# SECRETS OF THE SOLAR SYSTEM 5'-01

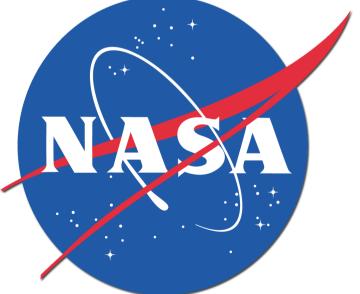
Wladimir (Wlad) Lyra Brian Levine

AMNH After-School Program

American Museumö Natural History



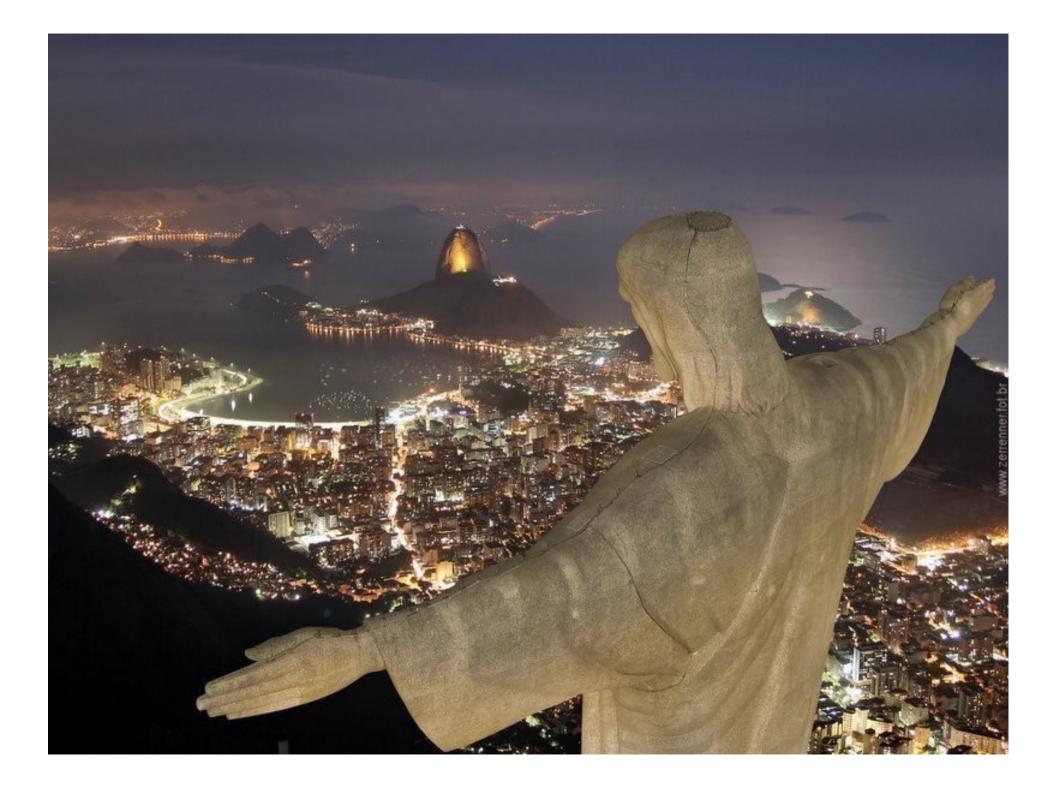
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### Quick Bio

Wladimir (Wlad) Lyra





### Quick Bio

### Wladimir (Wlad) Lyra

B.Sc. in Astronomy, Federal University of Rio de Janeiro (Brazil), 1999-2003.

Research Assistant 2003-2004 Space Telescope (Baltimore MD - USA). Cerro Tololo Inter-American Observatory CTIO (La Serena - Chile). European Southern Observatory ESO (Munich - Germany). Lisbon Observatory (Lisbon - Portugal).

Ph.D. in Astronomy, Uppsala University (Uppsala - Sweden), 2004-2009. Nordic Institute for Theoretical Physics (Stockholm - Sweden).

#### Postdoctoral Researcher

Max-Planck Institute for Astronomy (Heidelberg - Germany), 2009. American Museum of Natural History (New York NY - USA), 2009-2011.

#### Stellar Astrophysics, Planetary Sciences

Solar-type stars, extrasolar planets, star and *planet formation*, hydrodynamics, plasma physics, turbulence.

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### Quick Bio

Brian Levine Astrophysics Educator

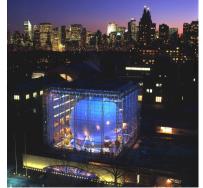


Brooklyn

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### Outline

Origins

Tour of the Solar System

## WHERE DO WE COME FROM?

### **Star Formation**

The space between stars is **NOT EMPTY**, it is just very low density

Some of it is gas (99%), some of it is dust (1%). This matter is called **INTERSTELLAR MEDIUM (ISM)** 

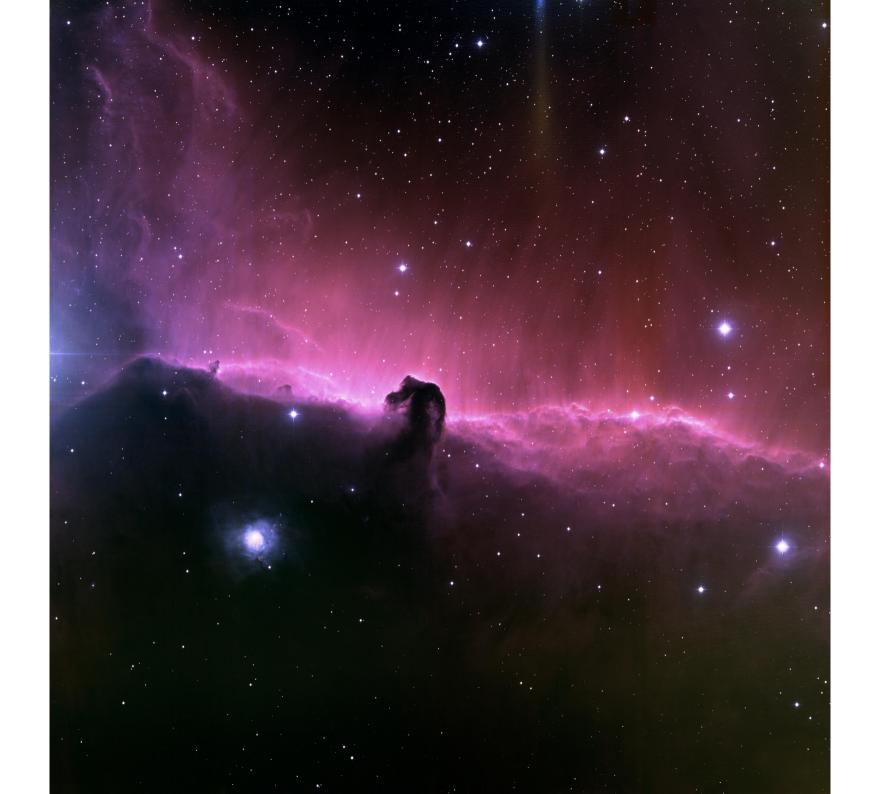


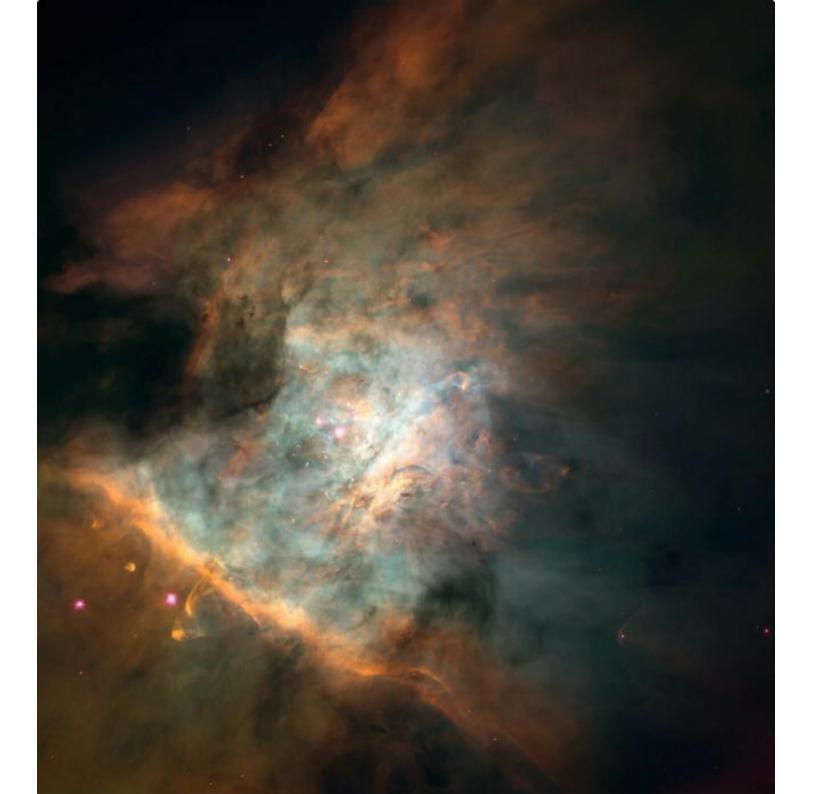
### **Star Formation**

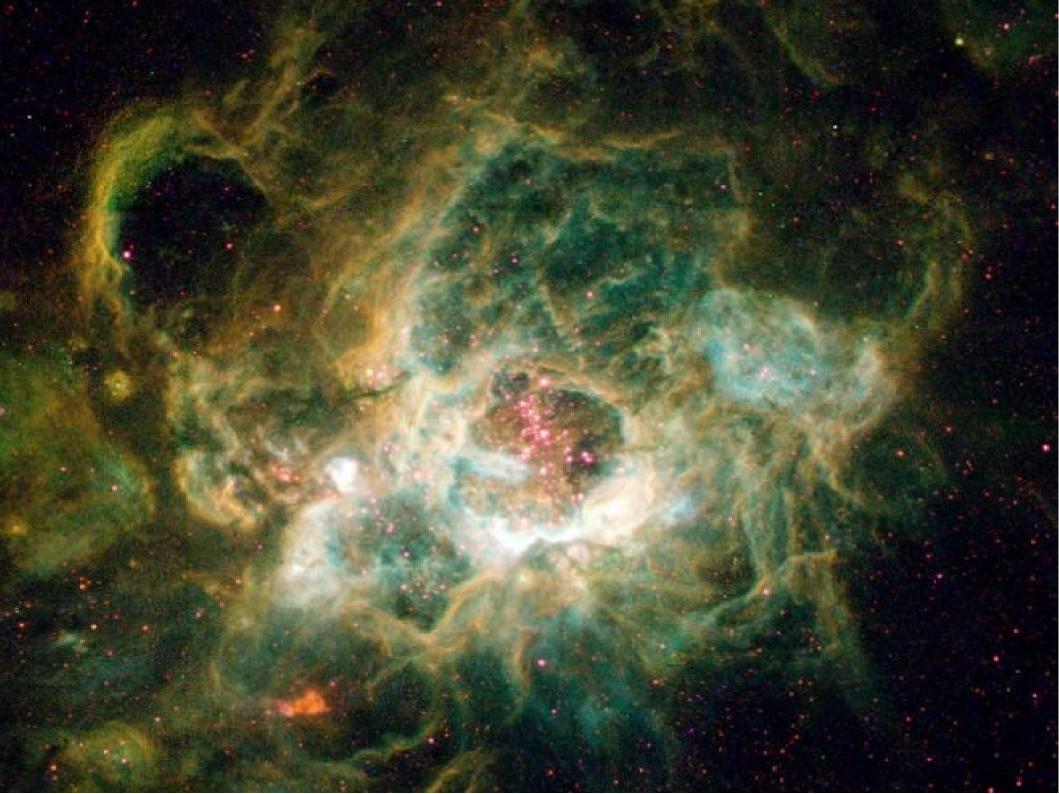
A process by which gas collapses gravitationally, to form stars.

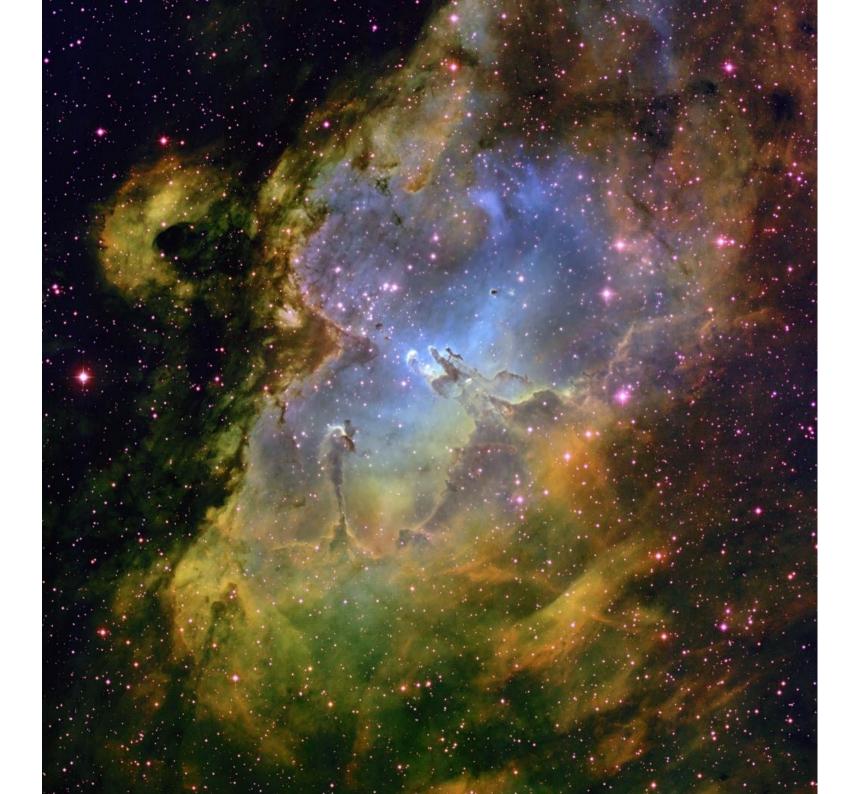
#### Star formation occurs in the *densest regions of the ISM*, called *Molecular Clouds*



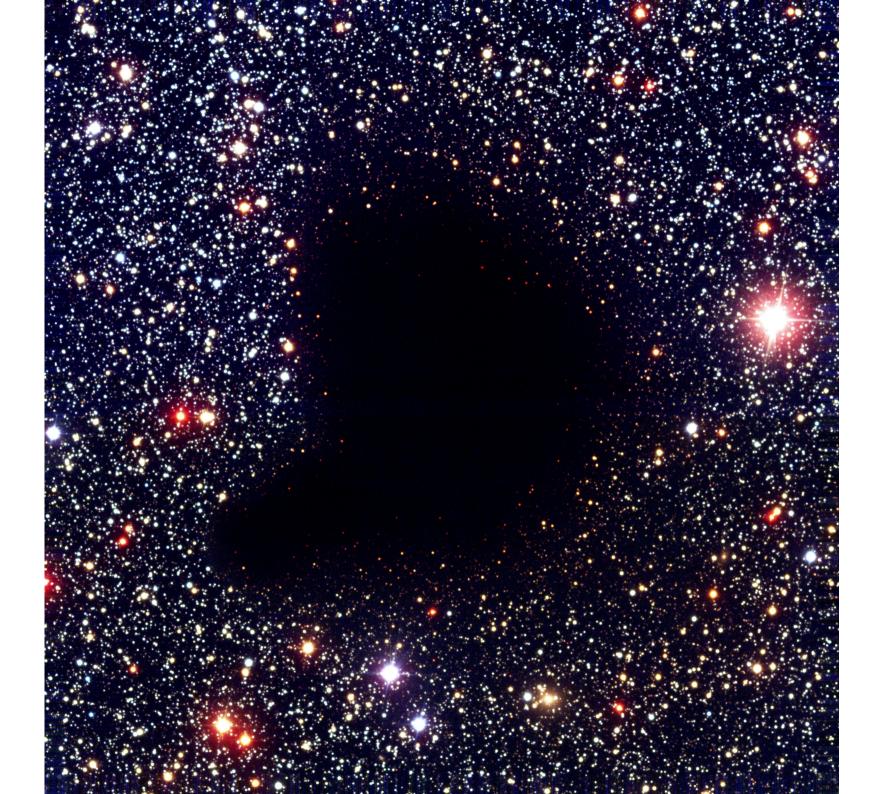


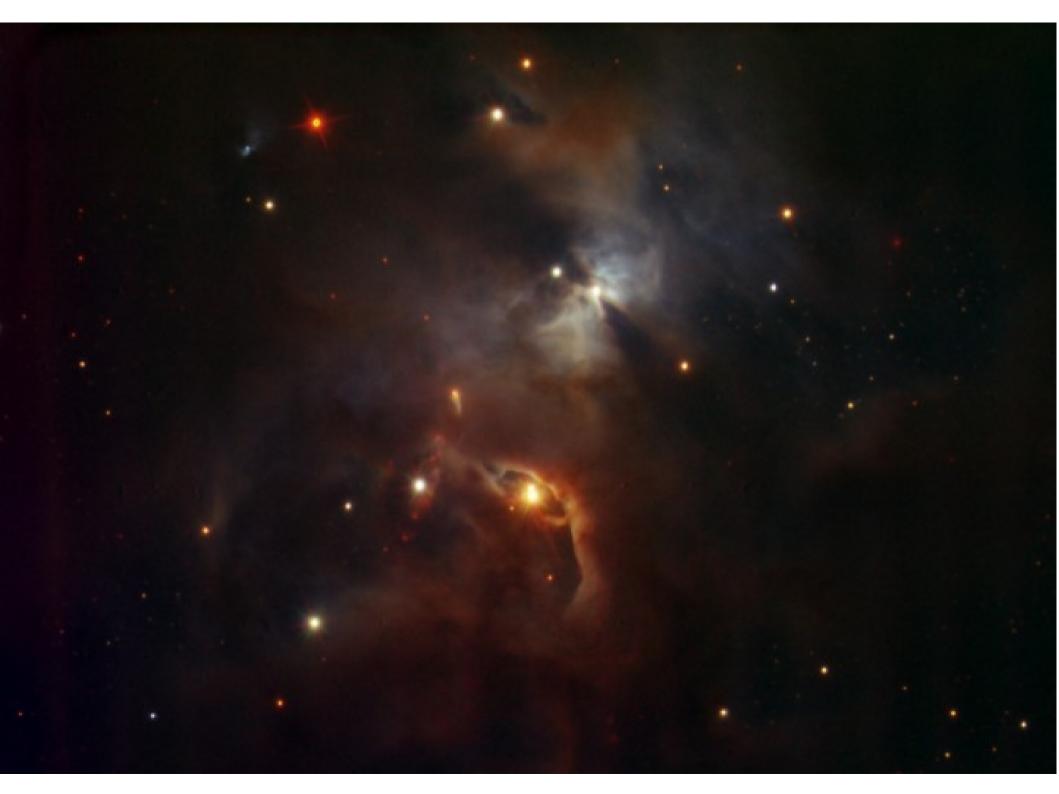




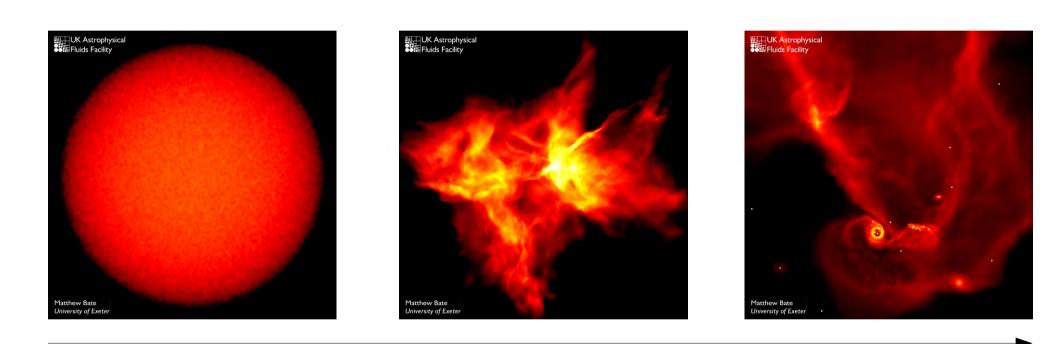








### A Simulation of Star Formation

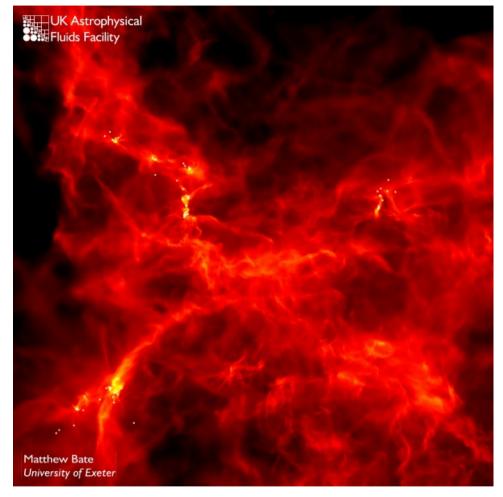


#### time

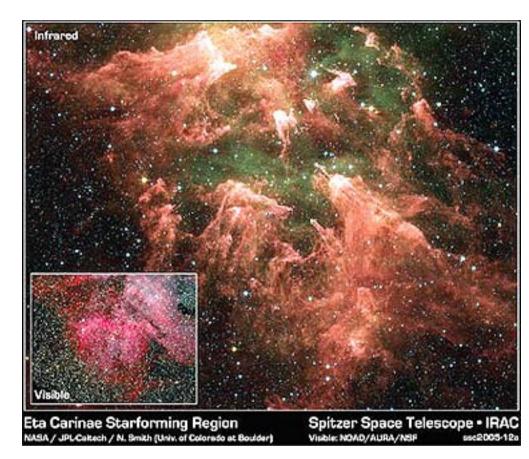
A Molecular cloud fragments into a clumpy structure of high and low density regions

