

HR DIAGRAM AND STAR SIZES!



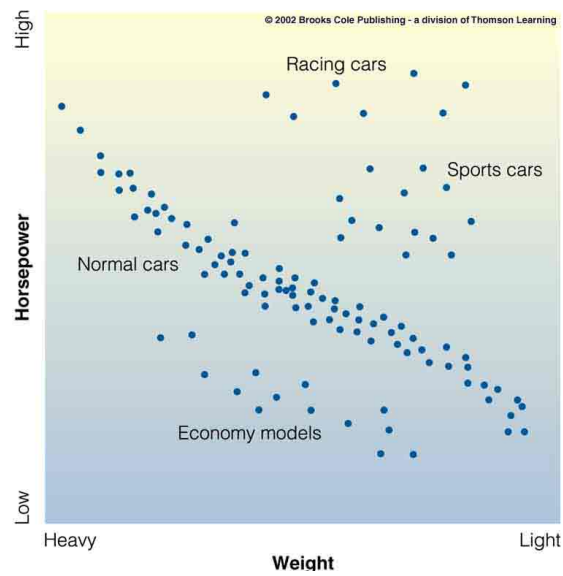
So, we know how to measure the
following properties of stars...

1. **Luminosity** (from Brightness and Distance)
2. **Mass** (from Doppler shifts in Binary Stars)
3. **Temperature** (from stellar spectrum- Blackbody curve)

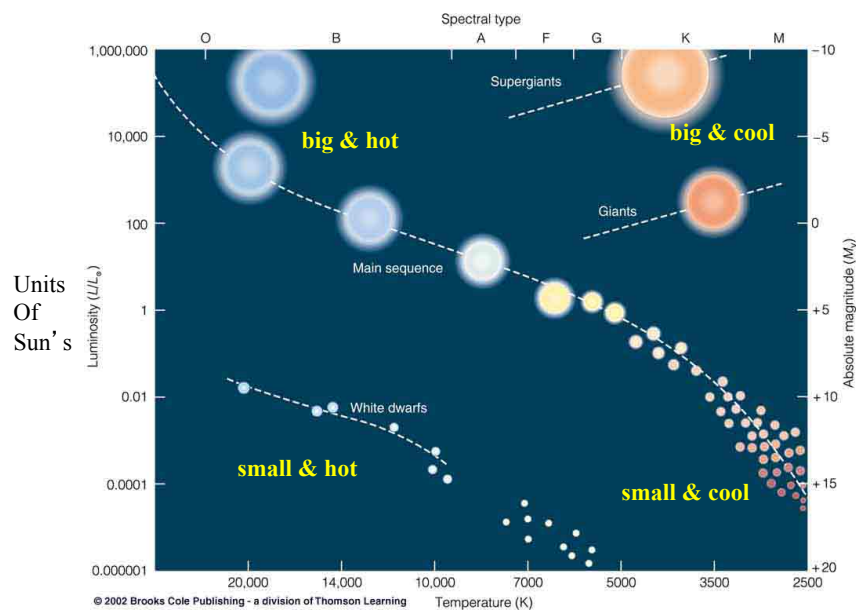
*NOW, we will learn how to measure the sizes of
stars from their luminosity and temperature!*

But first, the HR Diagram

HR Diagram for Automobiles



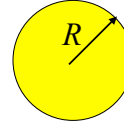
HR Diagram for Stars



For a given star, the **Luminosity depends upon two properties**:

1. The total surface area that emits the light
2. The light emitted per unit surface area (flux)

Surface area is *radius dependent* $\text{Area} = \text{const} \times R^2$



Flux is the light energy radiated from every unit area of the star's surface, it is *temperature dependent* $\text{Flux} = \text{const} \times T^4$

$$\text{Luminosity} = k \times \text{Area} \times \text{Flux}$$

We have

$$L = k R^2 T^4$$

k is a constant

$$L = k R^2 T^4$$

The above relation is shown as lines of constant radius as a function of L and T on the HR diagram...

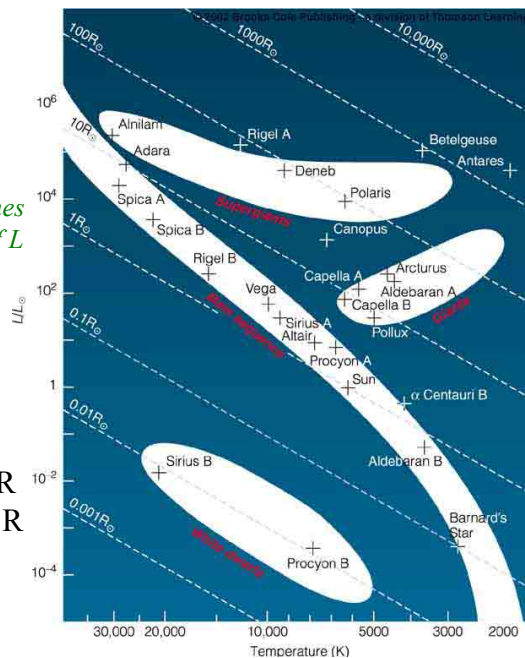
Main sequence:

- O Stars 10x Sun's R
- M stars 1/10x Sun's R

Giants 1000-10,000x Sun's R

White Dwarfs 1/100x Sun's R

WDs about size of Earth!



$$L = k R^2 T^4$$

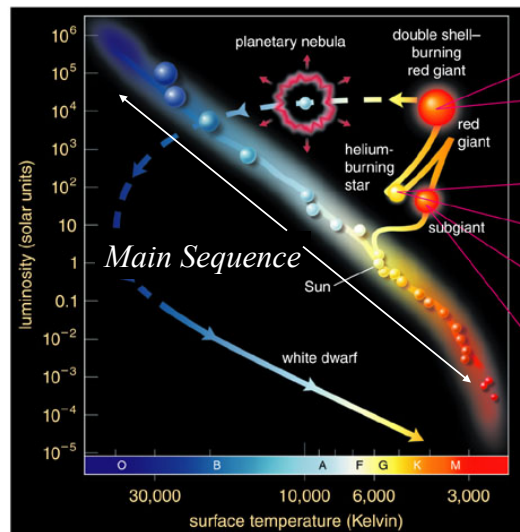
Compared to the Sun....

2 x R gives 4x L
 3 x R gives 9x L
 5 x R gives 25x L
 $\frac{1}{2}$ x R gives $\frac{1}{4}$ x L

2 x R and 2 x T gives 64 x L
 3 x R and $\frac{1}{2}$ x T gives $\frac{9}{16}$ x L = 0.56L
 $\frac{1}{2}$ x R and 3 x T gives $8\frac{1}{4}$ x L = 20.25L

2 x T gives 16x L
 3 x T gives 81x L
 $\frac{1}{2}$ x T gives $\frac{1}{16}$ x L

And, so it goes...



2nd stage red giant

1st stage red giant

Let's Focus:

**THE
MAIN
SEQUENCE**

Copyright © 2004 Pearson Education, publishing as Addison Wesley.

Main Sequence Is

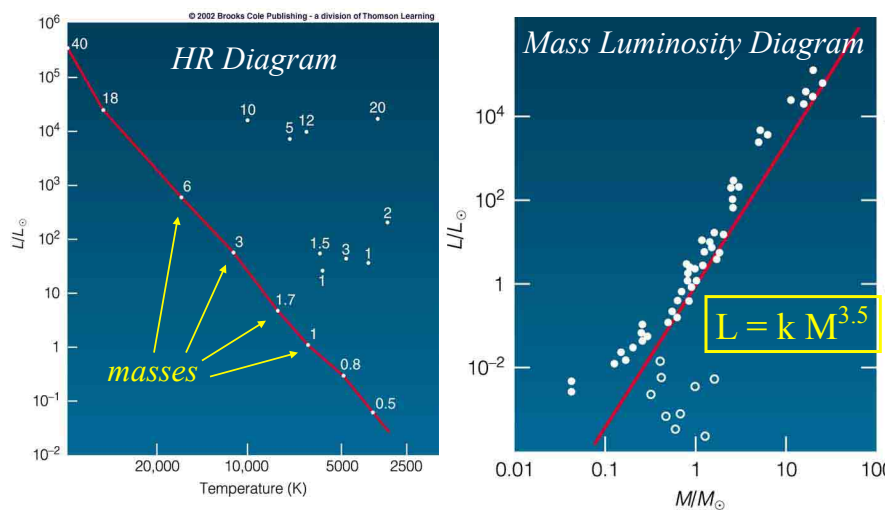
O, B, A, F, G, K, M

increasing temperature $\leftarrow$ $\rightarrow$ decreasing temperature
 increasing size $\leftarrow$ $\rightarrow$ decreasing size
 increasing luminosity $\leftarrow$ $\rightarrow$ decreasing luminosity

The above applies to the Main Sequence Only
 Giants and White Dwarfs do NOT obey all these trends.

Main Sequence is also a Mass Sequence

As temperature increases, mass increases! As mass increases, luminosity increases!



$$L = k M^{3.5}$$

Compared to the Sun....

2 x M gives	$2 \times 2 \times 2 \times \sqrt{2} =$	11.3 L
3 x M gives	$3 \times 3 \times 3 \times \sqrt{3} =$	46.7 L
5 x M gives	$5 \times 5 \times 5 \times \sqrt{5} =$	279 L
10 x M gives	...	= 3162 L
50 x M gives	...	= 883,900 L

In other words, massive stars have very very very very large luminosities

Main Sequence
Is

O, B, A, F, G, K, M

increasing temperature $\leftarrow$ $\longrightarrow$ decreasing temperature
 increasing size $\leftarrow$ $\longrightarrow$ decreasing size
 increasing mass $\leftarrow$ $\longrightarrow$ decreasing mass
 increasing luminosity $\leftarrow$ $\longrightarrow$ decreasing luminosity

This applies to the Main Sequence Only
 Giants and White Dwarfs do NOT obey all these trends.