The Shakura-Sunyaev model

Turbulence is a source of VISOSIty p7 u, = pF, - 2, (ps, + pu, u,)

For mean quantites

Decompor in mean and fluctuation

ρ ly(x,+u,)=ph-)(ph,+ph,+ + (x,+u,')(x,+u,'))

Averep out (11)=0

p 2 = u, = pf - 2, (p &1, + p v, u, + p <u, n, >)=0

This term behaves like a viscous tensor, and transports angular momentum

Effect vely, it's a diffusion of momentum, coming from the colvection term, advecting momentum at of a gaussian surface.

Rey wolds straw behave like a VISOUS ferm

$$\frac{\partial L}{\partial t} + \frac{1}{R} \frac{\partial}{\partial r} (R \log r) = \frac{1}{R} \frac{\partial}{\partial r} (v L \frac{d \ln \Lambda}{d \ln R})$$

$$\frac{\partial}{\partial t} (\Sigma r^2 \Lambda) + \frac{1}{2} (r(\Sigma r^2 \Lambda) ur) = \frac{1}{2} \int \nabla T r d\phi$$

$$T = 2\pi r \quad v \Sigma r d\Lambda \quad r \quad r \times Fv$$
where $r \times Fv$
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$$\frac{\partial}{\partial r} (\Sigma r^2 \Lambda) + \frac{\partial}{\partial r} (r(\Sigma r^2 \Lambda) ur) = \frac{1}{2} \int \Lambda \left(\frac{\partial v}{\partial r} + \frac{\partial v}{\partial r}\right) dr$$

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$$\frac{\partial v}{\partial r} = \frac{1}{2} (r \Lambda) = \Lambda + r d\Lambda = \frac{1}{2} \int \Lambda r d\Lambda$$
Shor
$$\frac{\partial v}{\partial r} = \frac{1}{2} \int \Lambda r d\Lambda$$

Suss continuity equation on at momentum of and x+ == =0 31 + 1 3 (Rlur) = 1 3 (VL dlur)

3+ + 1 3 (+ 1 + 1) = 1 2 (1 b (1 b (1 b) r2p (n/ up)= vl dln/l

$$v = -\left(\frac{d \ln \Lambda}{d \ln S}\right) \frac{R^{+r}}{p \Lambda}, R^{+r} = \rho \langle u_r^{\dagger} u_{\phi} \rangle$$

Equivalence turbulence and visasity This momentum difficion's in fact a pressure so, we are write

S in fect a pressure so, we are write

$$R^{\phi r} = \times P$$

The stress is proportional to the pressure The consolity is correct here The stress generates a pressure V = - (alnr) xp = 2x Rcs?

Two different ways to write it Both physical The visco-

15 otropic length The size of the addres is H The speads are will then med, &

If a is constant, it's possible to solve andytically for several dish phenomena

Time scales
$$T = \frac{1}{2} = \left(\frac{H}{R}\right)^{-2} \frac{1}{4R}$$

For $\left(\frac{H}{R}\right) \approx 0.05$ $4 \approx 10^{-2}$

Turbulance

Raylagh on term what leads to turblence -

3+ (0 D) 0 = - 1 Db + d

Decompose the flow into base (hime-indepent) and penturbation $u(\vec{x},t) = \bar{u}(\vec{x}) + \dot{u}(\vec{x},t)$, with $\bar{u}_r = 0$ and $\hat{u}_{\varphi} = L \Gamma >> u_{\varphi}$

 $\frac{\partial r}{\partial r} + n \theta \frac{\partial}{\partial r} n - \frac{n \theta}{n \theta} = V_{SL} \xrightarrow{n = n + n} \frac{\partial r}{\partial r} - N_{SL} - 5 \sqrt{n \theta} = -N_{SL}$

(20r - 2 R rs = 0)

3t + nr 3 nt + none =) 3nt + nr 3 (nr) + 8xx

The perturbation equations are thus $N = N_o \left(\frac{r_o}{r}\right)^{\frac{q}{r}}$

9 + Mr, - 5 V Nb= 2 and 3+ Mp + V(5-4) nr = 0 Now consider the perturbation as a found mode 4'= \$ e-7 (wt-Kx)

 $-1 \omega \hat{u}_r - 2 \Omega \hat{v}_{\phi} = 0 \qquad \Rightarrow \hat{u}_{\phi} = -1 \omega \hat{v}_r$ $-1 \omega \hat{v}_{\phi} + \Omega(2-\alpha) \hat{u}_r = 0 \qquad \Rightarrow 2 \Omega$

$$-i\omega\left(\frac{1}{2}\omega\right) \hat{v}_{r}^{2} + \Omega(2-q)\hat{v}_{r}^{2} = 0$$

$$-\omega^{2} \hat{v}_{r}^{2} + \Omega(2-q)\hat{v}_{r}^{2} = 0$$

$$\omega^{2} - 2(2-q) \quad \Omega^{2} = 0$$

$$\omega^{2} = 2\Omega^{2}(2-q) \quad \text{if } q(2) \text{ then the flow is stable}$$

$$This fegural is called epicyclic frequency, $x = \Omega\sqrt{2(2-q)}$

$$q = -\frac{MnR}{d\ln r} < 2 \text{ is the case for keplenen Disks So, they are stable}$$

$$\left| \frac{d\ln \Omega}{d\ln r} \right| < 2 \quad \text{Stability}$$

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$$\frac{d\ln \Omega}{d\ln r} + 2\Omega r$$

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=
$$\Re \left(2 + \frac{d \ln \Lambda}{d \ell} \right) = \Re \left(2 - q \right)$$