The densest clumps are massive enough to undergo gravitational collapse and form stars

### A Simulation of Star Formation

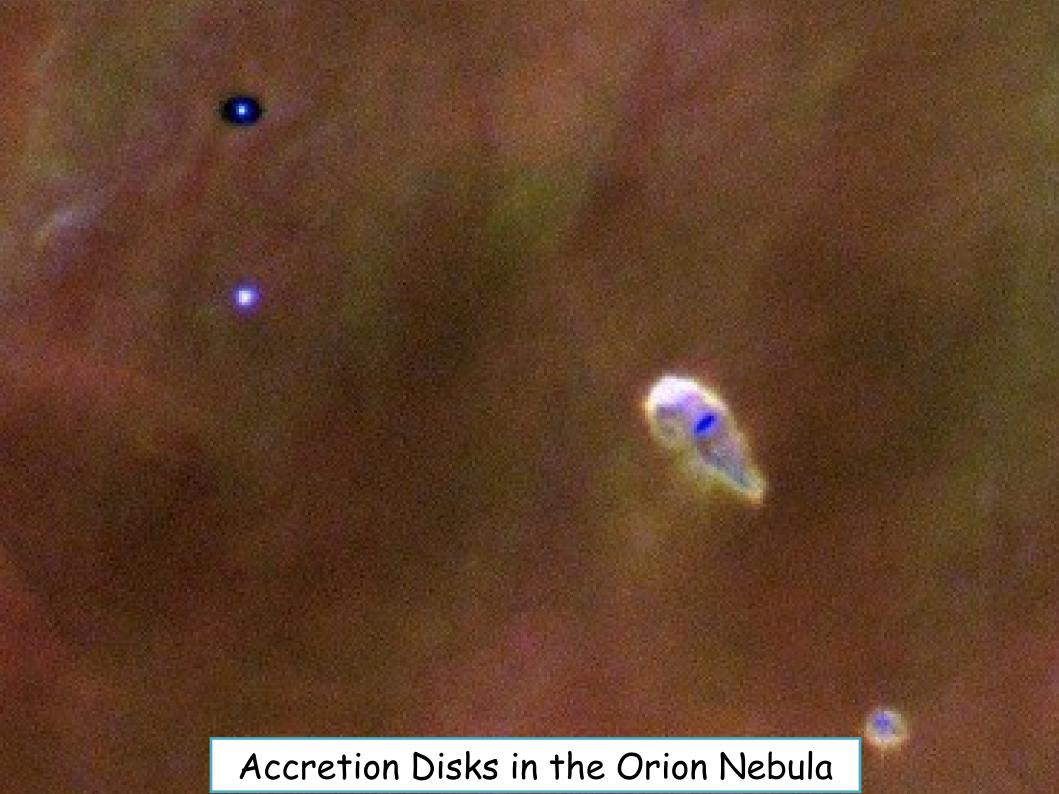


Computer simulation

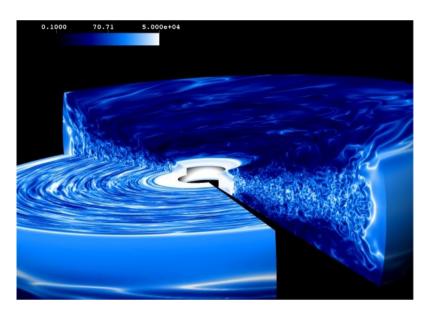


Observation

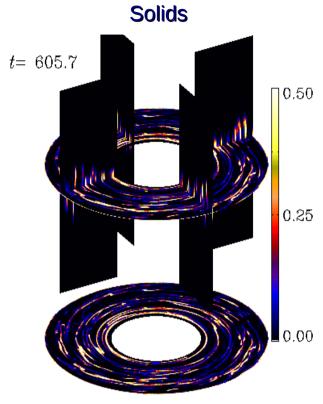




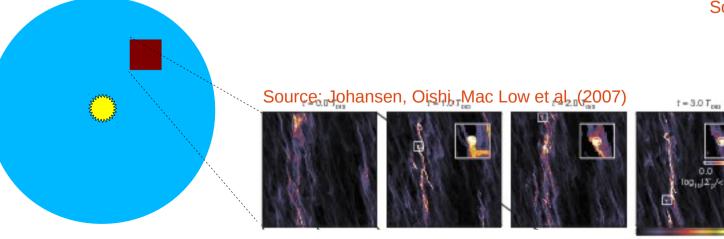
### Simulations of Planet Formation



Source: Flock et al. (2011)

