

```
In [67]: %matplotlib inline

import numpy as np
import matplotlib.pyplot as plt
from scipy.integrate import quad
from scipy.interpolate import interp1d
import pdb
from tools import plots
```

First, create a function to compute an arbitrary Sersic profile.

```
In [68]: def profile(r, I_e, R_e=1., n=1., mag=False) :
    """
    routine to return Sersic profile

    Args:
        r: array-like
            input array of radii at which to calculate profile
        I_e: float
            surface brightness at effective radius, in intensity or magnitude
            units, depending on mag
        R_e: float, optional
            effective radius, default=1.
        n: sersic index (default=1, i.e. exponential)
        mag: bool, optional
            return in mag units if True, default=False
    """

    bn=1.99*n-0.327
    I_r = np.exp (-bn * ((r/R_e)**(1./n) - 1) )           # Sersic function
    if mag :
        return I_e - 2.5 *np.log10(I_r)
    else :
        return I_e * I_r
```

Write a function to calculate flux in an annulus :  $2 \pi r SB(r)$

```
In [69]: def flux(r, I_e=0., R_e=1., n=1) :
    """ Helper routine to calculate annular flux at given radius for integrator
    """
    return 2*np.pi*r*profile(r,I_e,R_e,n=n)
```

Function to calculate and plot profiles at a range of different Sersic indices. Optionally calculate and plot cumulative profile.



Function to show multiple profiles with different  $R_e$  and Sersic  $n$ , normalized at some radius, used to construct Straatman Figure 1  
plot

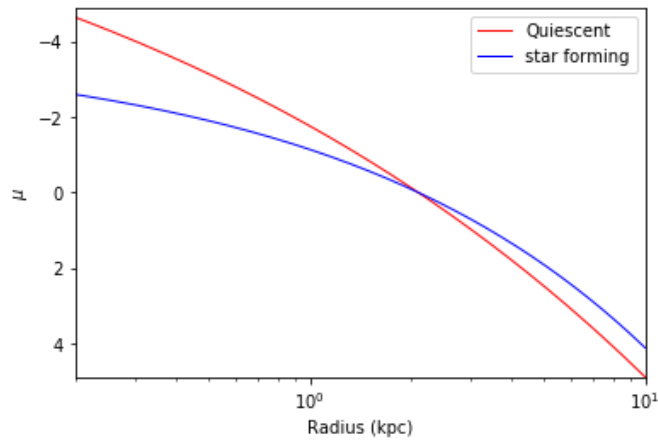
```
In [72]: def straatman(R_es=[0.85,2.62],ns=[4.14,2.17],labels=['Quiescent','star forming'],colors=['red','blue'])
    """ Plot multiple Sersic profiles, normalized at some radius
    """
    r=np.arange(0.,10.,0.01)
    fig,ax=plt.subplots(1,1)
    for R_e,n,label,color in zip(R_es,ns,labels,colors) :
        prof = profile(r,I_e=0.,R_e=R_e,n=n,mag=True)

        # construct an interpolating function
        interp=interp1d(r,prof)

        # interpolate at normalizing radius and plot
        norm = interp(rnorm)
        plots.plot1(ax,r,prof-norm,label=label,yr=[5,-5],xr=[0.1,10.1],
                    xt='Radius (kpc)',yt='$\mu$',color=color)

    ax.set_xscale('log')
    ax.legend()

straatman()
```



Routine to calculate Petrosian radius for a Sersic profile

```

In [75]: def petro(R_e=1,n=1,eta=0.2, plot = False) :

    """ Calculate Petrosian radius for Sersic profile with index n

    I_e : float, optional
           SB at effective radius, default=1.
    n : float, optional
        Sersic index, default=1.
    eta : float, optional
          Critical value of Petrosian ratio to determine Petrosian radius
    """

    # input radii to 10 R_e
    r=np.arange(0.,10.*R_e,0.01)
    ratio=[]
    I_e=1 # note Petrosian ratio independent of I_e!
    for rr in r :
        # integrate flux internal to rr
        tot=quad(flux,0.,rr,args=(I_e,R_e,n))[0]

        # Petrosian ratio = SB / average_internal_flux
        ratio.append(profile(rr,I_e,R_e,n=n)/(tot/(np.pi*rr**2)))

    # interpolate to desired ratio, eta
    interp=interp1d(ratio,r)
    petrorad = interp(eta)

    if plot :
        fig,ax=plots.multi(1,1)
        plots.plot1(ax,r,ratio,xt='Radius',yt='Petrosian ratio')
        ax.plot([petrorad,petrorad],plt.ylim())
        ax.grid()
        ax.set_title('R_e: {:.2f}   n: {:.2f}   R_petro : {:.2f}'.format(R_e,n,petrorad))

    return petrorad

```

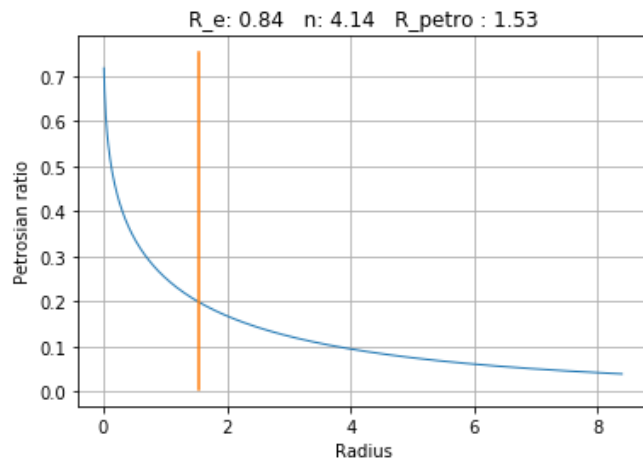
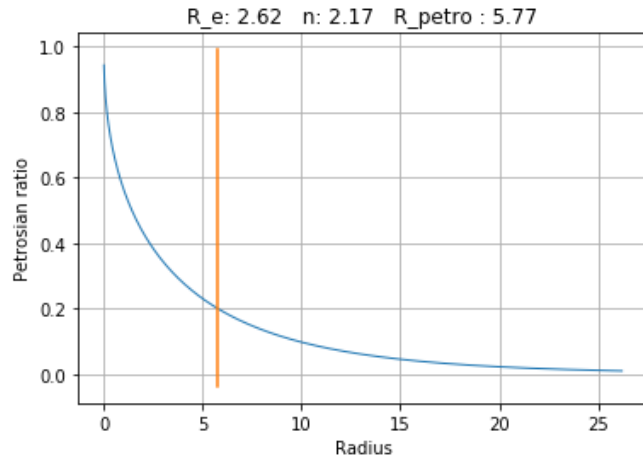
Calculate Petrosian radius for Straatman et al quiescent and star forming stack model parameters

```
In [76]: print('star forming: R_p = {:.2f}'.format(petro(R_e=2.62,n=2.17,plot=True)))  
print('quiescent: R_p = {:.2f}'.format(petro(R_e=0.84,n=4.14,plot=True)))
```

/Users/holtz/anaconda3/lib/python3.7/site-packages/ipykernel\_launcher.py:22: RuntimeWarning: invalid value encountered in double\_scalars

star forming: R\_p = 5.77

quiescent: R\_p = 1.53



In [ ]: