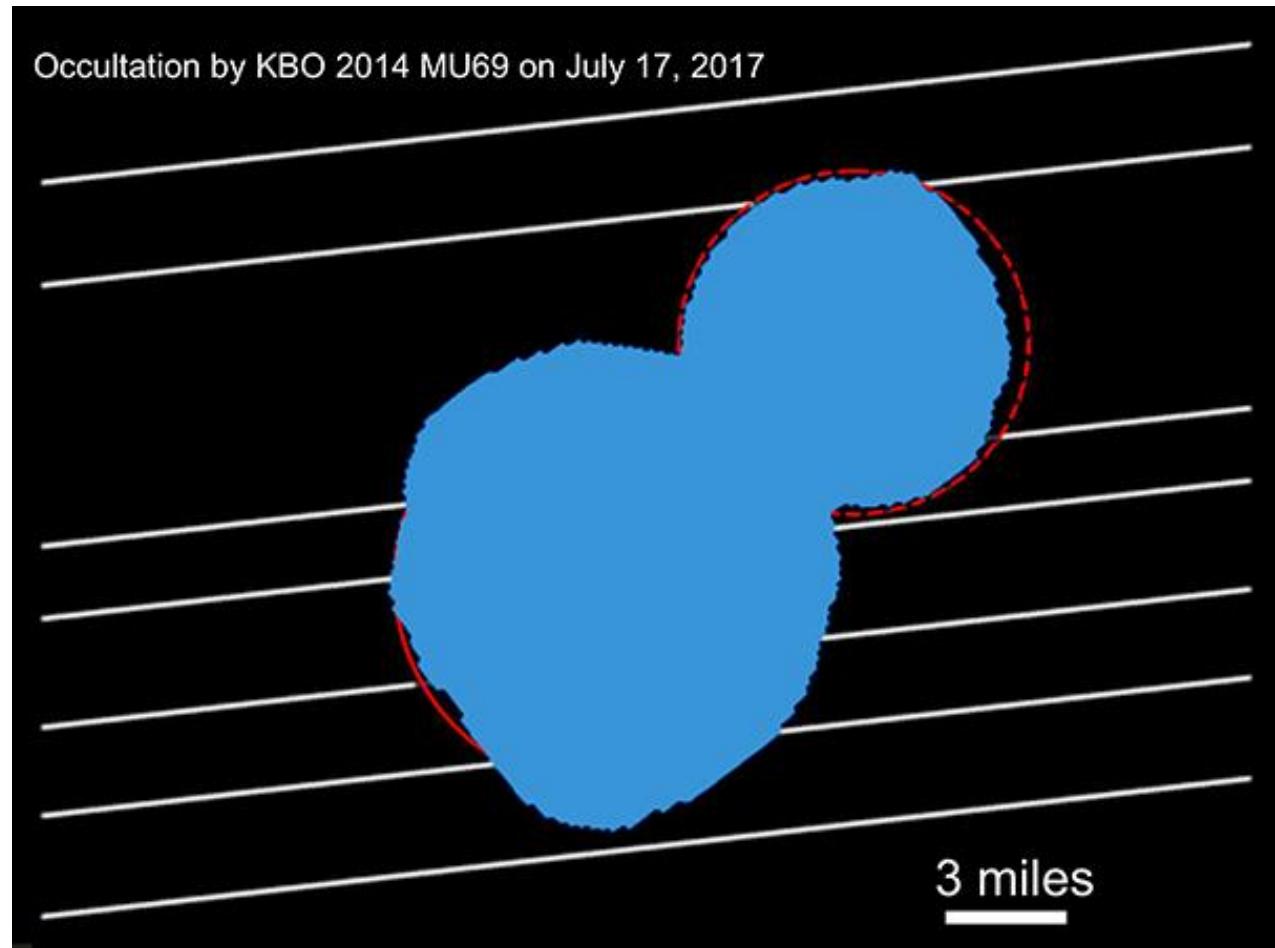
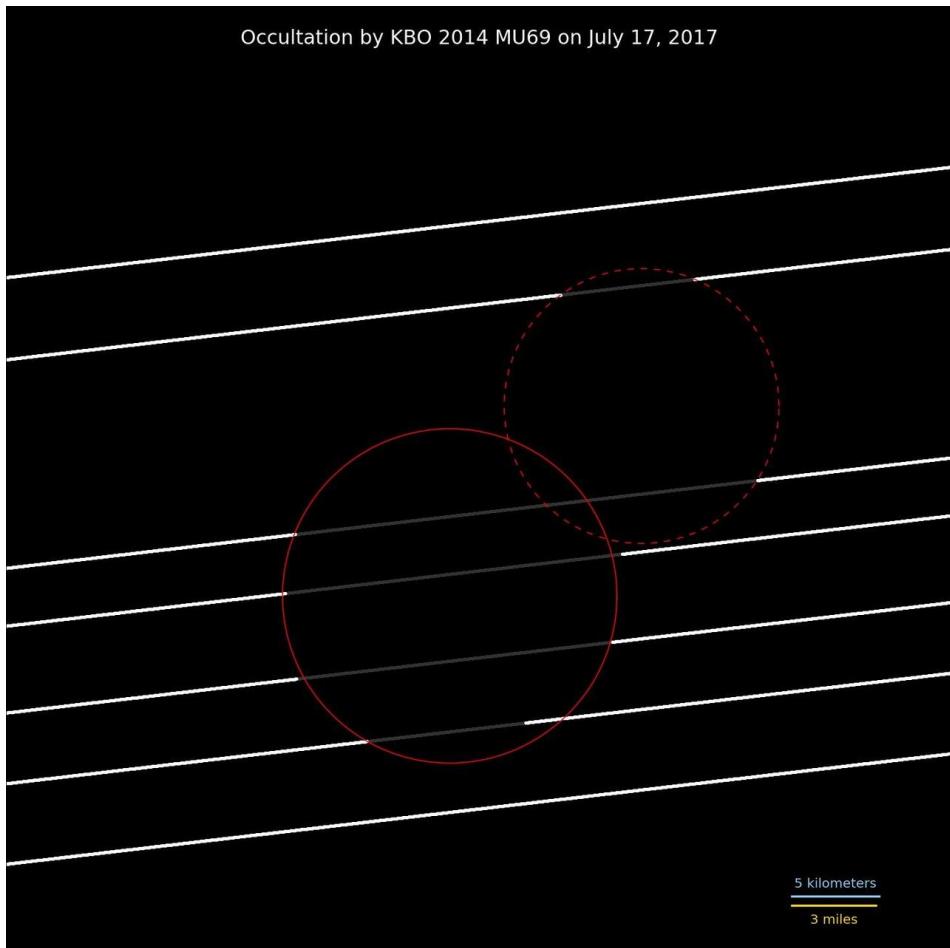


Arrokoth: Discovery

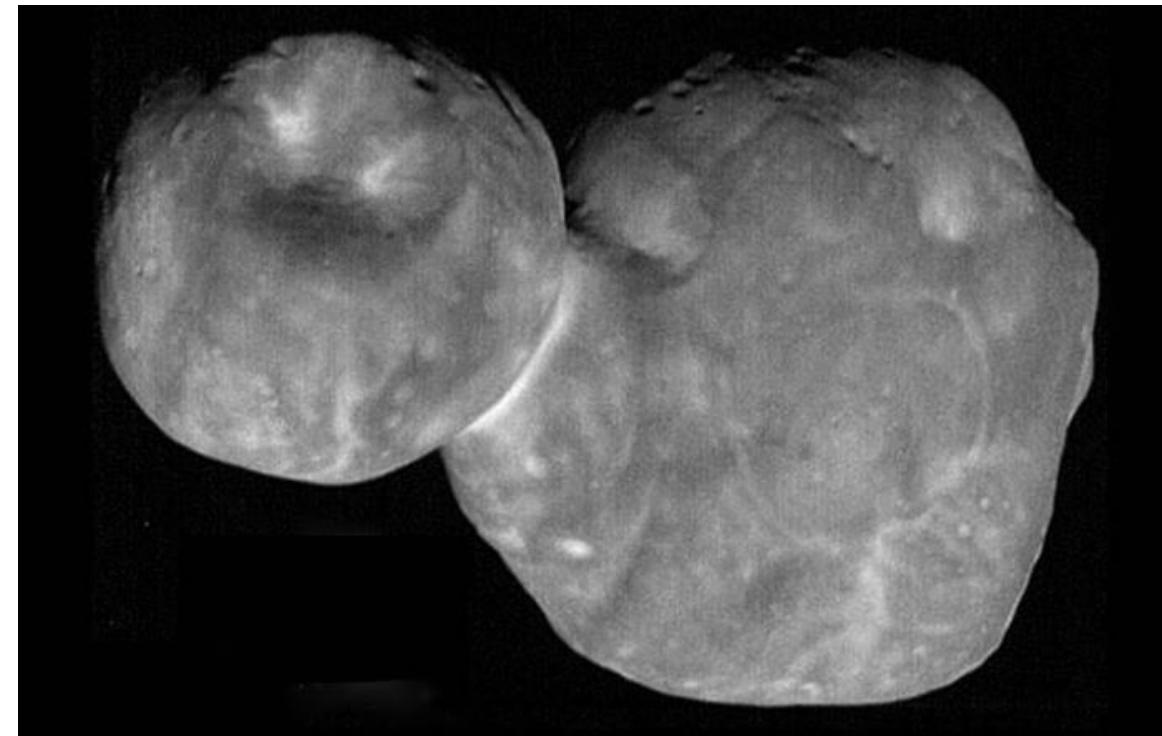
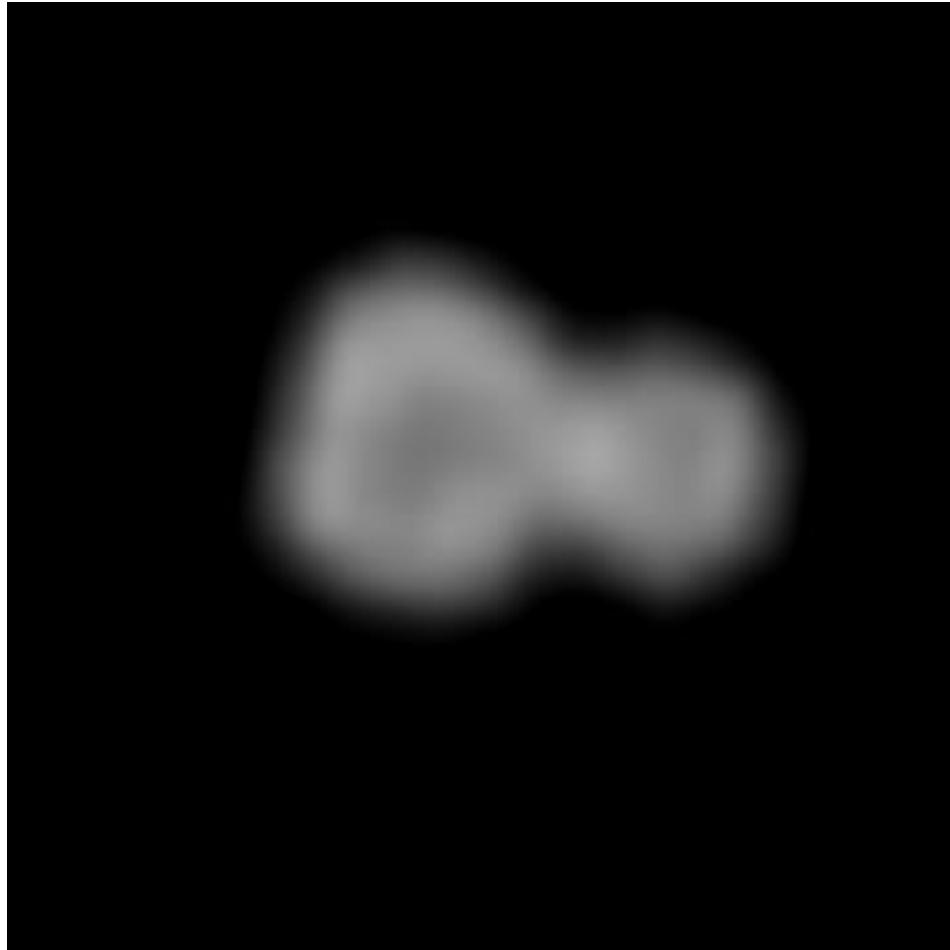


Discovery: Marc Buie. Video credit: Bruce Murray

Occultation data suggests binary



Approach sequence: Contact Binary

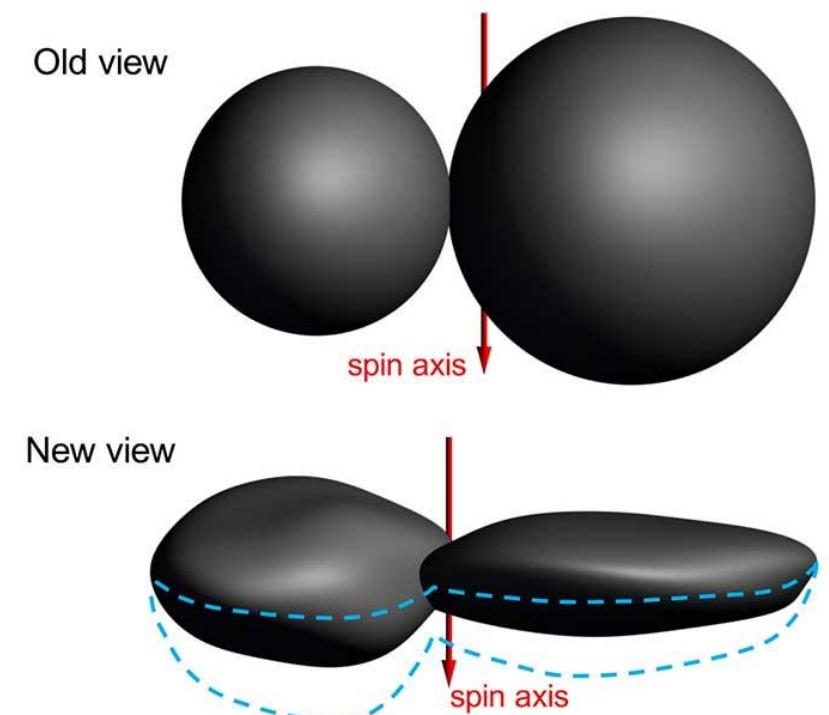
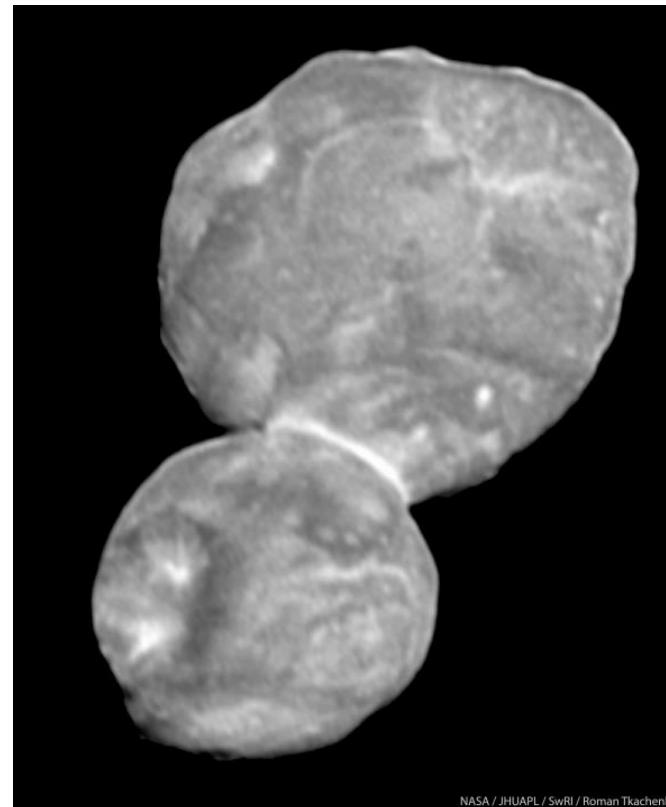


“The Snowman”



↑
21 miles
(33 km)
↓

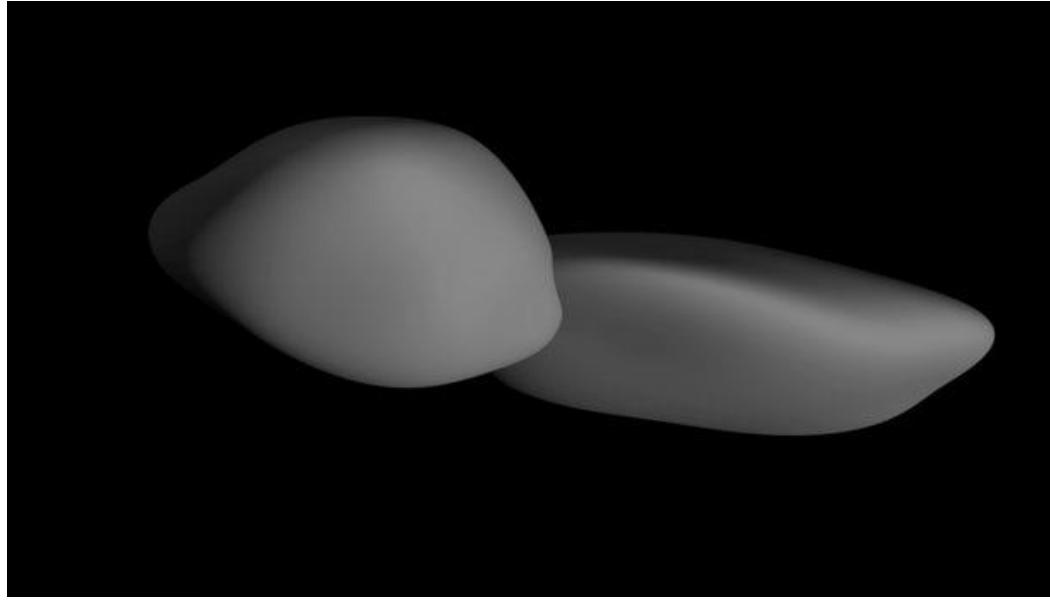
Departure sequence: Shape



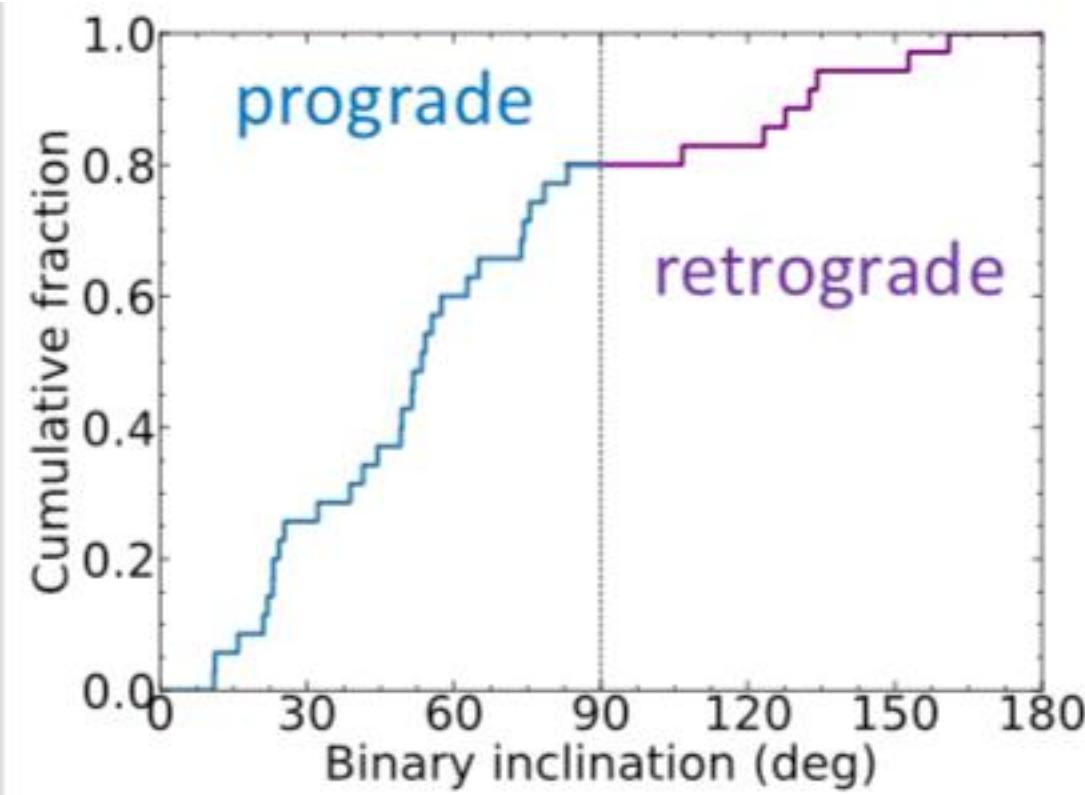
“The Snowman”



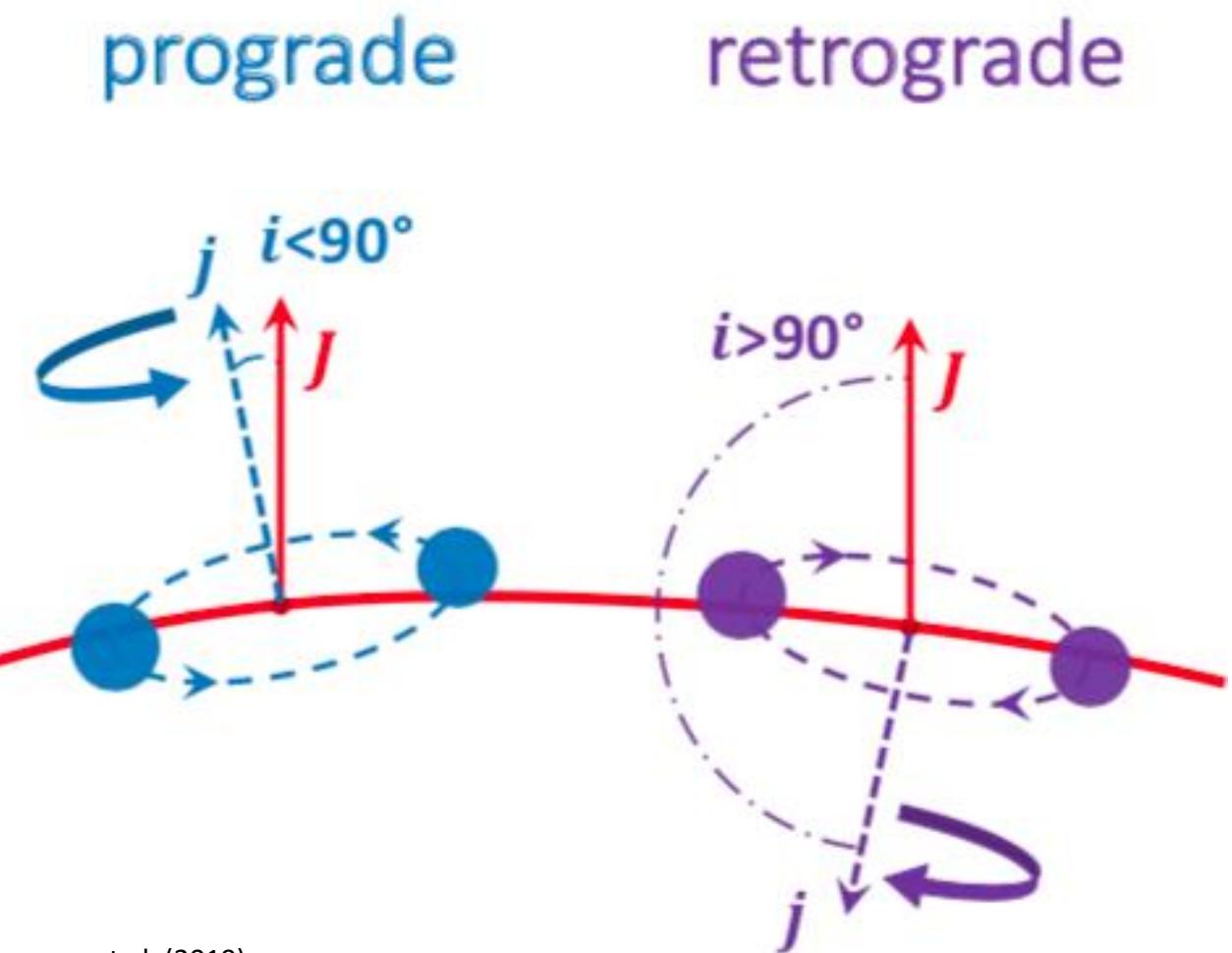
↑
21 miles
(33 km)
↓



Preference for Prograde

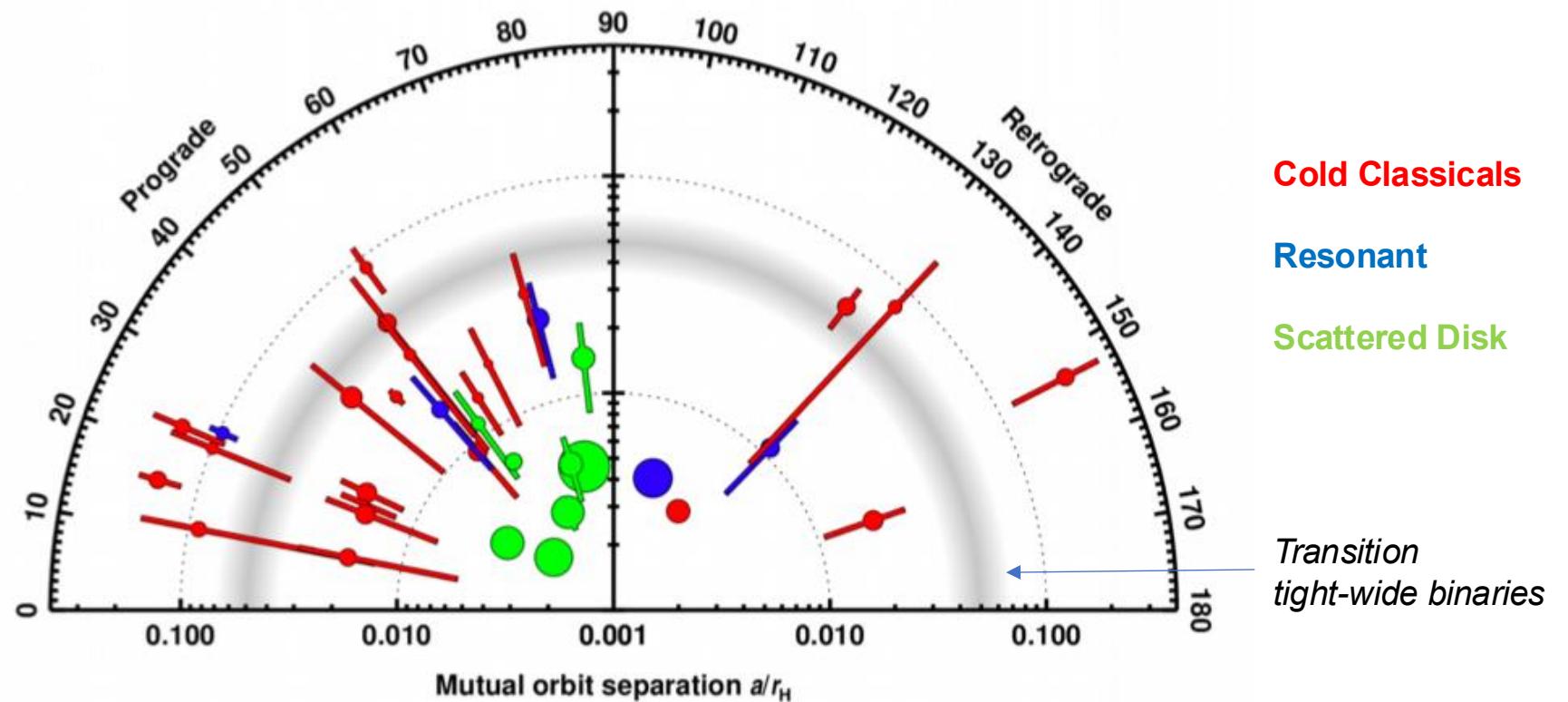


orbiting around the Sun

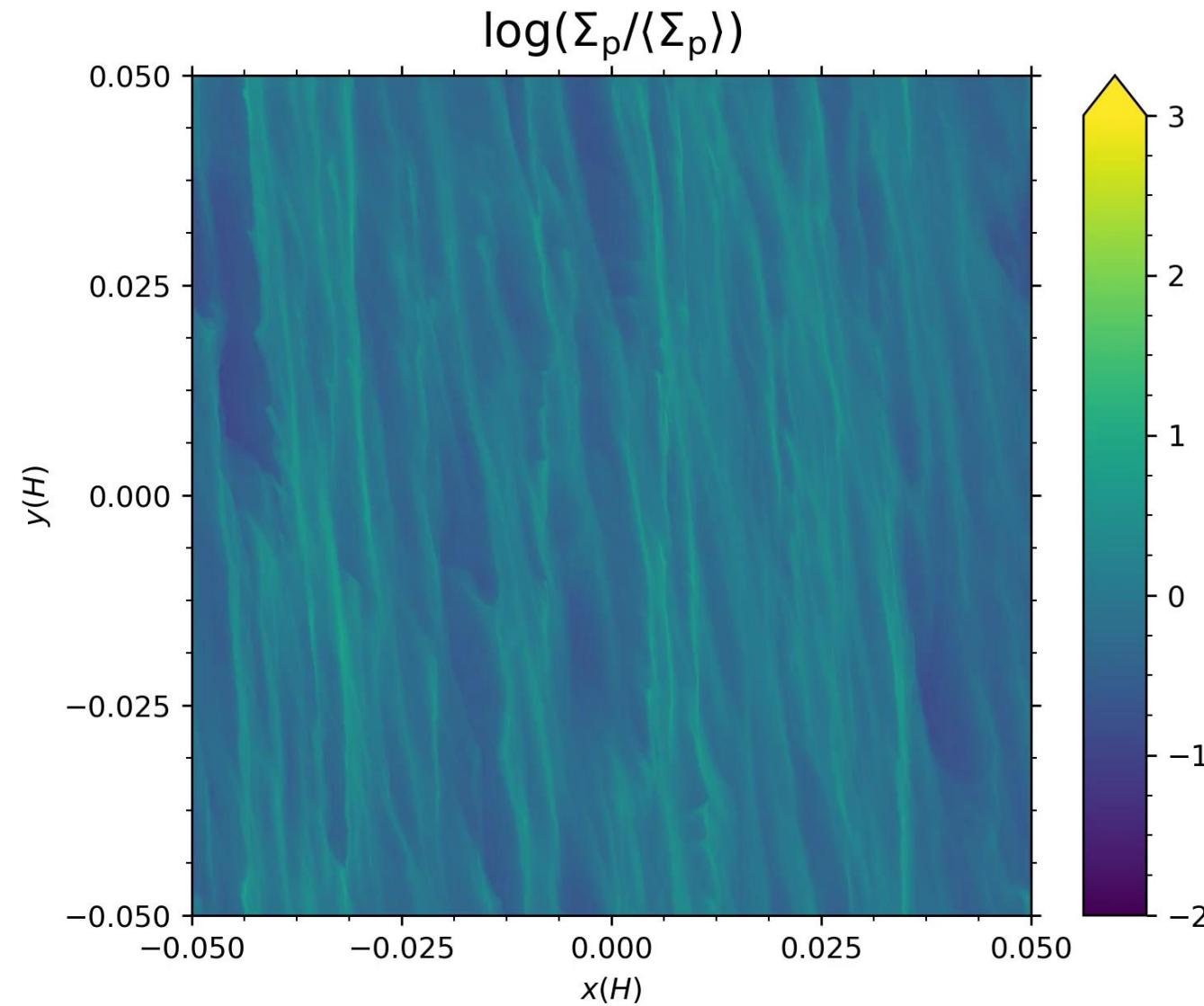


Prograde vs Retrograde

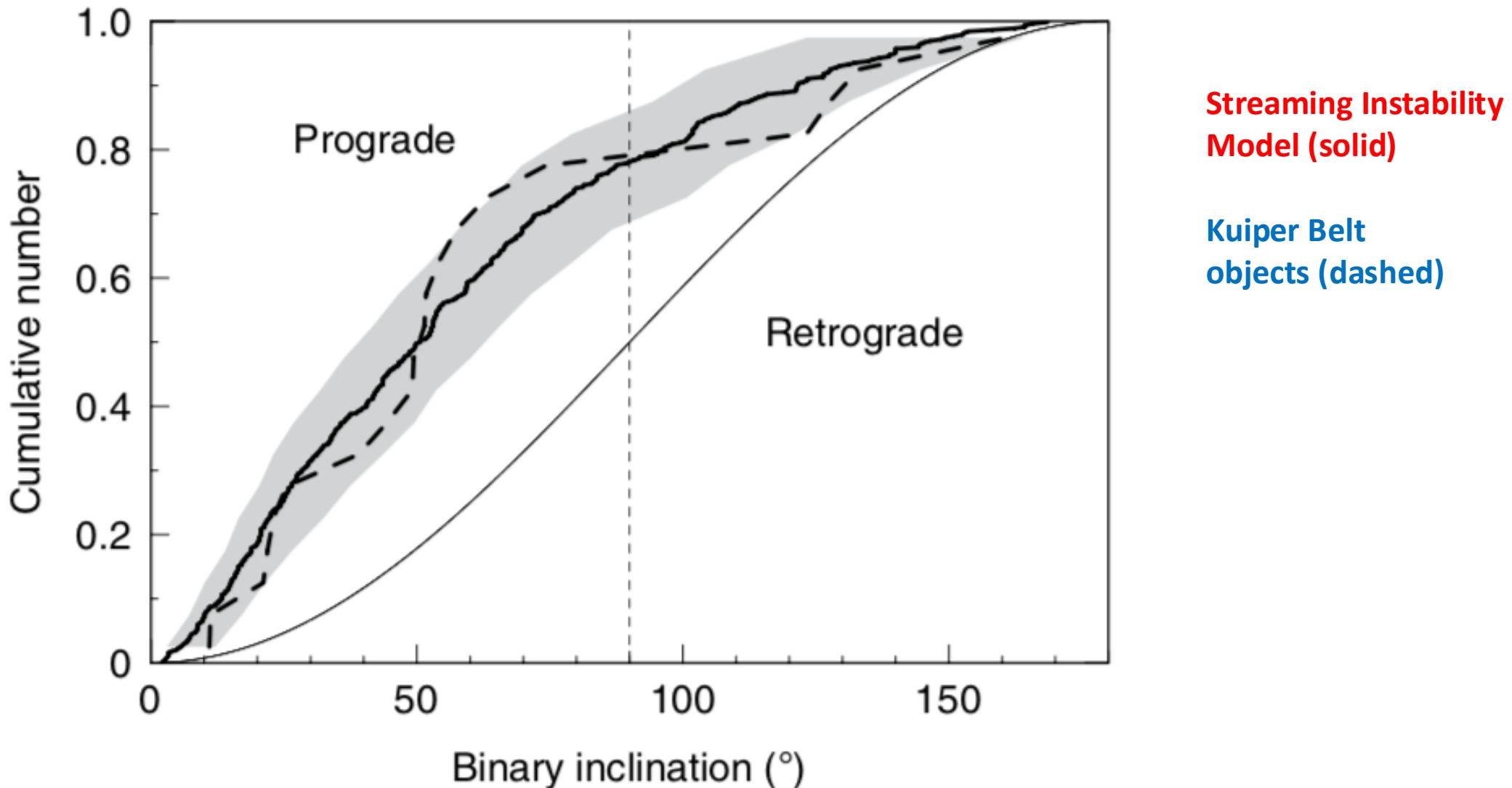
Protractor Plot



Finding binaries

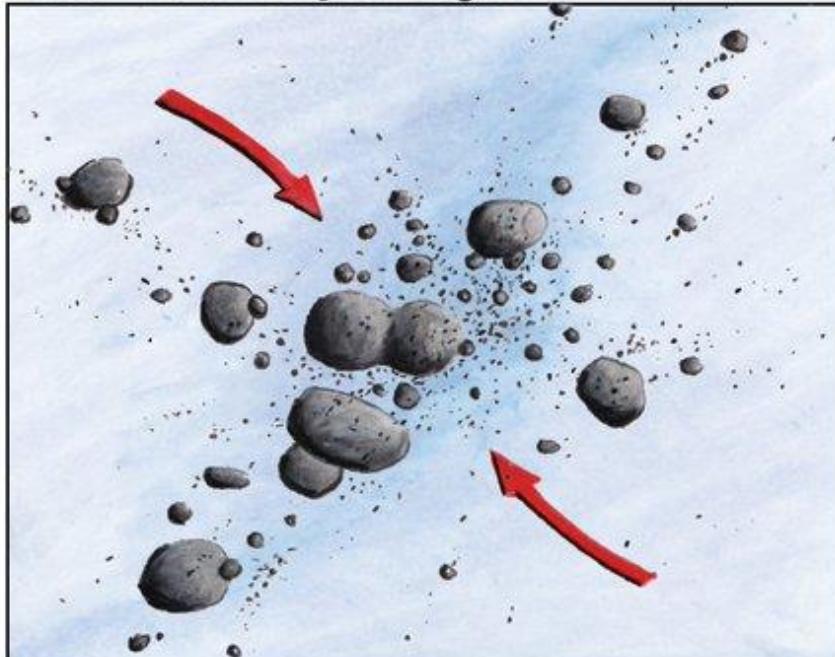


Preference for Prograde (~80%)



The Formation of Arrokoth

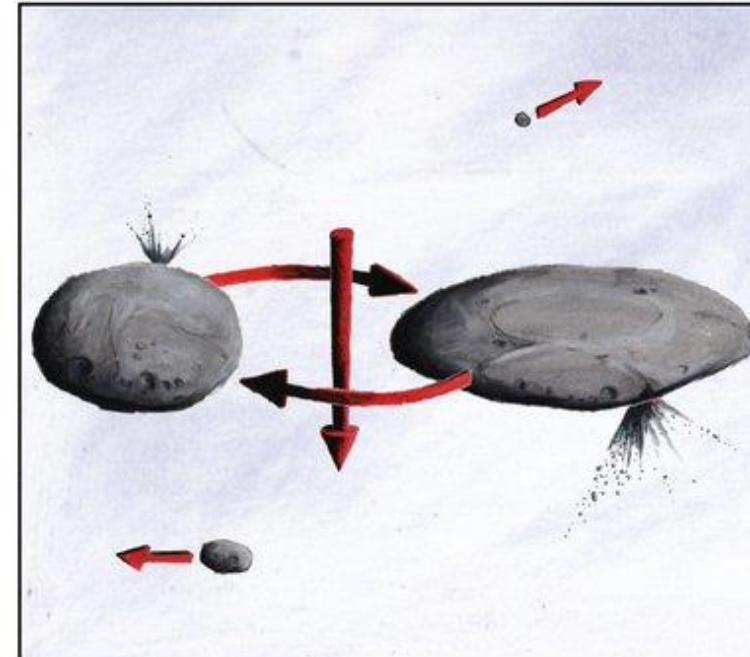
About 4.5 billion years ago...



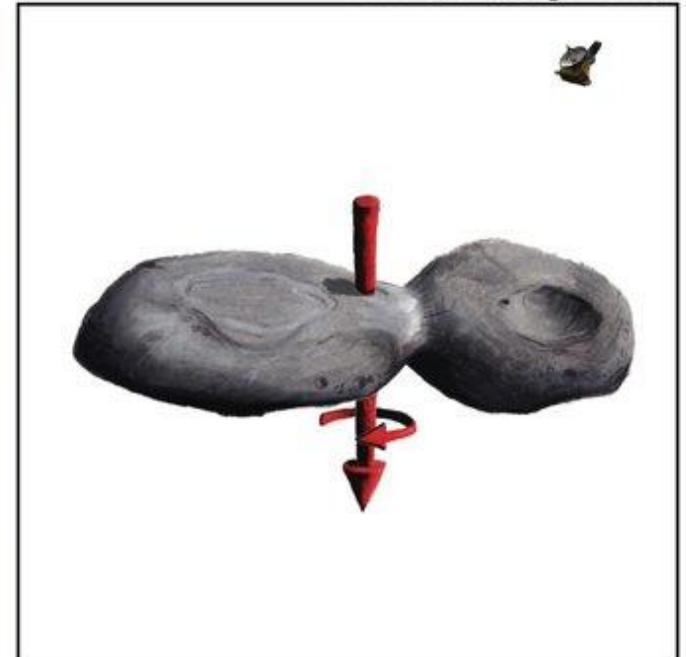
A rotating cloud of small, icy bodies starts to coalesce in the outer solar system.

 New Horizons / NASA / JHUAPL / SwRI / James Tuttle Keane

...1 January 2019.



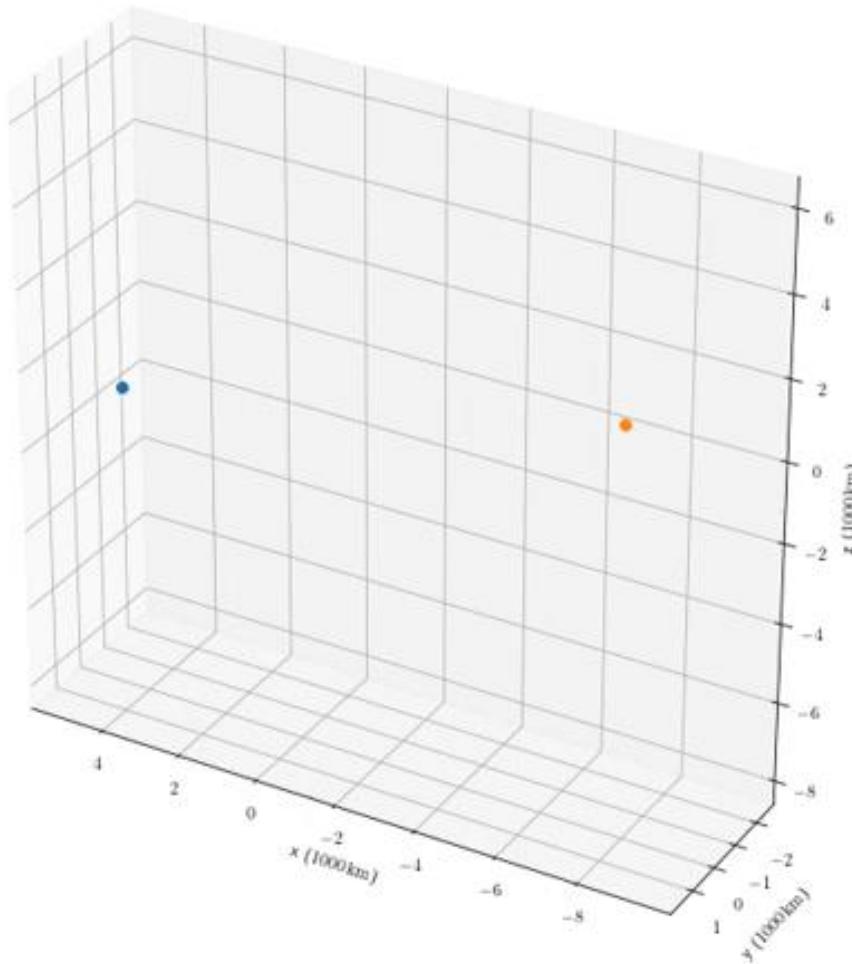
Eventually two larger bodies remain.



The two bodies slowly spiral closer until they touch, forming the bi-lobed object we see today.

time = -1.3 kyr

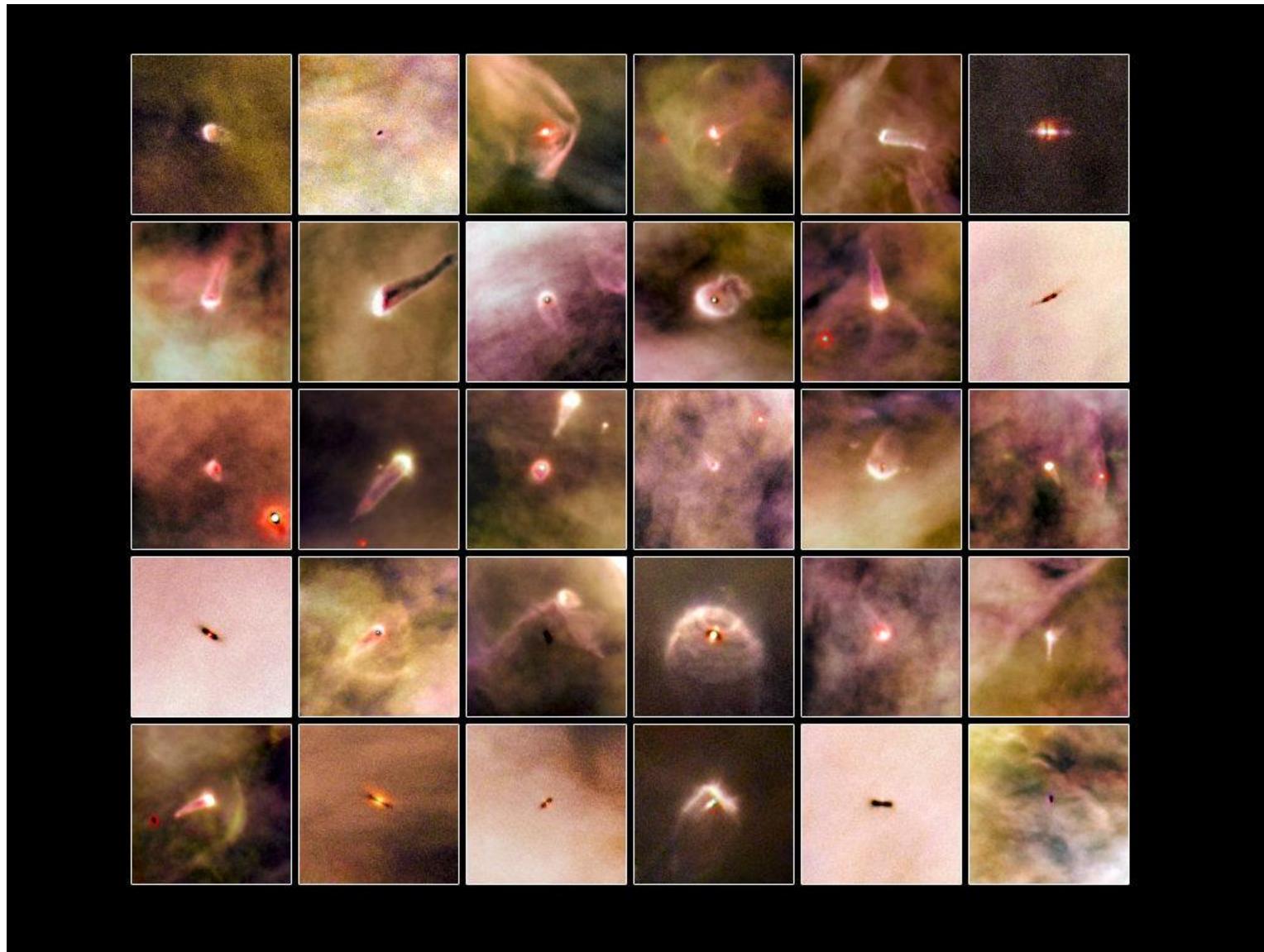
Wenu - Weeyo



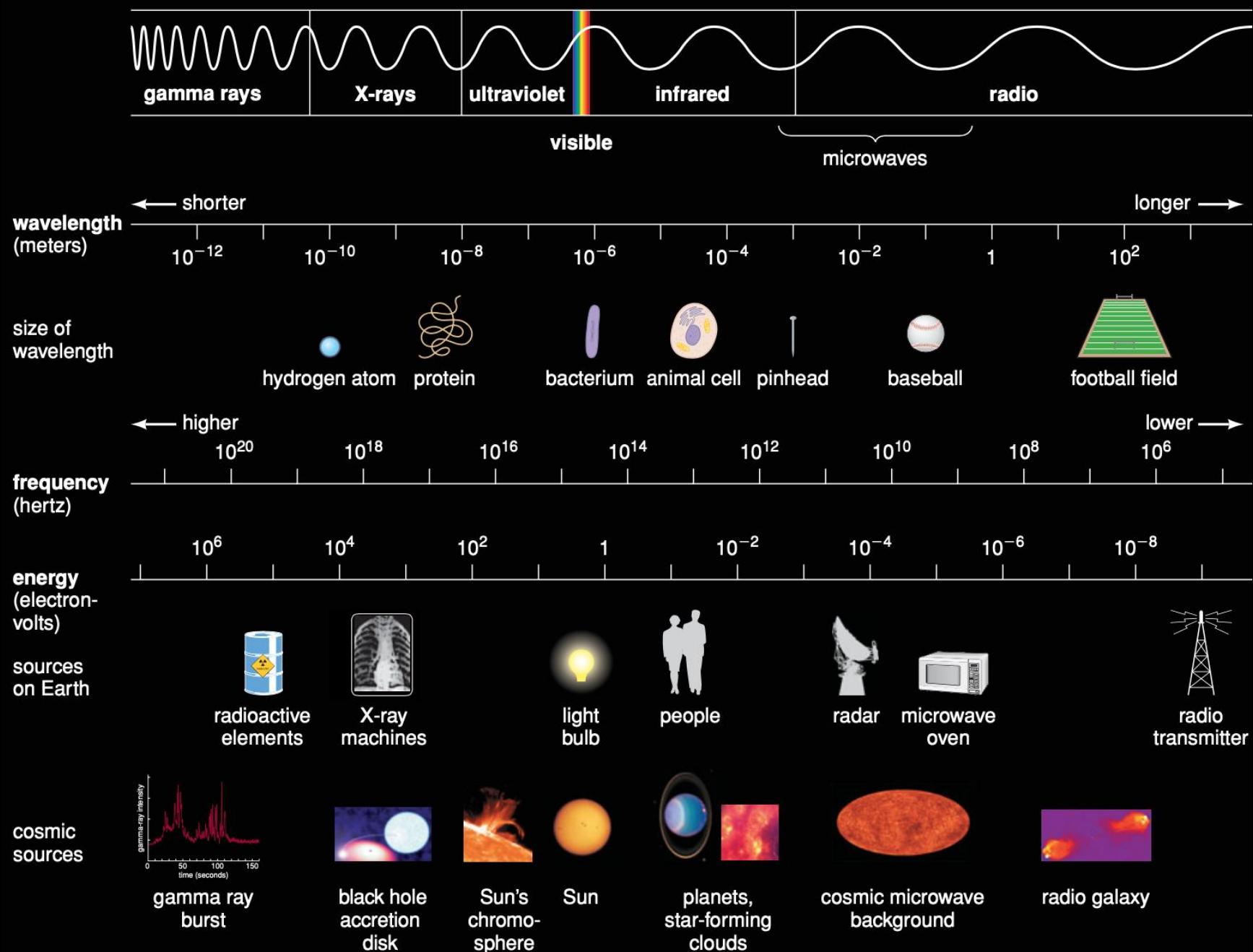
© Alexander Heger (2023)



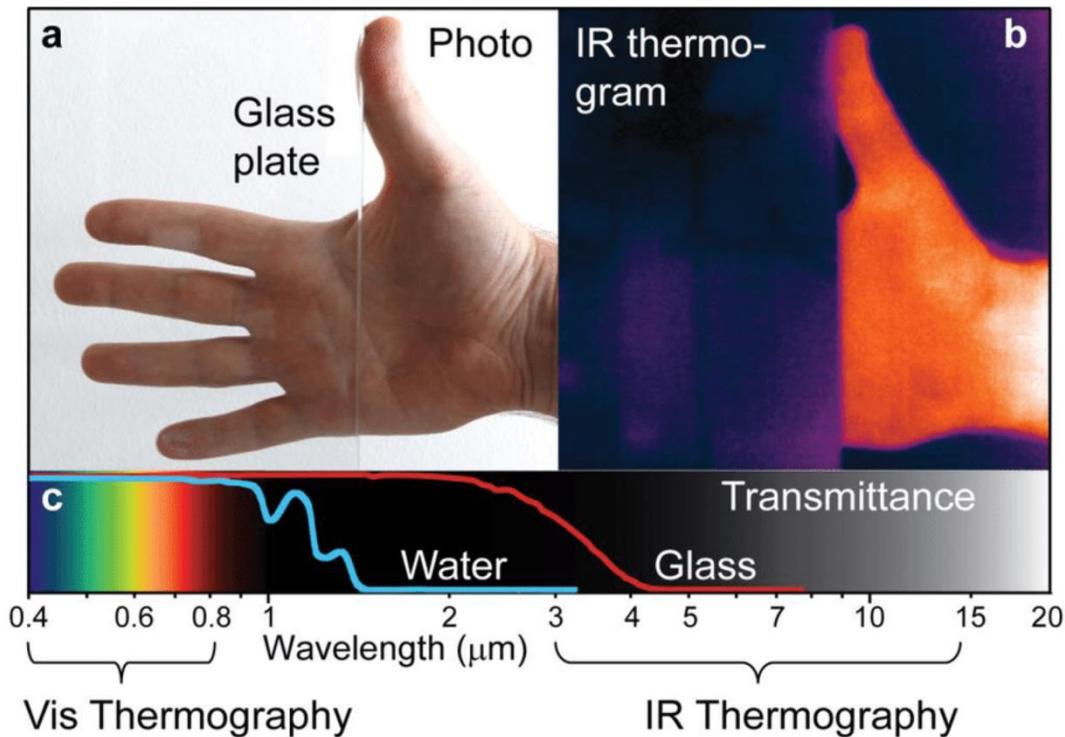
The *Hubble Space Telescope* view of disks in Orion



The Electromagnetic Spectrum



Optical vs Infrared



Optical vs Infrared

Hubble (color)



Credits : NASA, C.R. O'Dell and S.K. Wong (Rice University)

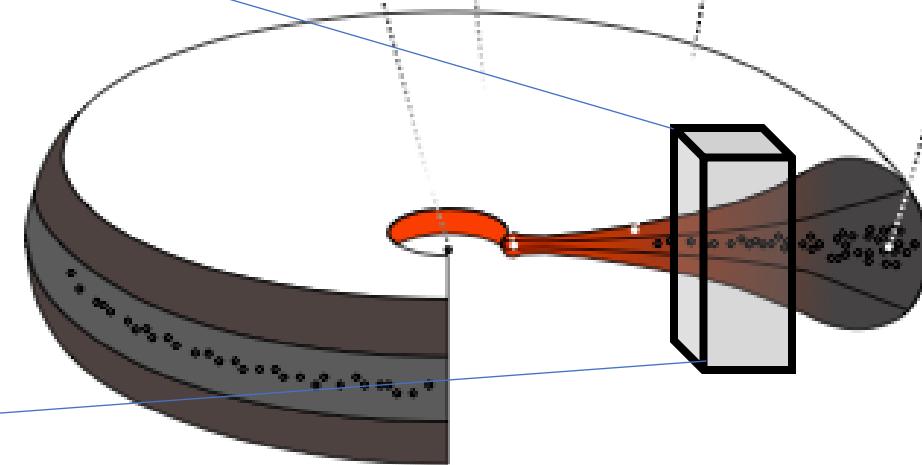
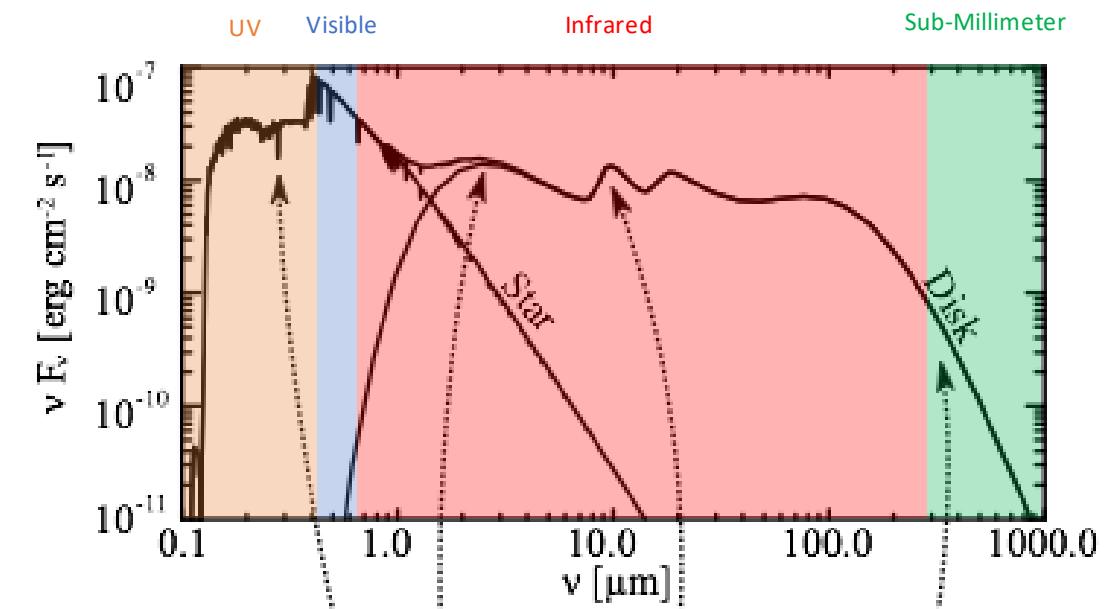
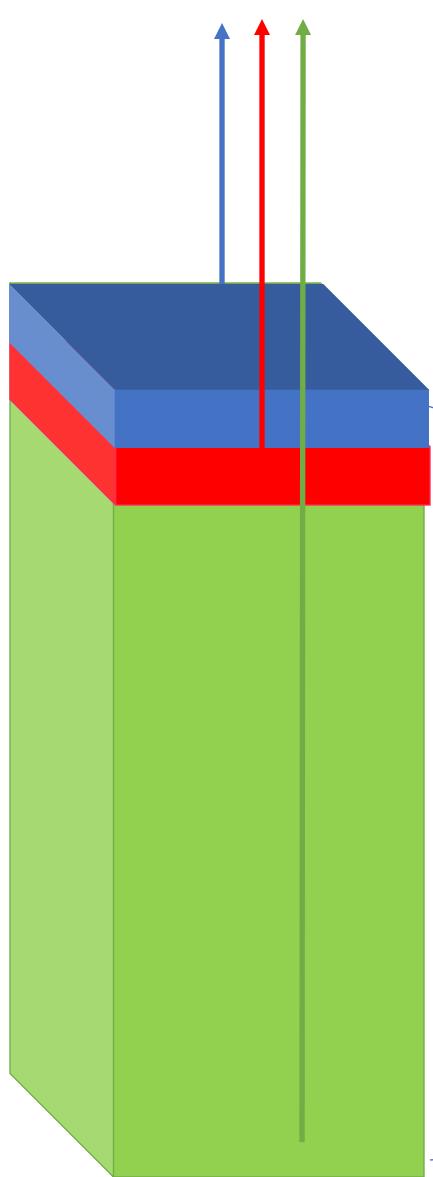
JWST (color)