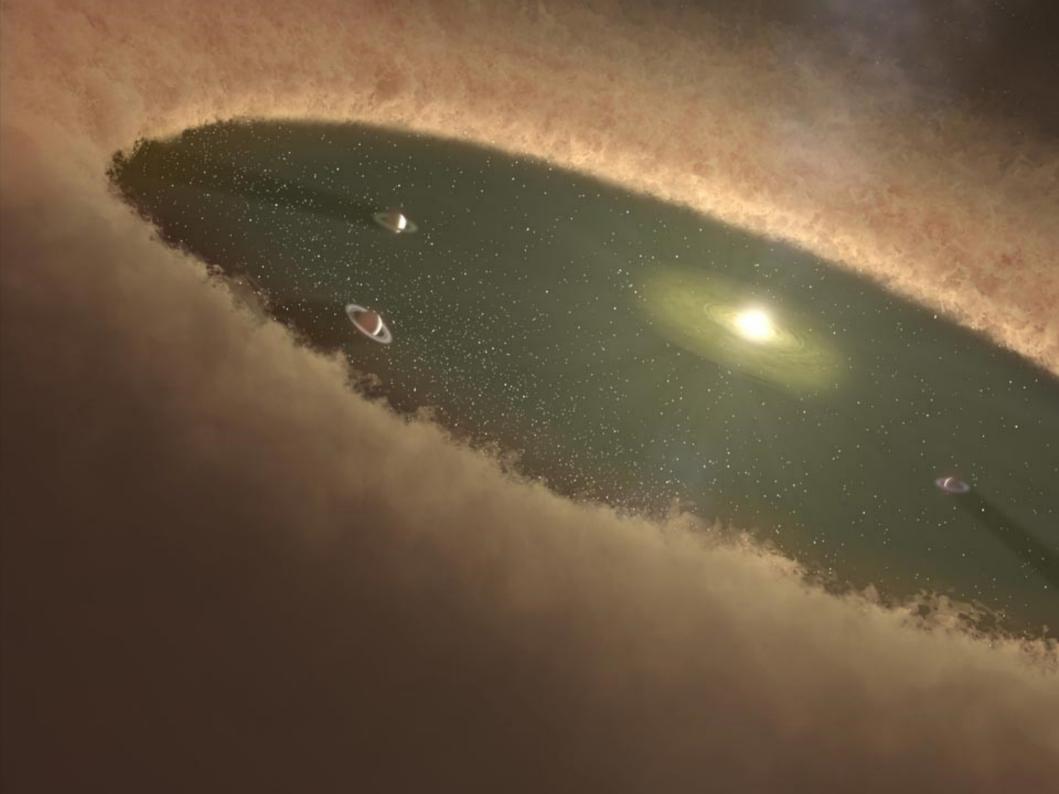


Source: Lyra et al. (2009)

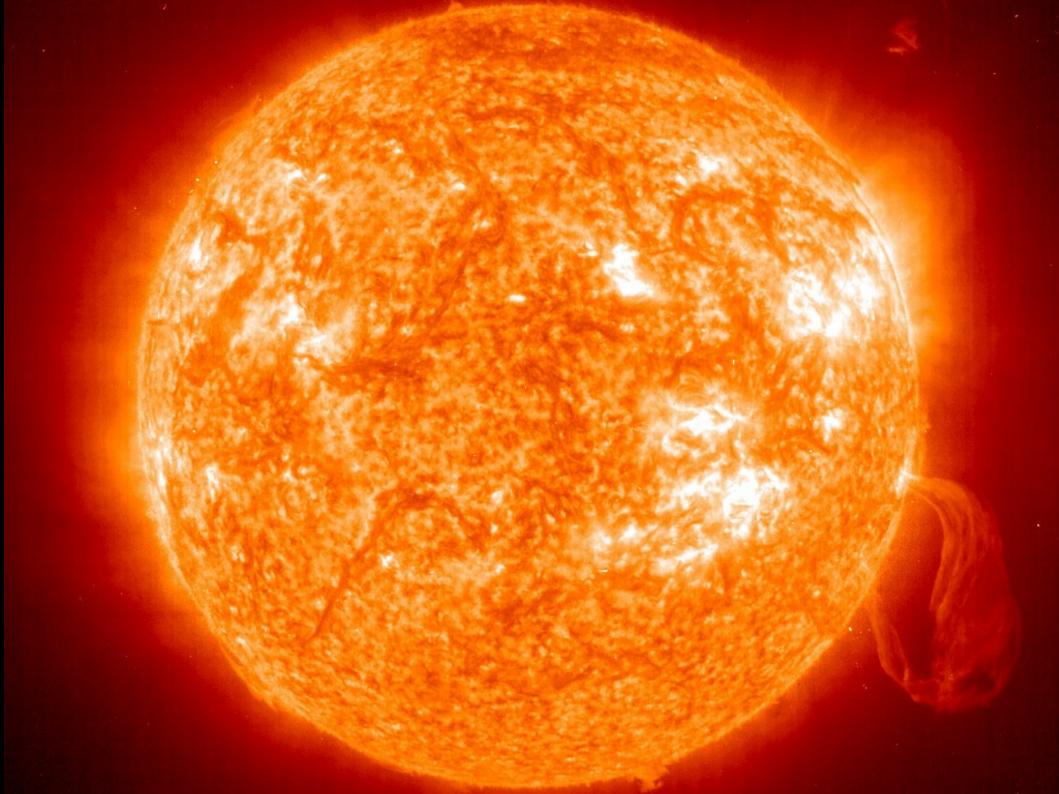


0.0 20.0 8480

 $\log_1 |\Sigma_0| < \Sigma$ 



...about four and a half billion years later...



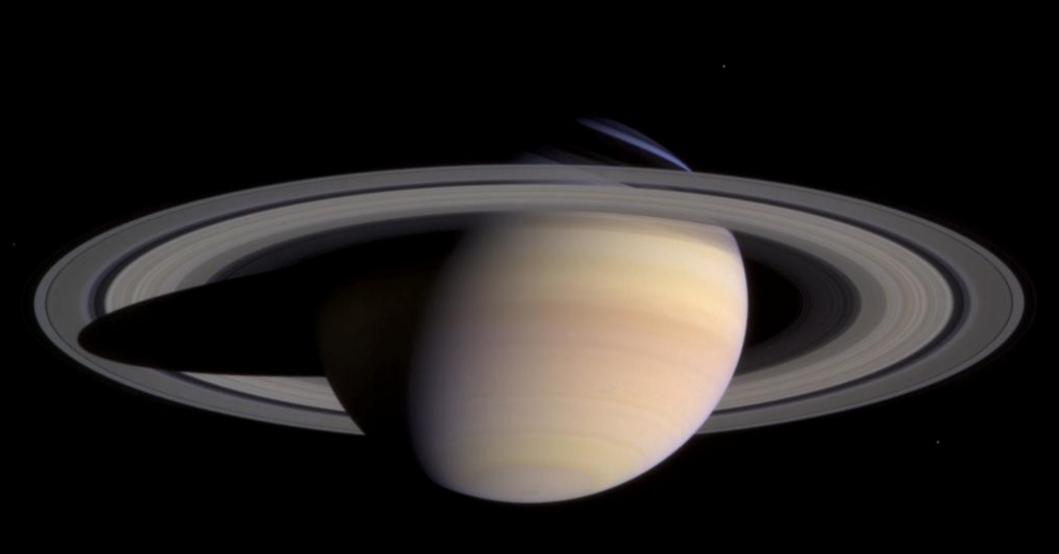


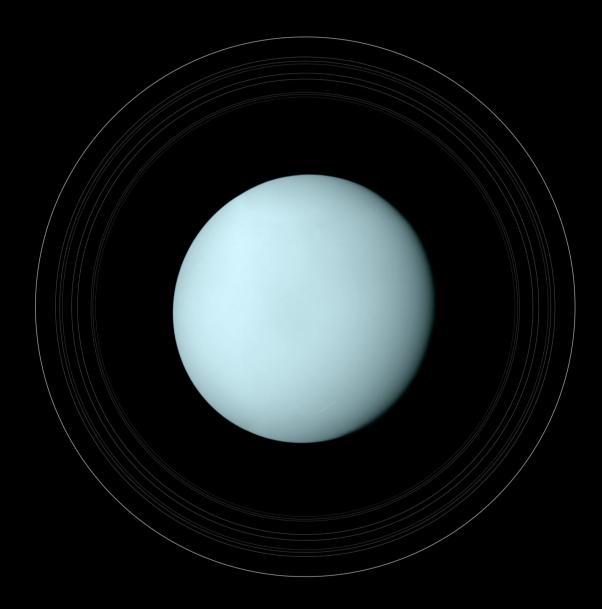


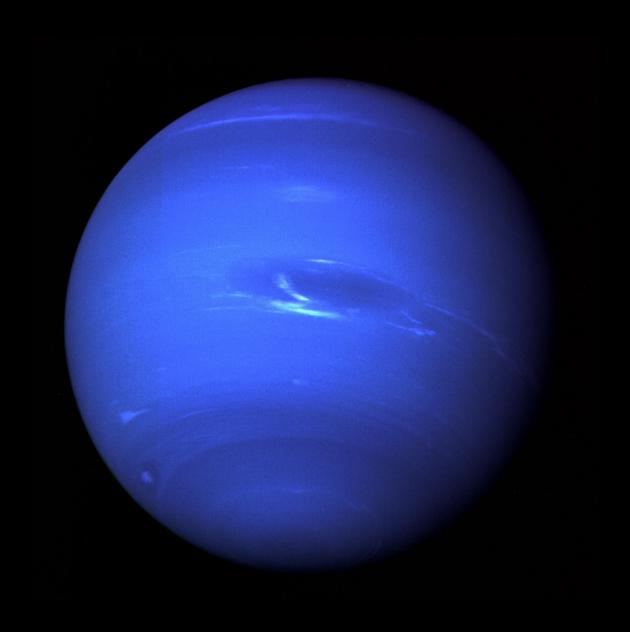




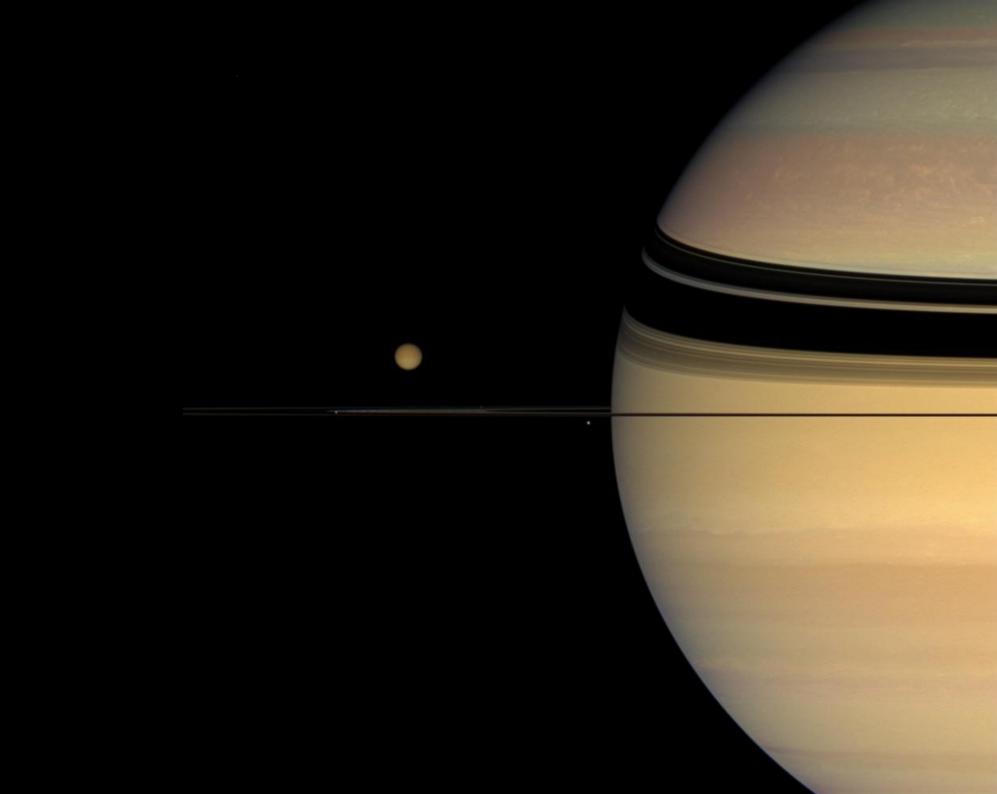








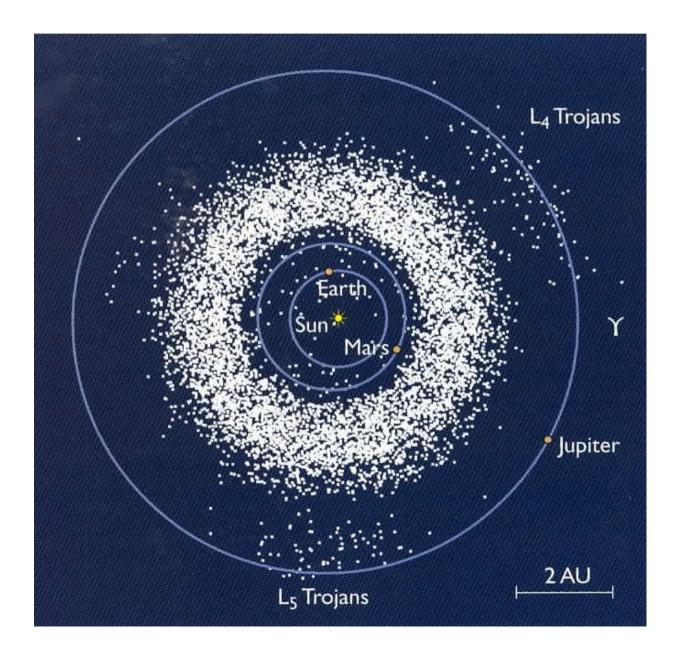




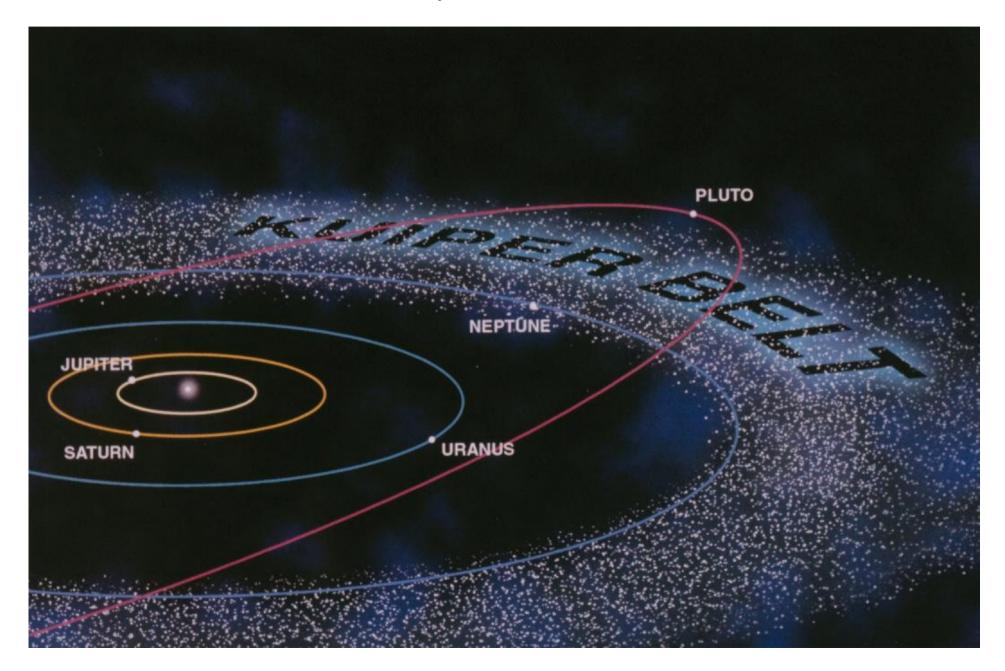


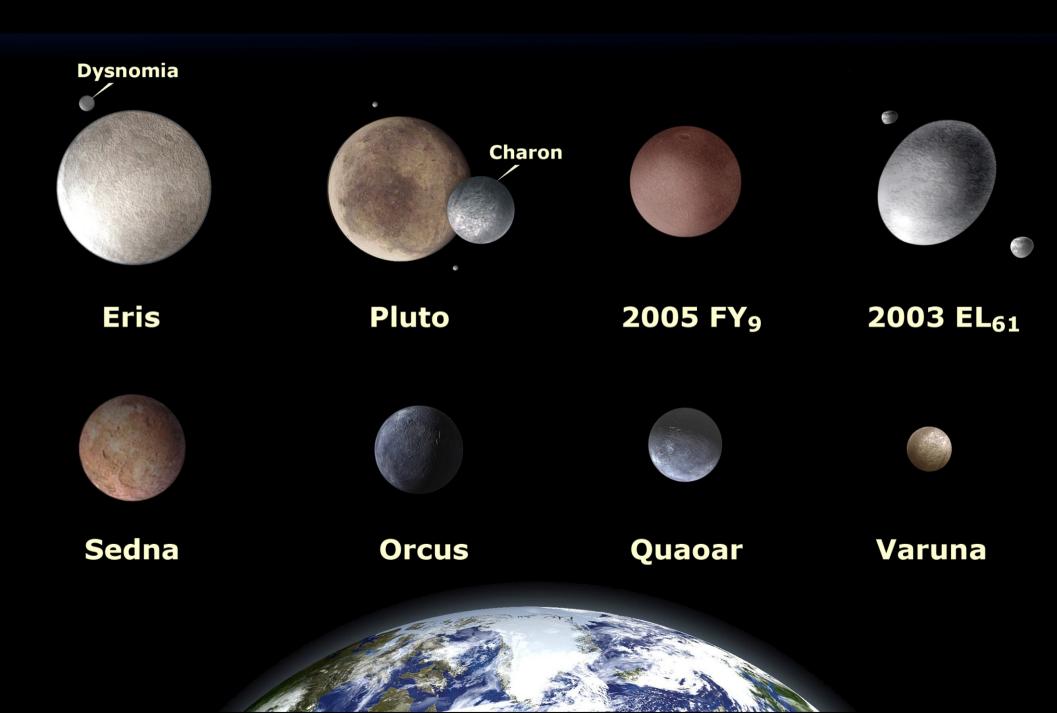


#### Asteroid Belt



### Kuiper Belt





## **Oort** Cloud 2 light-years **Oort Cloud Kuiper Belt** Pluto's Orbit

#### Tour of Time Scales of the Solar System

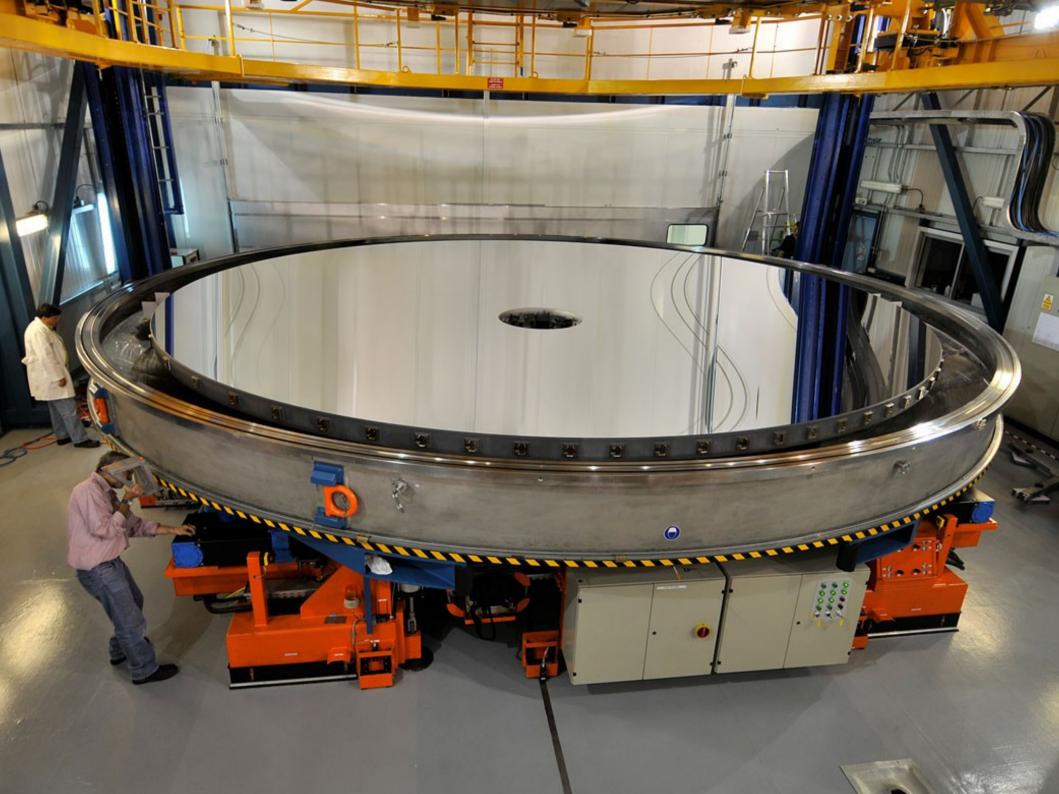
# How high could you jump on different objects in the Solar System?

