

## Minimum Mass Solar Nebula

class 3

The refractories composition of the planets is chondritic (show plot meteorites vs solar)  
Add volatiles until solar composition is reached.

MMSN: Jupiter's enrichment  $\rightarrow$  2-40 times solar far Widenessmilling  $\rightarrow$  zone's boundaries  $\rightarrow$  arithmetic mean of adjacent orbit

$$\Sigma(r) = 1.7 \times 10^3 \left( \frac{r}{\text{AU}} \right)^{-3/2} \text{ g cm}^{-2}$$

Integrating this to 30 AU

$$M = 2\pi \int_{r_0}^{r_1} \Sigma r dr = 2\pi \Sigma_0 \int_{r_0}^{r_1} \left( \frac{r}{\text{AU}} \right)^{-3/2} \cdot r dr = 2\pi \Sigma_0 \text{ AU}^{3/2} \int_{r_0}^{r_1} r^{-1/2} dr$$
$$= 2\pi \Sigma_0 \text{ AU}^{3/2} 2r^{1/2} \Big|_{r_0}^{r_1}$$

$$M = 4\pi \Sigma_0 \text{ AU}^{3/2} (\sqrt{r_1} - \sqrt{r_0}) = 4\pi \Sigma_0 \text{ AU}^2 \sqrt{30} \approx 2.6 \times 10^{31} \text{ g}$$

$$M_{\text{MMSN}} \approx 0.01 M_{\odot}$$

Minimum mass to form the planets.

Point out it assumes 100% efficiency, and ignores migration.

Estimate pressure:

$$\Sigma = 1500 \text{ g/cm}^2$$

$$T = 280 \text{ K}$$

$$P = \rho (f-1) c_v T$$

$$= \rho (f-1) c_v \cdot T$$

$$\rho = \frac{\Sigma}{H} = \frac{1500}{1 \times 10^{12}}$$

$$c_v \sim R \quad (R \sim 8. \times 10^7 \text{ erg K mol}^{-1})$$

$$P = \frac{1500}{10^{12}} \cdot 10^8 \cdot 200 = \frac{10^3 \cdot 10^8 \cdot 10^2}{10^{12}} = 10 \frac{\text{dyne}}{\text{cm}^2}$$

$$P = 10^{-5} \text{ bar}$$

$$\approx 1 \text{ Pa}$$

$$c_p - c_v = \frac{R}{M} = \frac{k_B}{m}$$

water sublimation temperature/pressure (show phase diagram)

Highlight the water delivery problem

Water ice under these pressures are at  $\sim 200 \text{ K}$ . At Earth, liquid water needs  $273 \text{ K} < T < 373 \text{ K}$ . 100 difference!

Planets in the habitable zone form from water-poor material!  
Water must come from somewhere.

$$\Sigma_{\text{rock}} \approx 7.1 \left( \frac{r}{\text{AU}} \right)^{-3/2} \text{ g cm}^{-2} \text{ for } r < 2.7 \text{ AU}$$

$$\Sigma_{\text{ice/rock}} \approx 30 \left( \frac{r}{\text{AU}} \right)^{-3/2} \text{ g cm}^{-2} \text{ for } r > 2.7 \text{ AU}$$

Evidence for water snowline at middle of asteroid belt:

Meteorites; Ceres; Jupiter Moons/Mars/Pressure

Deuterium to hydrogen in water from Earth: 156 ppm

Carbonaceous chondrites: 154  $\pm$  10 ppm

Port Cloud comets: 300 ppm

Water on Earth:  $3 \times 10^{-4} M_{\oplus}$

Mass of asteroid belt:  $3 \times 10^{-4} M_{\oplus}$

10% of carbonaceous chondrites is water, so only

$3 \times 10^{-5} M_{\oplus}$  of water at present.

Asteroid belt should be AT LEAST 10x more massive before, if all flux arrived at Earth. Another factor 10 at least, so originally belt should have been a few  $\times 10^{-3} M_{\oplus}$ , 100x more massive than today.