Credits : NASA / ESA / CSA / PDRs4All team S. Fuenmayor



Disk spectra

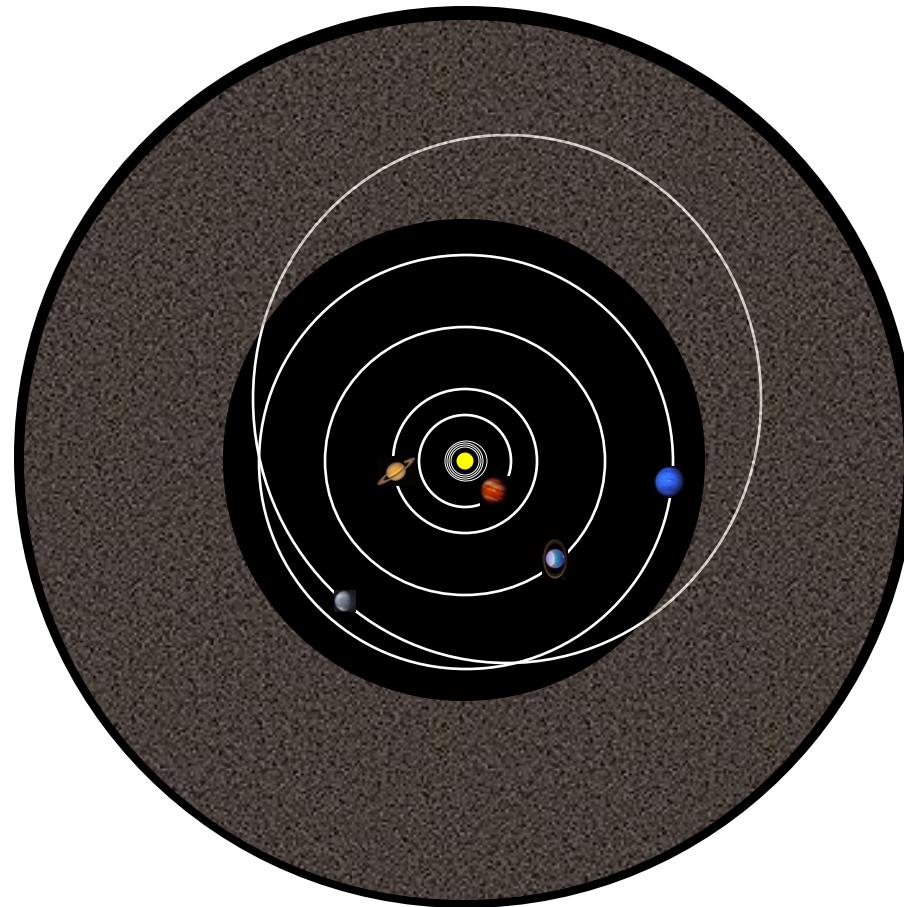


**The Atacama Large (sub-)Millimeter Array
(ALMA)**

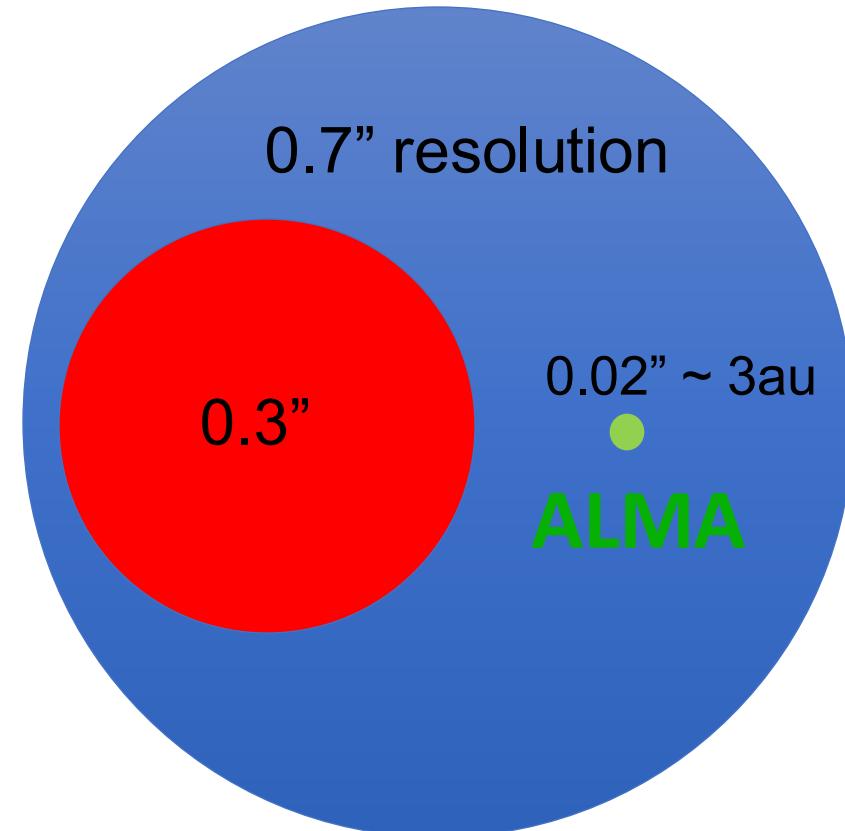


**The Very Large Array
(VLA)**

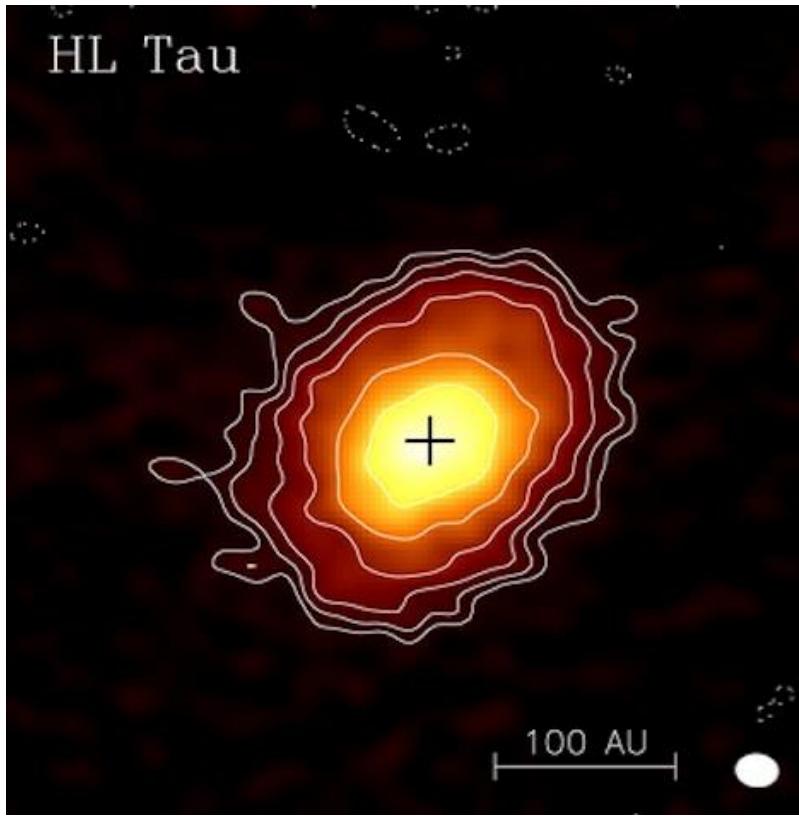




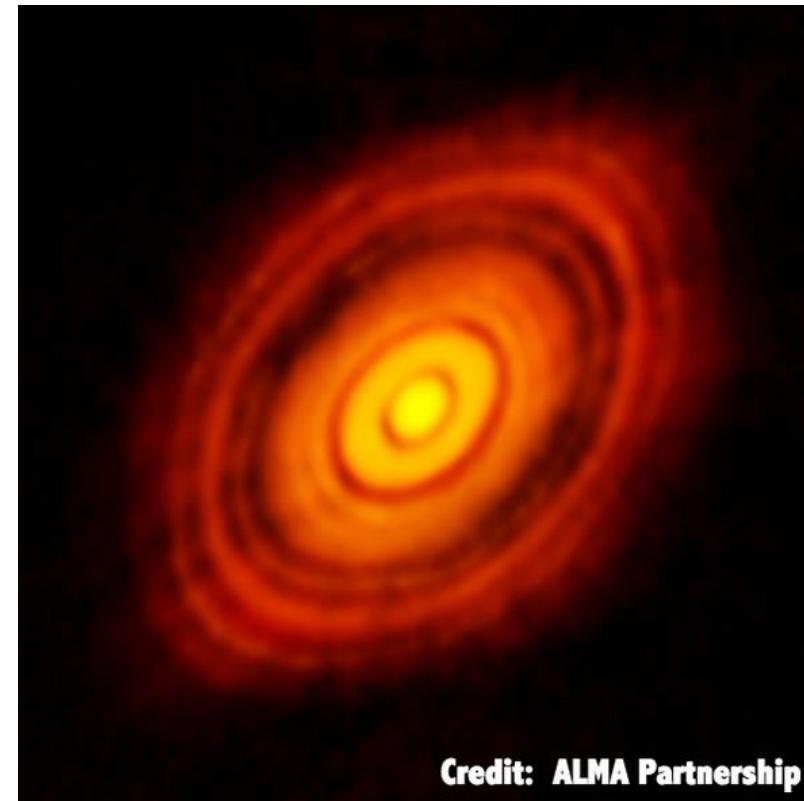
At 450 light-years



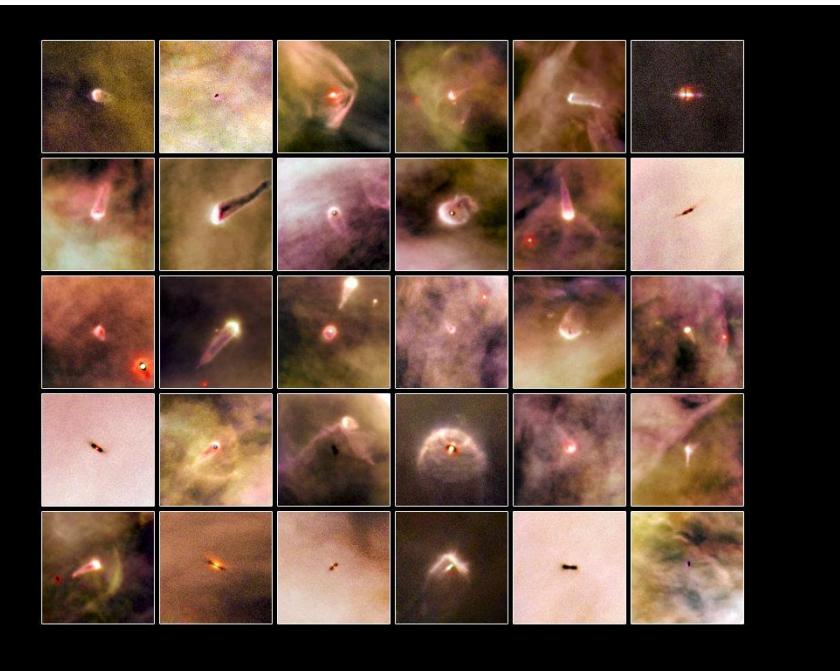
Before ALMA



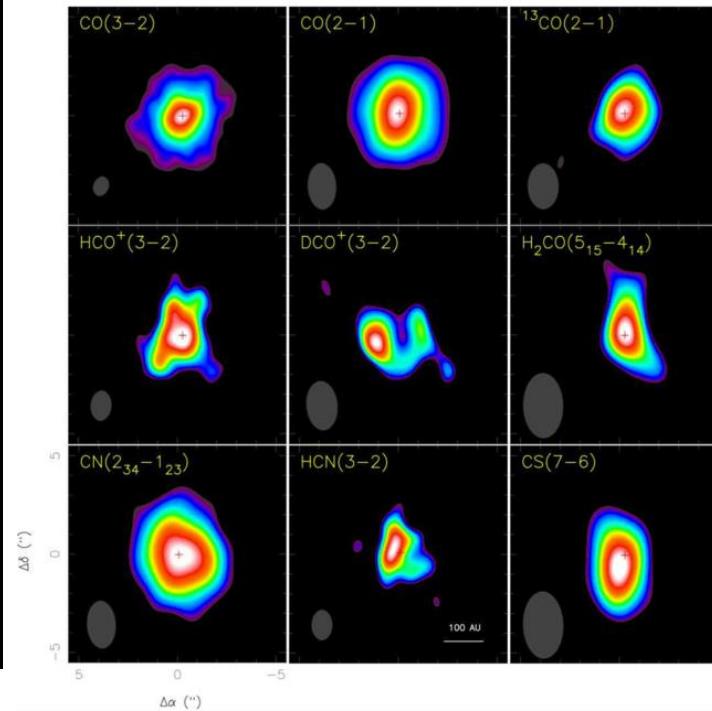
ALMA



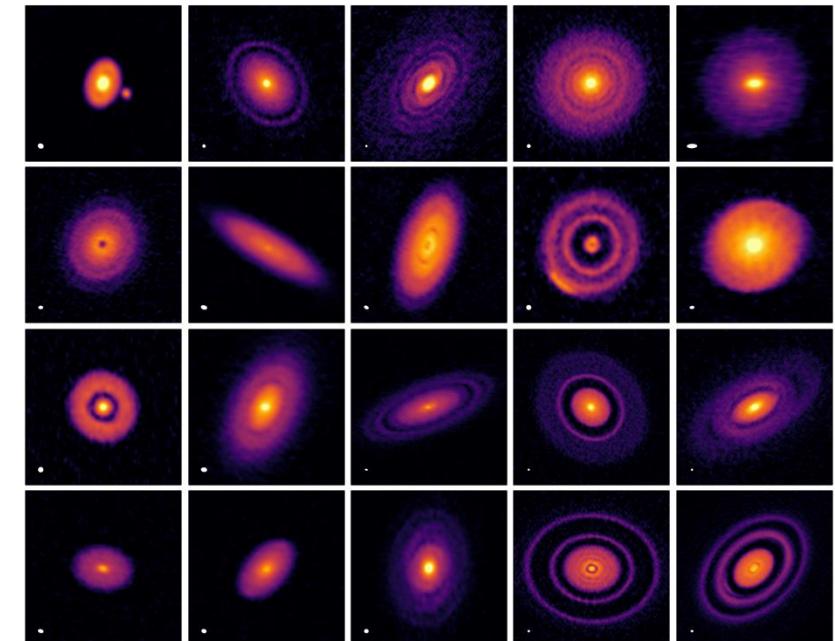
**Hubble Space Telescope
(HST)**

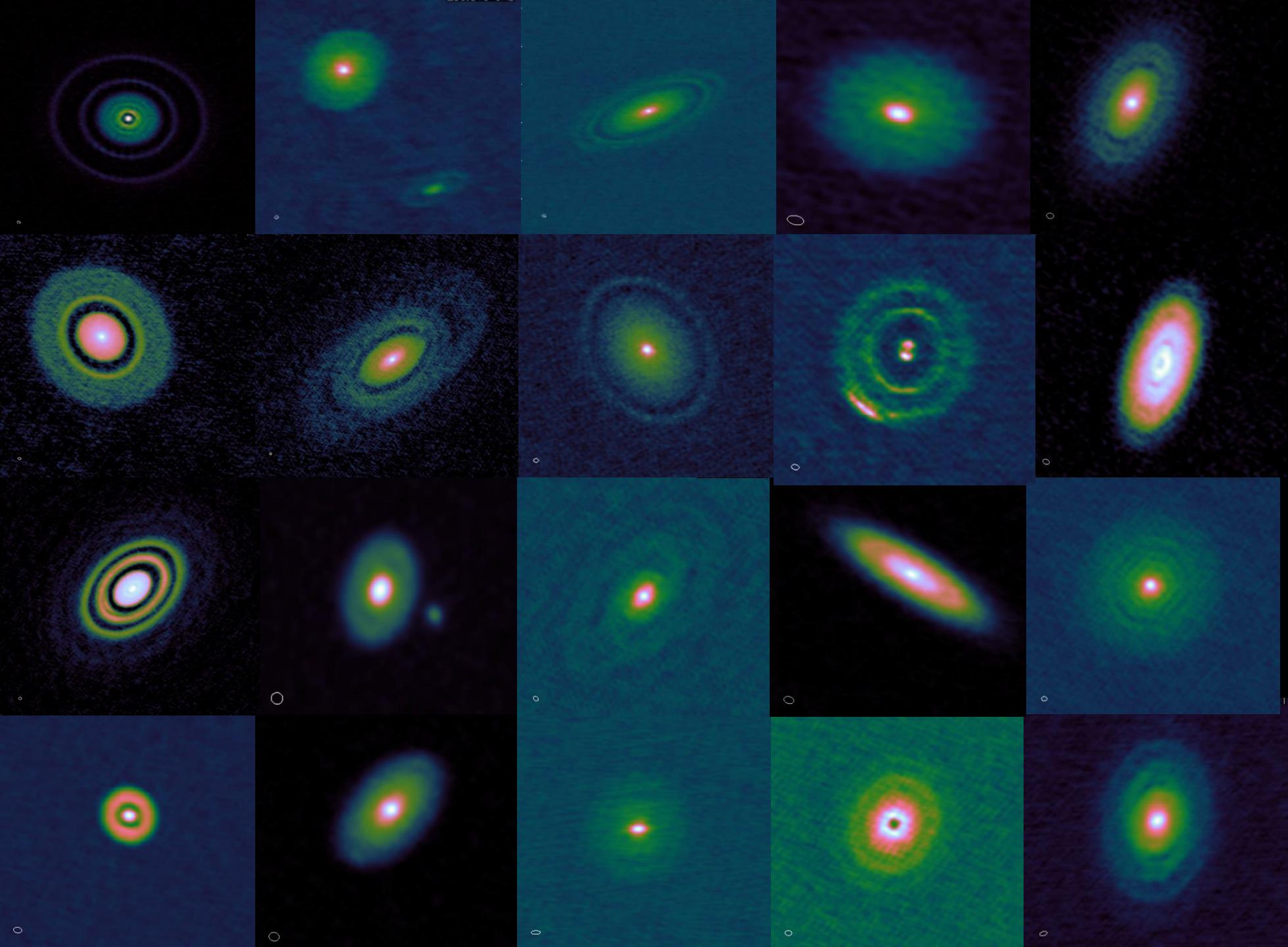


**Sub-Millimeter Array
(SMA)**



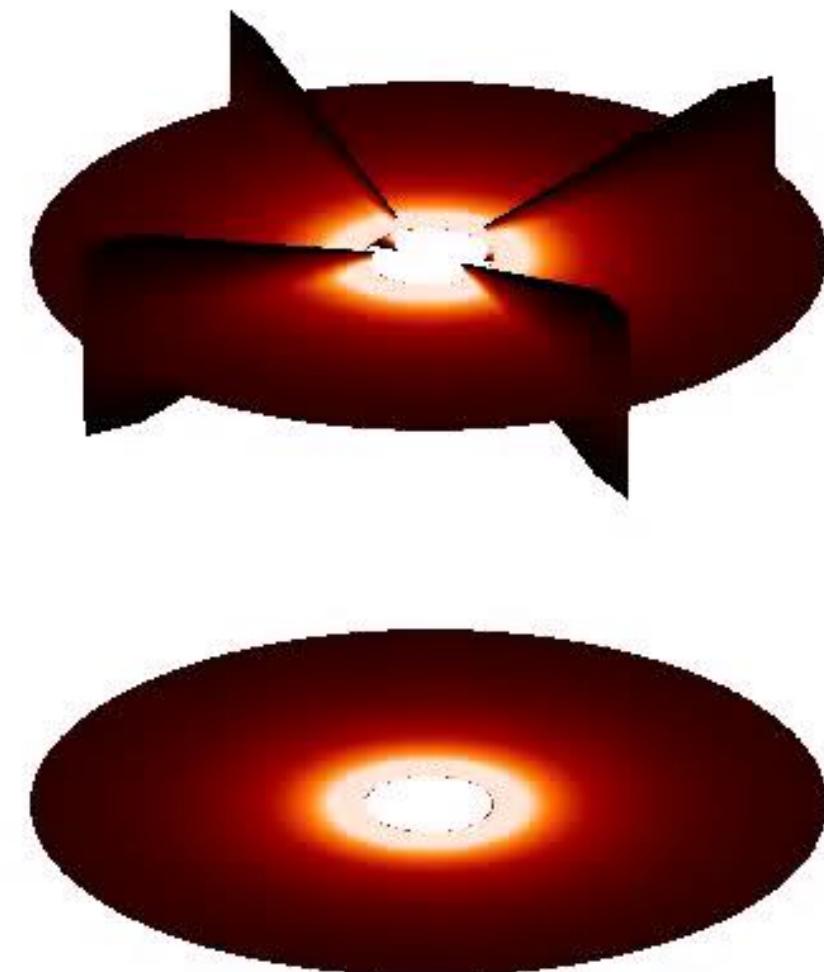
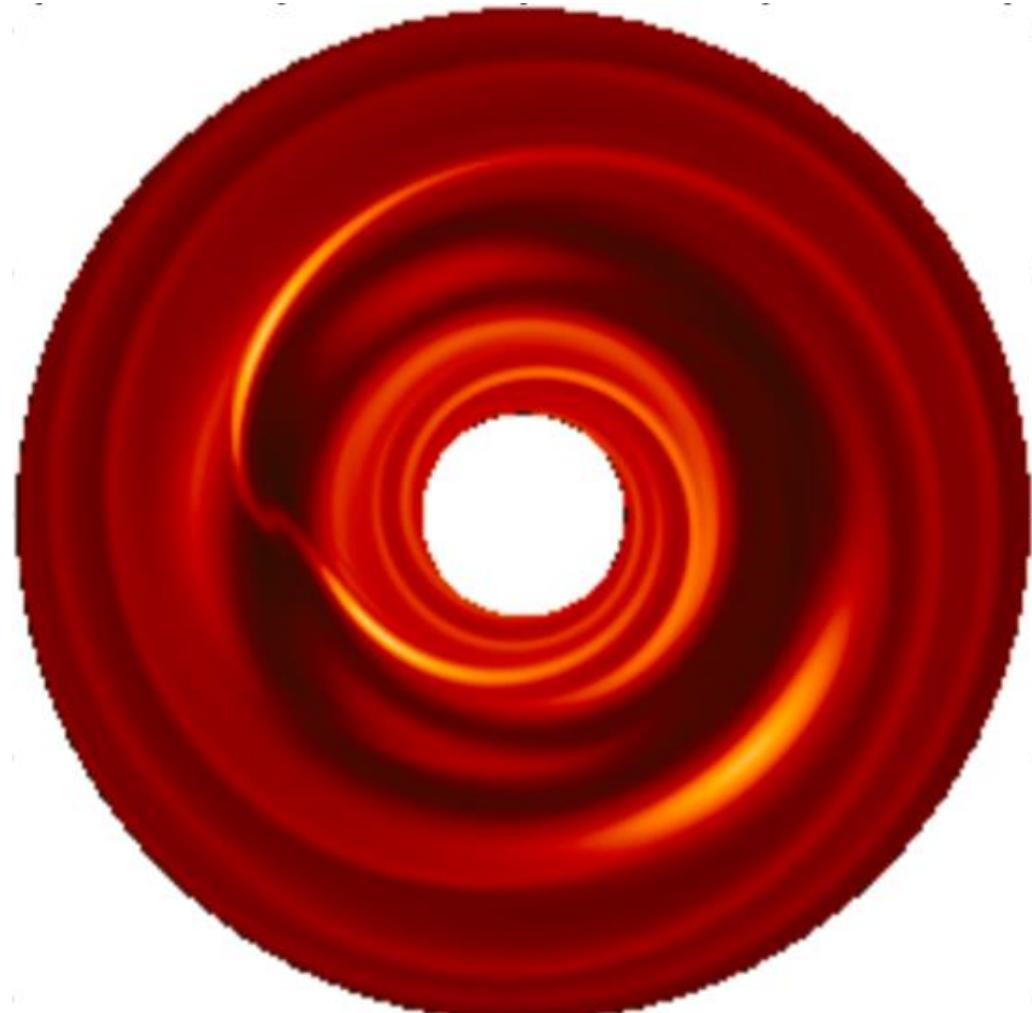
**Atacama Large
(sub-)Millimeter Array
(ALMA)**