#### How do we know this?

#### We have been searching...

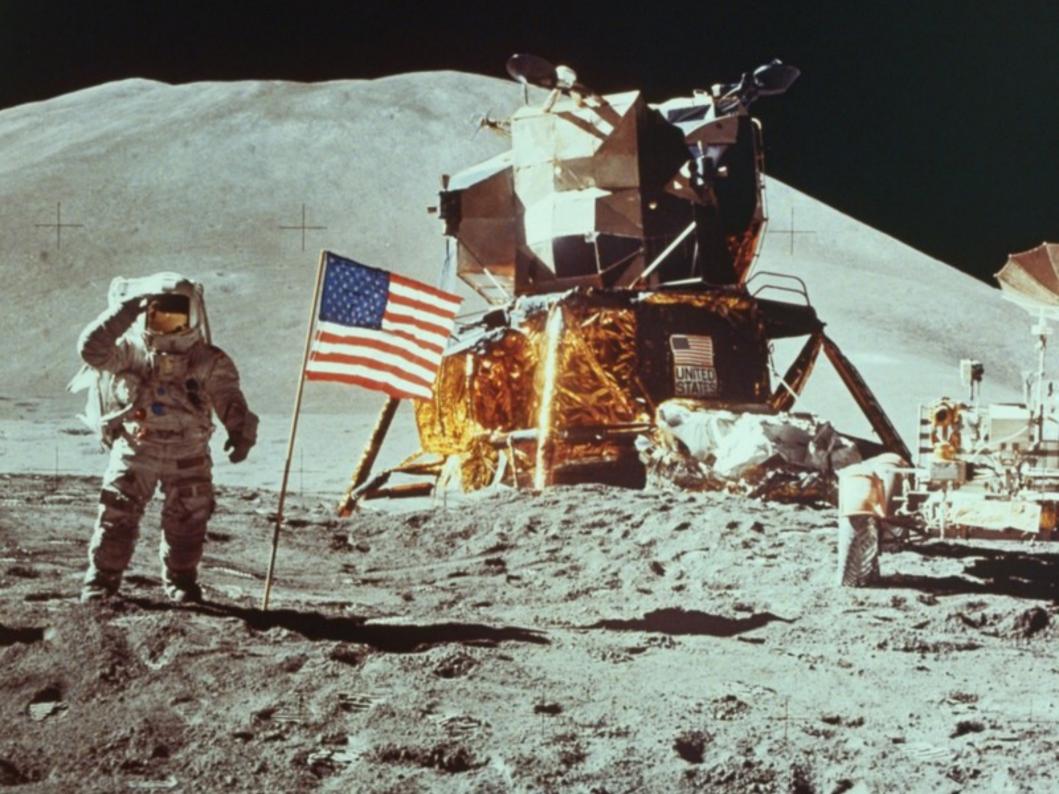


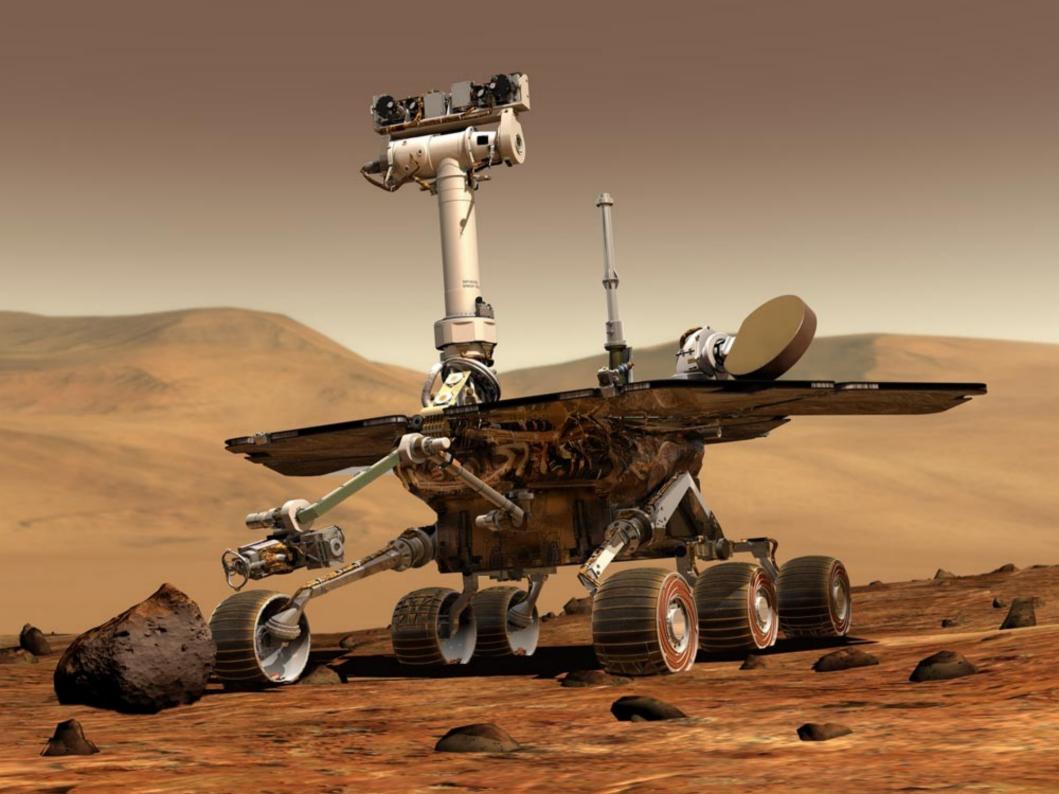






#### We've gone places...





#### We brought things back...





#### We've found clues...





#### PIECES IN THE PUZZLE OF THE 'SECRETS OF THE SOLAR SYSTEM'