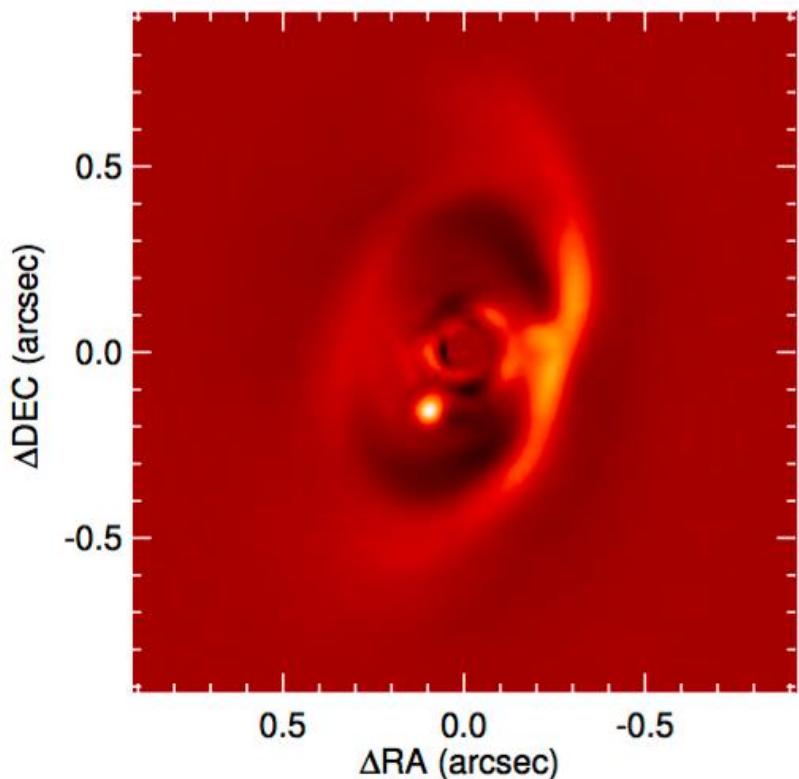
Planets carve gaps in disks

$t = 0.1$

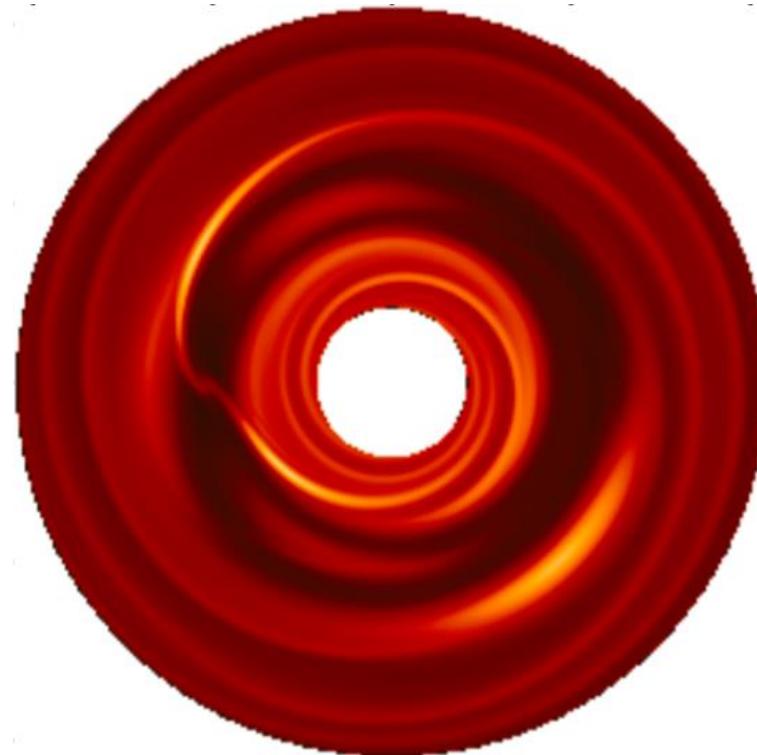


PDS 70b: A young planet spotted inside a gap

PDS 70 and PDS 70b

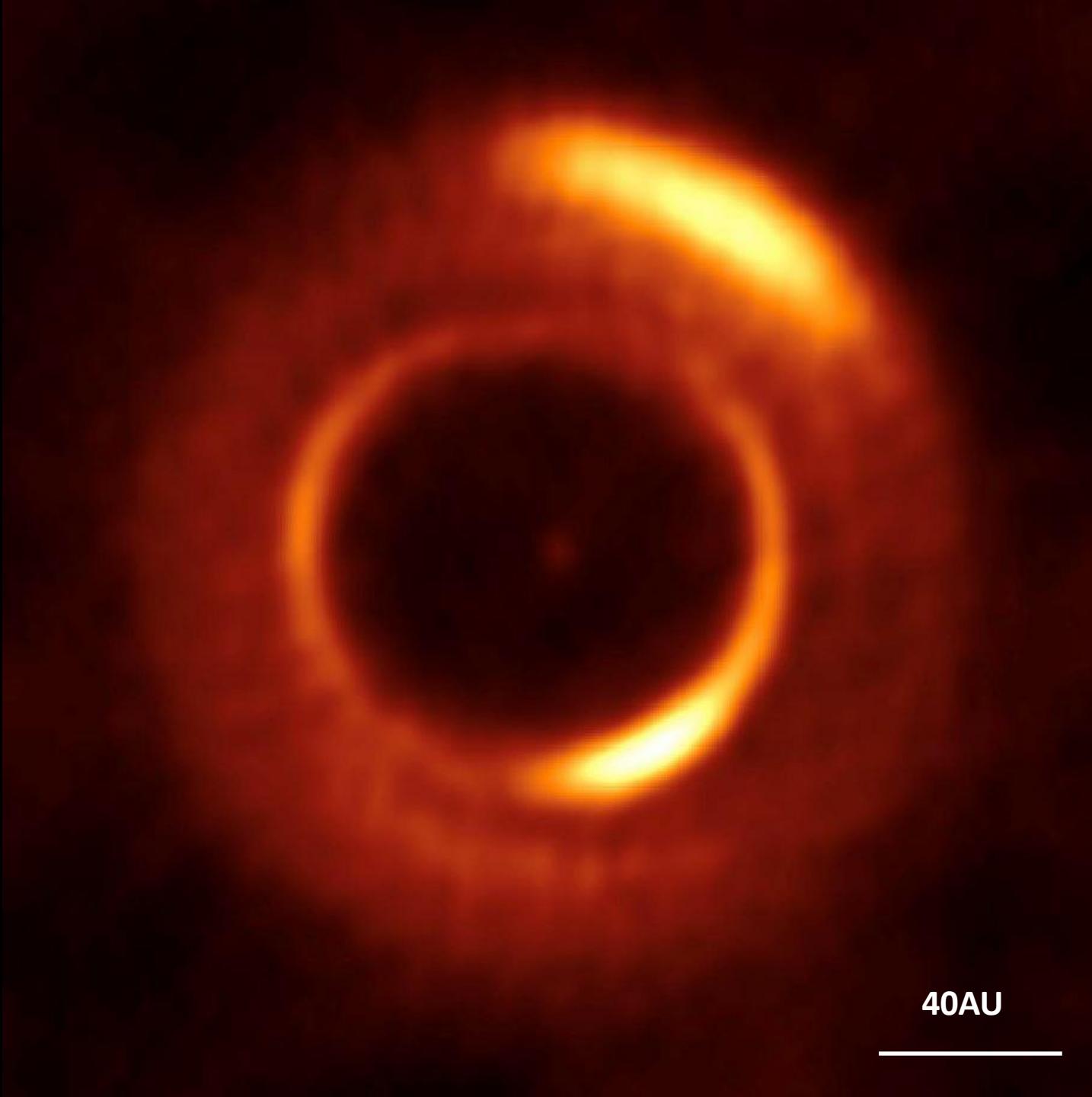


Keppler et al. (2018)
© ESO / A. Müller, MPIA

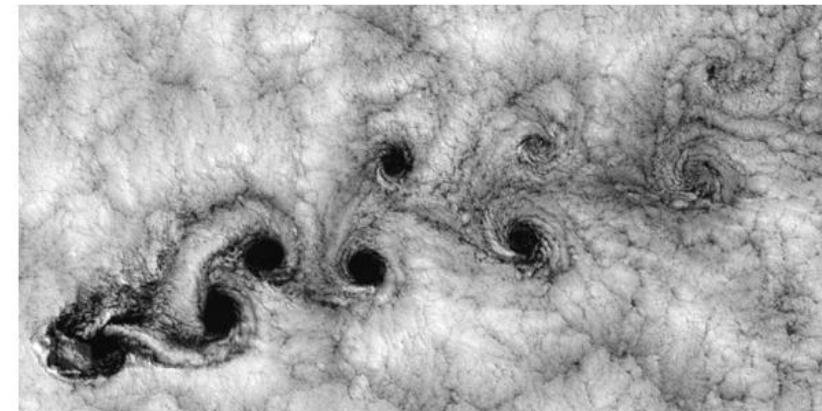


Lyra et al (2009)

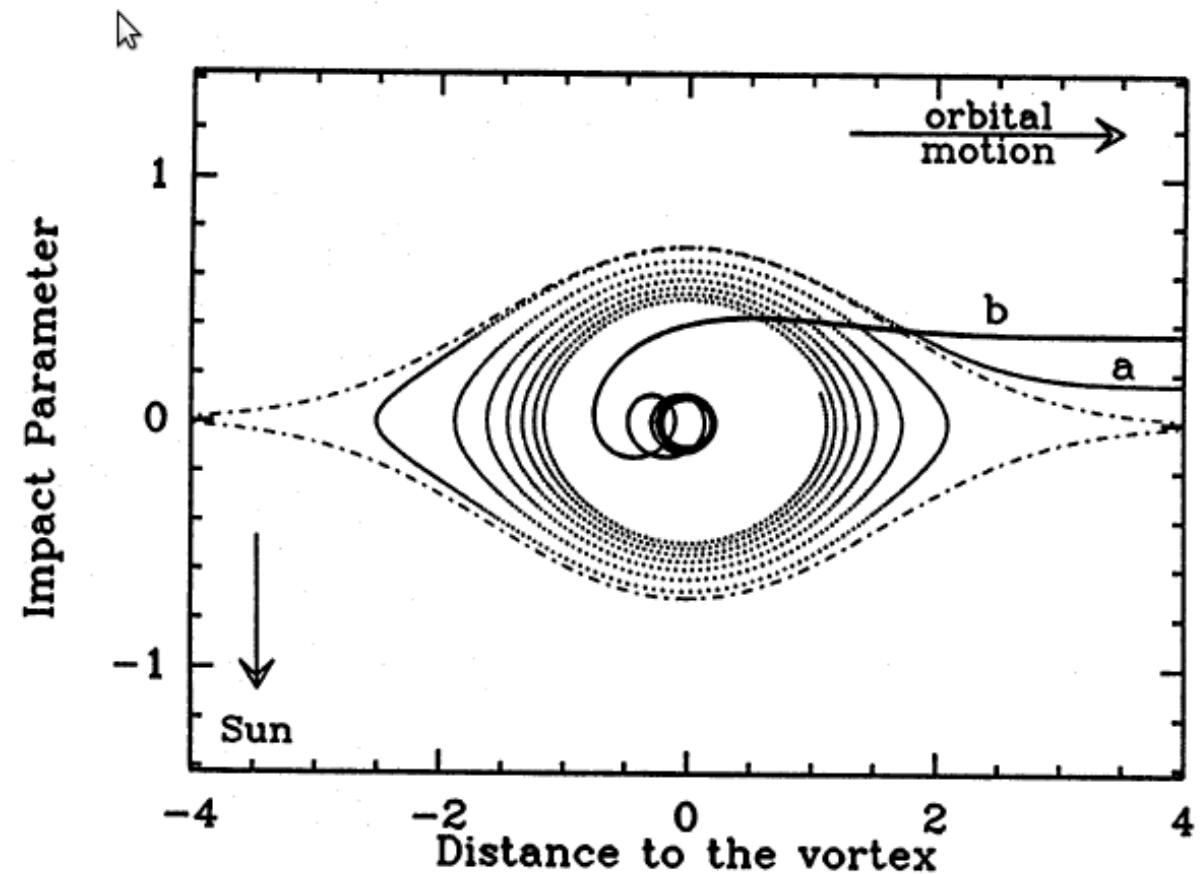
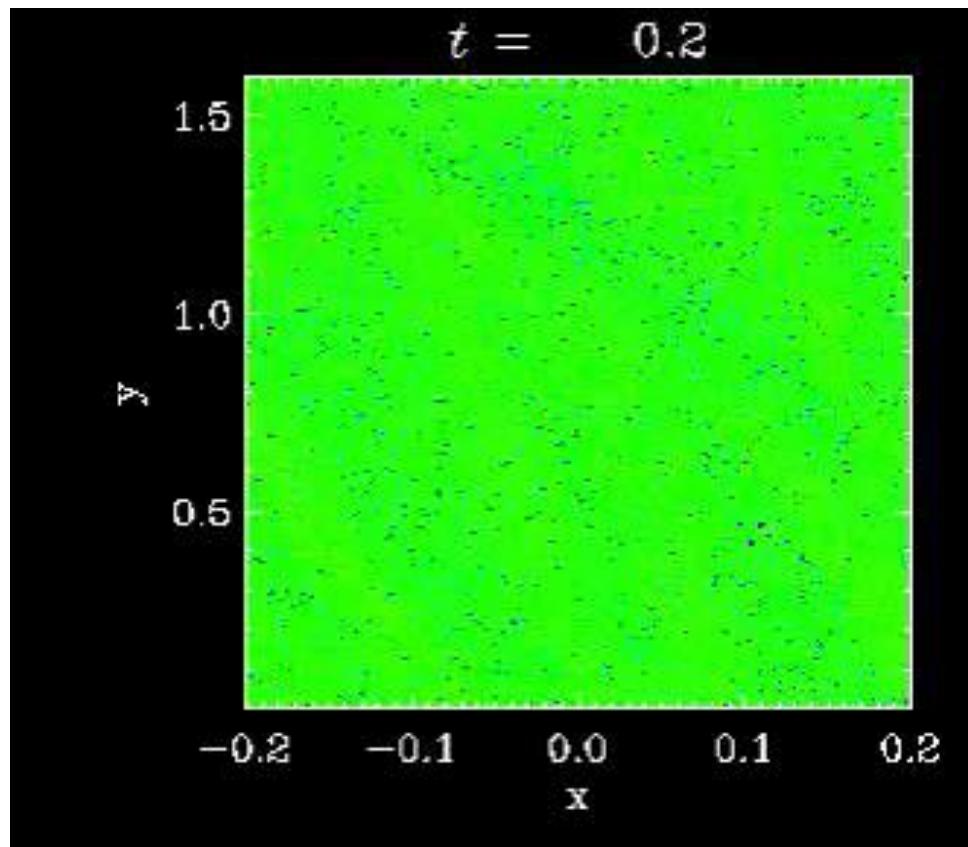
MWC 758



Vortices – an ubiquitous fluid mechanics phenomenon



Vortex Trapping



Video credit: Natalie Raettig

Barge & Sommeria (1995)

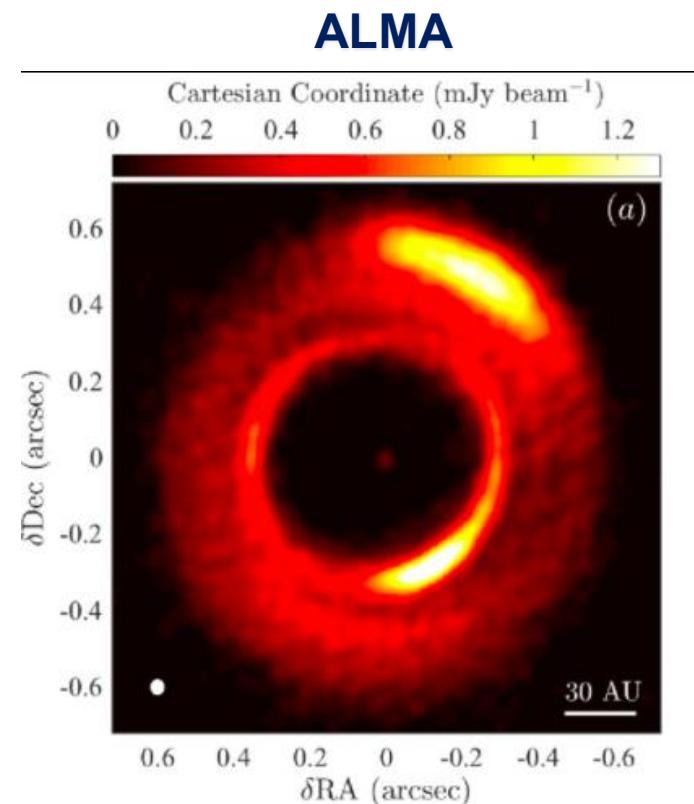
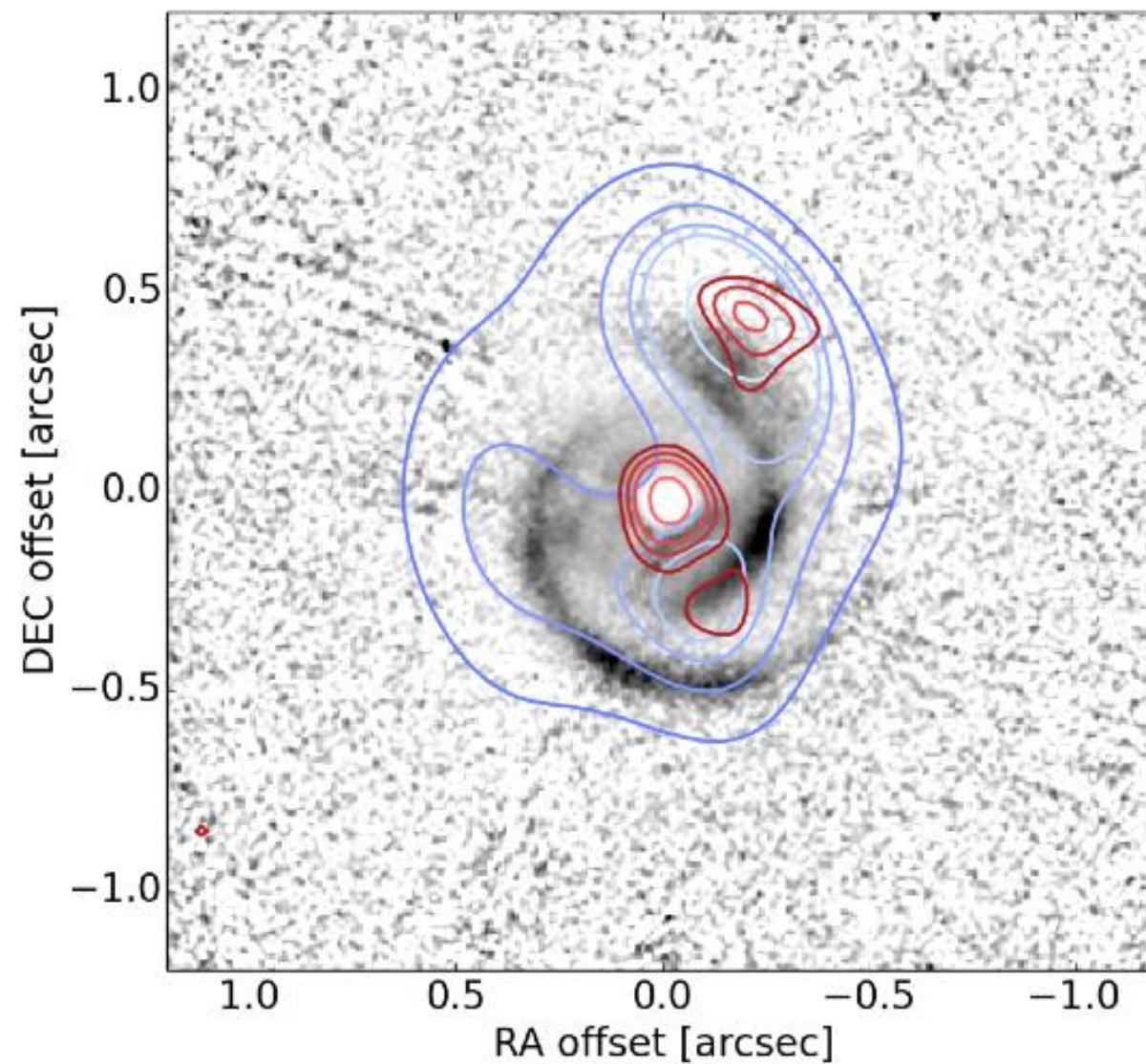
Disk Tomography

SPHERE-ALMA-VLA overlay of MWC 758

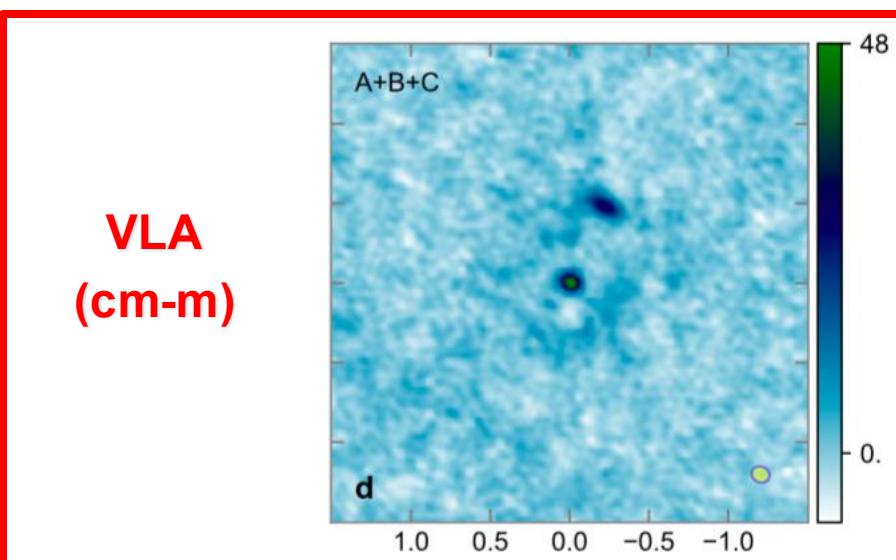
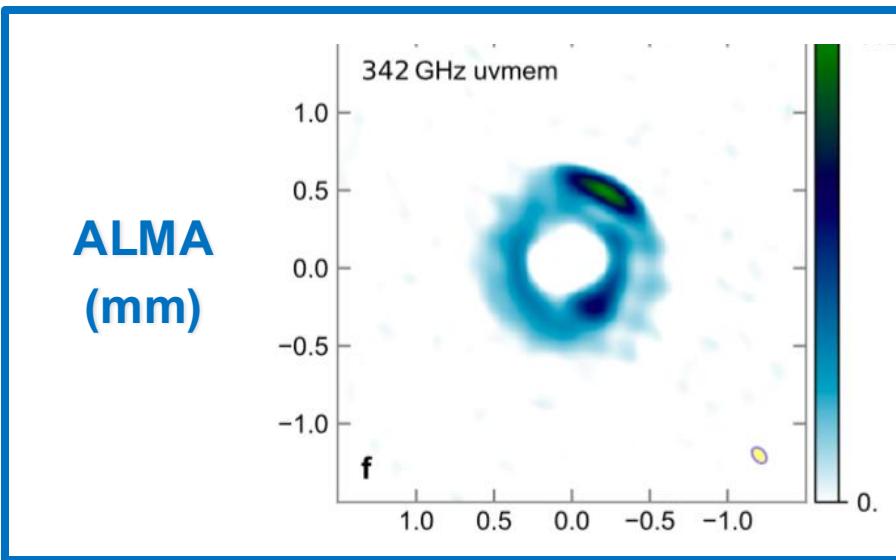
SPHERE (μm)

ALMA ($\sim \text{mm}$)

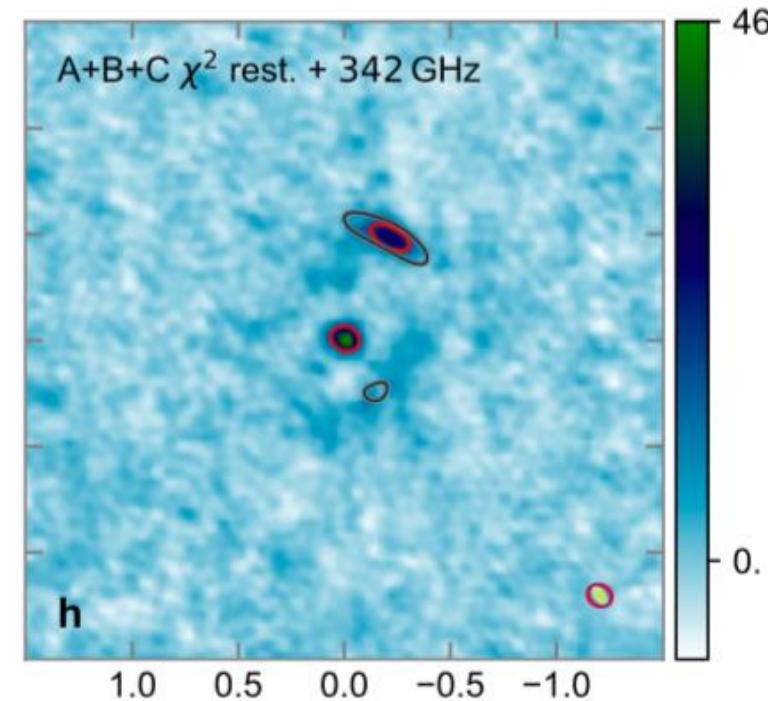
VLA (cm-m)



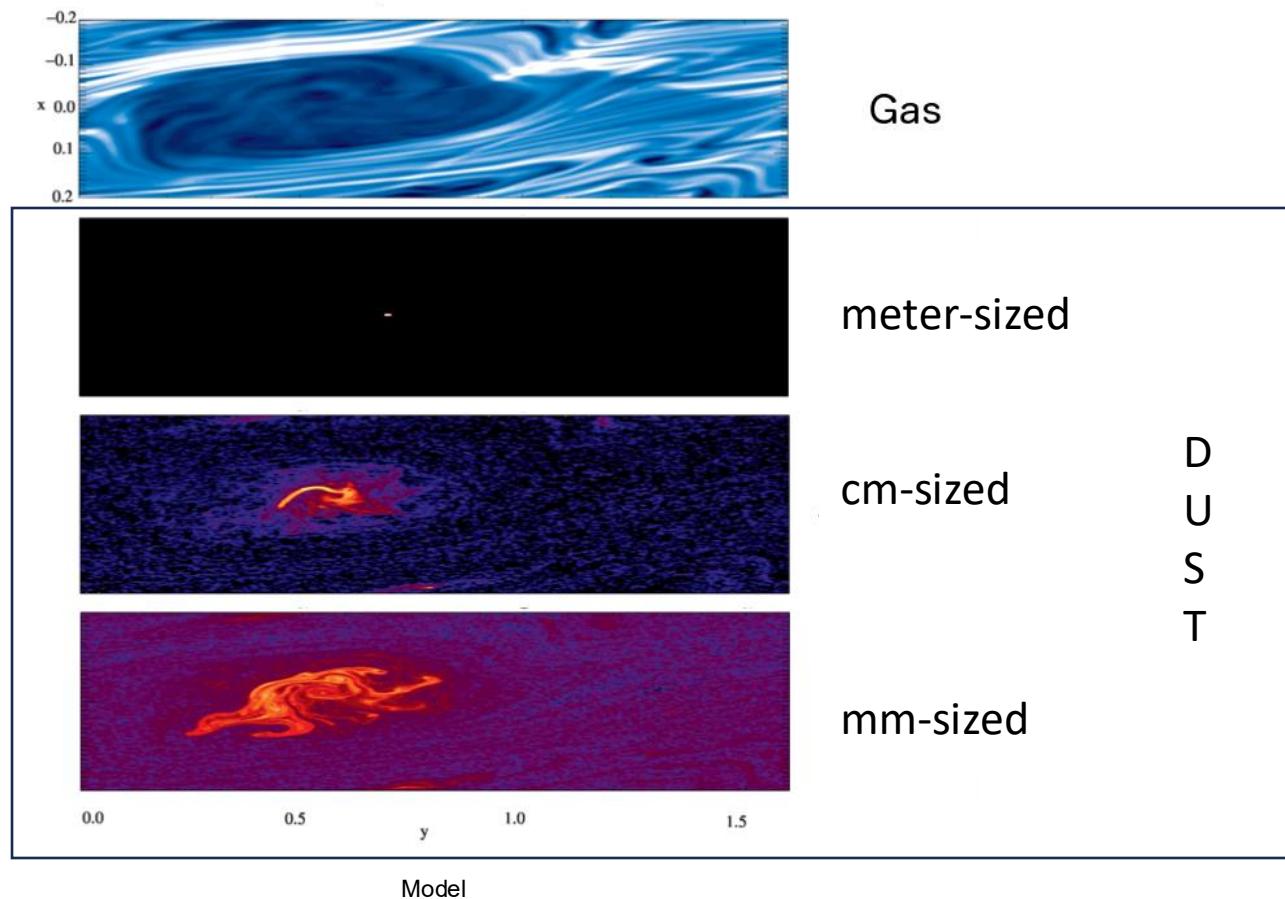
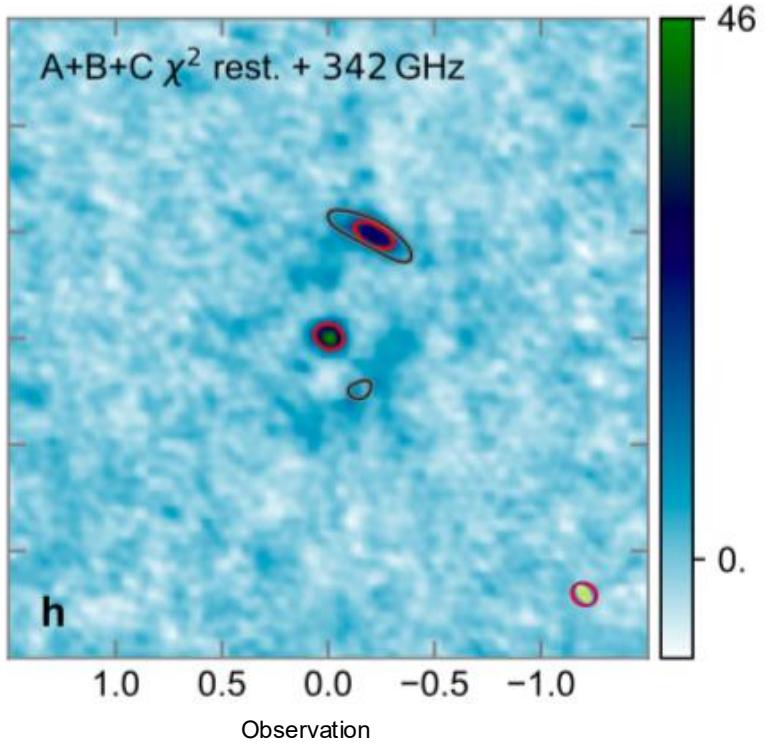
Vortex trapping



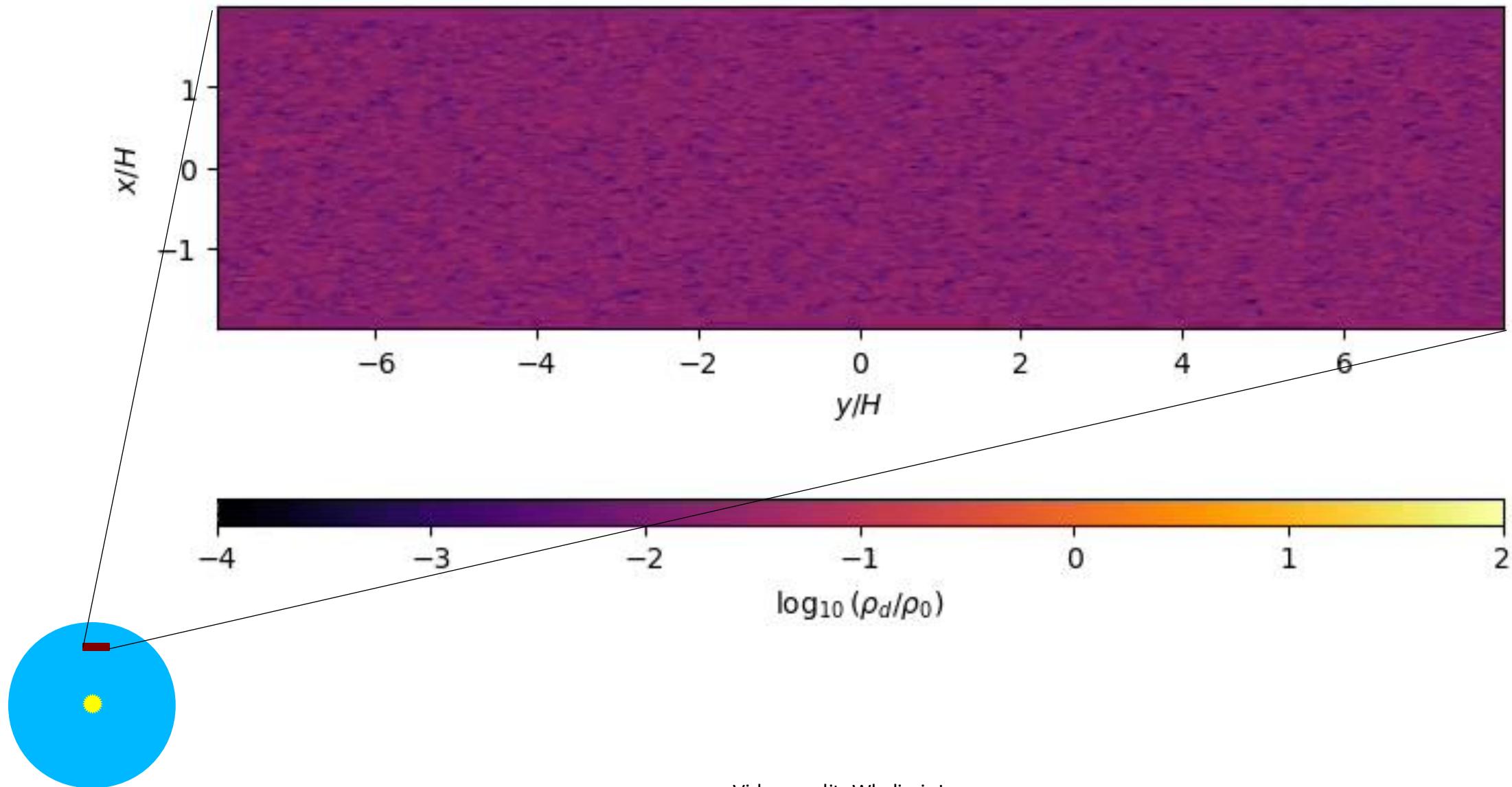
Overlay



Model vs Observation

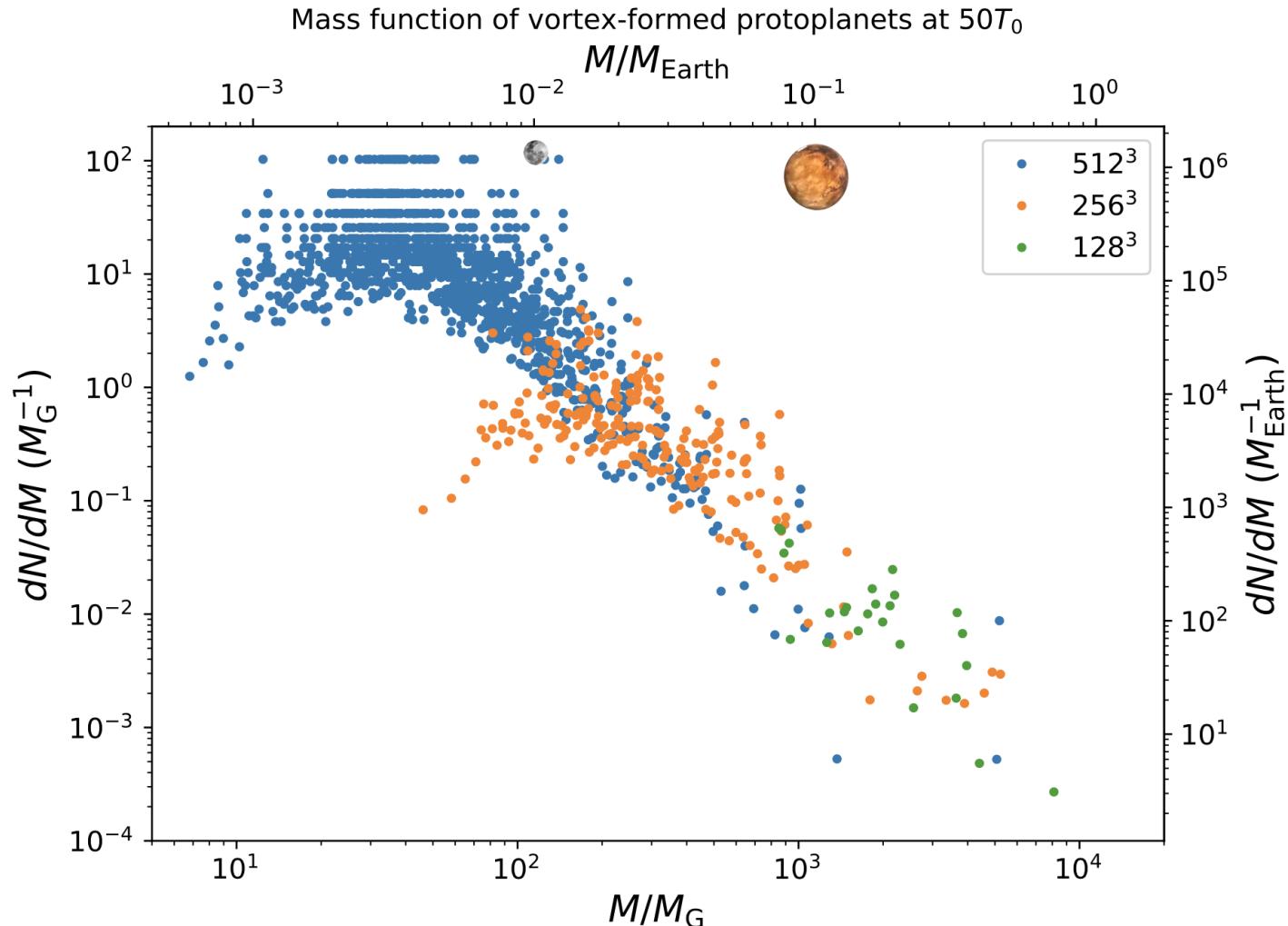


Vortex Trapping

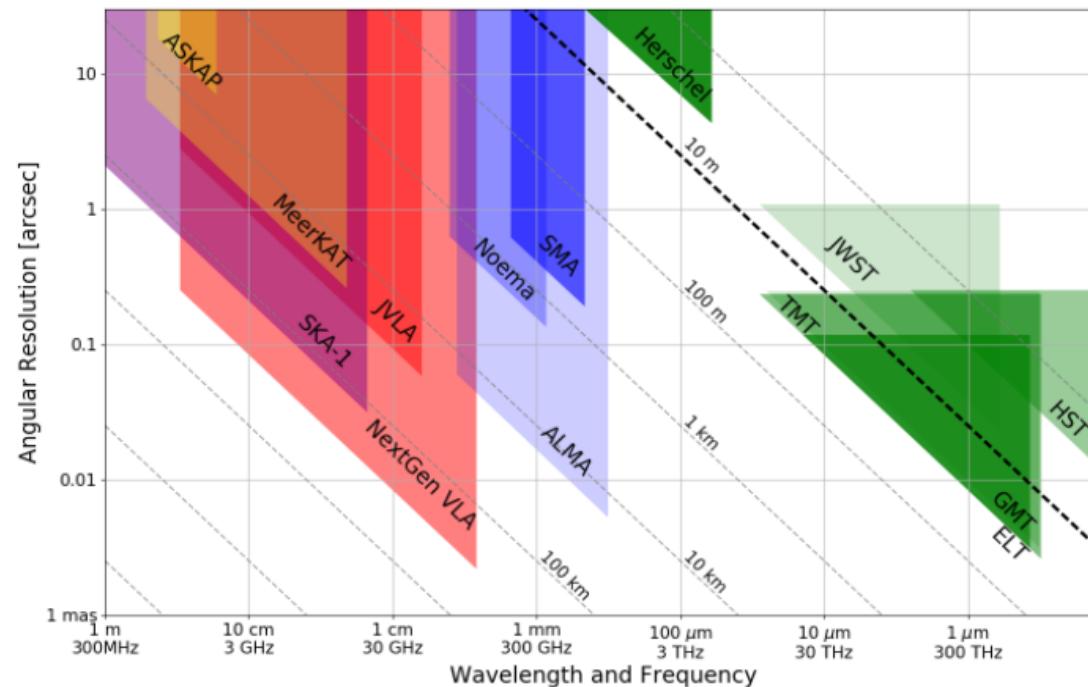


Video credit: Wladimir Lyra

Mass Function



Next Generation Very Large Array (ngVLA)



Planets at 5AU

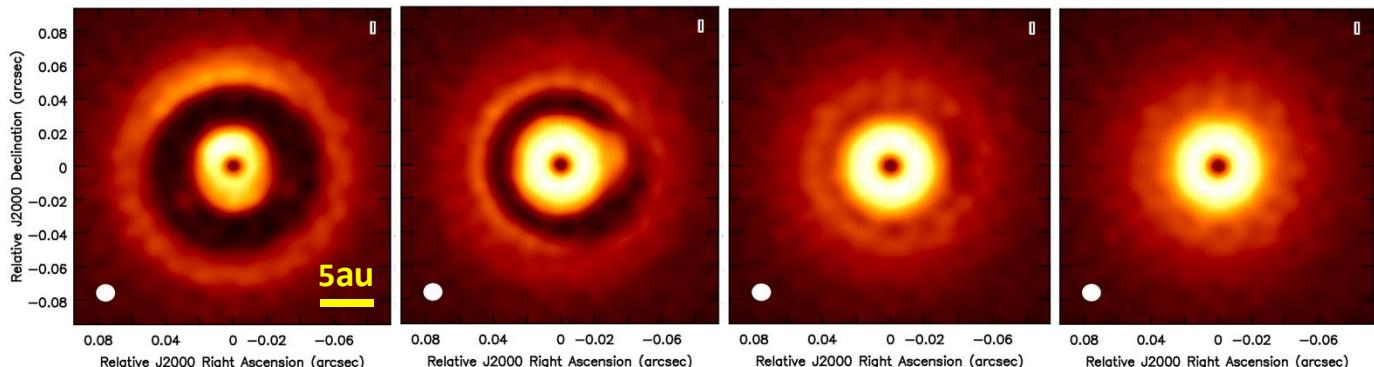
Jupiter

Saturn

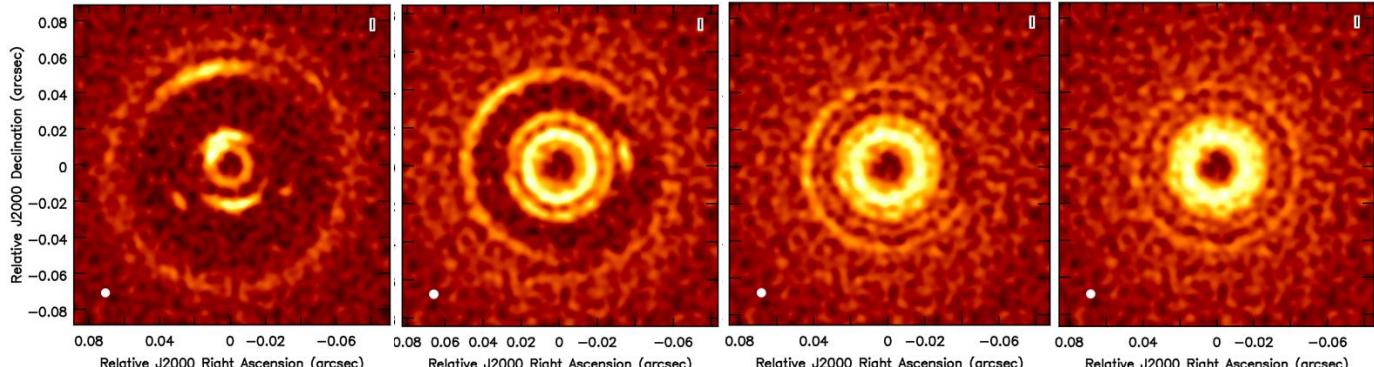
Neptune

$10 M_{\text{Earth}}$

ALMA at 0.87mm



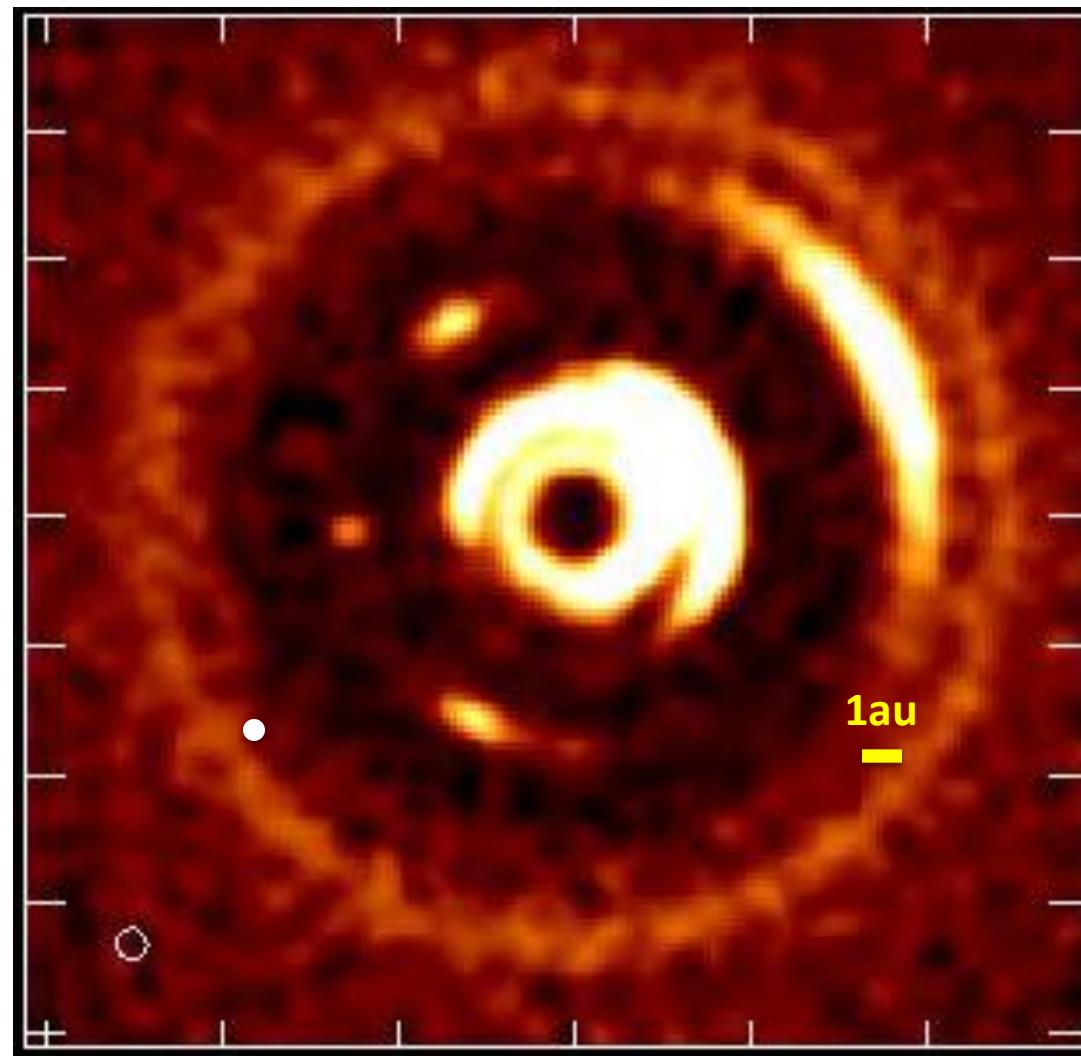
ngVLA at 3mm



ngVLA will identify gaps/substructures down to $\sim 5\text{-}10 M_{\text{Earth}}$

ngVLA

Jupiter at 5 AU – 500 light years away



Credit: Luca Ricci